



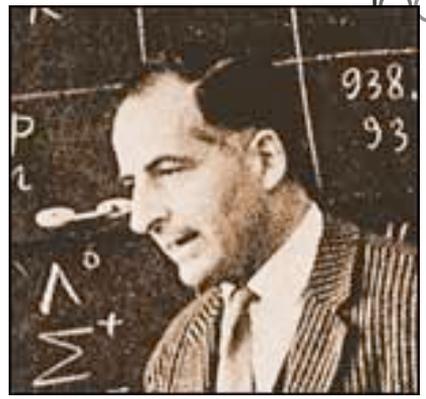
# First $\nu_{\mu} \rightarrow \nu_e$ Oscillation Results from MiniBooNE

*Morgan Wascko  
Imperial College London*

*UCL HEP Seminar  
May 11, 2007*

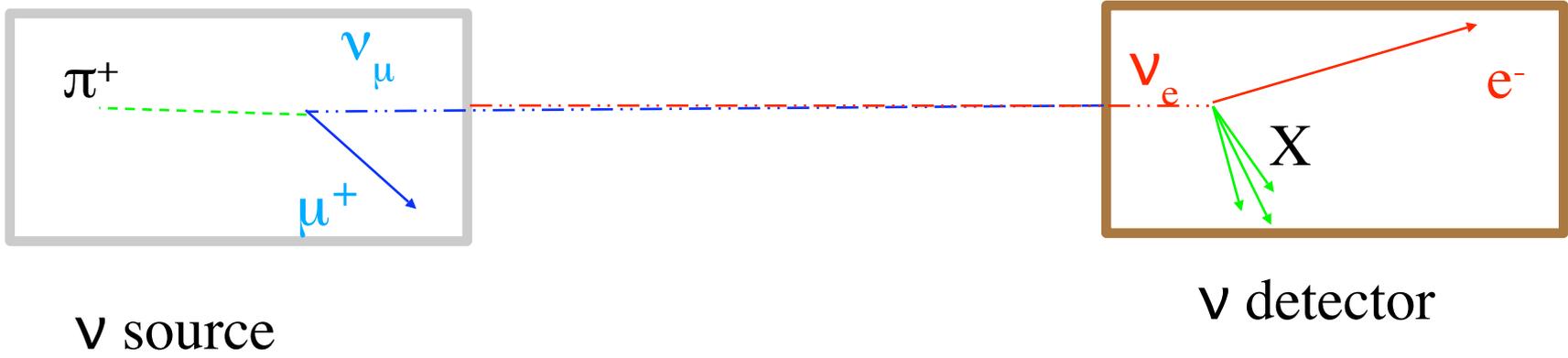
1. Motivation & Introduction
2. Description of the Experiment
3. Analysis Overview
4. Two Independent Oscillation Searches
5. First Results

# Motivation: Neutrino Oscillations



*Pontecorvo, 1957*

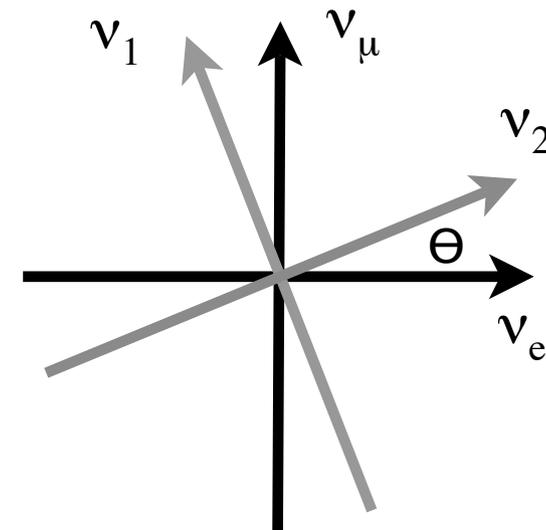
if neutrinos have mass, a neutrino that is produced as a  $\nu_\mu$  (e.g.  $\pi^+ \rightarrow \mu^+ \nu_\mu$ ) has a non-zero probability to oscillate and some time later be detected as a  $\nu_e$  (e.g.  $\nu_e n \rightarrow e^- p$ )



# Motivation: Neutrino Oscillations

In a world with 2 neutrinos,  
if the weak eigenstates ( $\nu_e, \nu_\mu$ )  
are different from the mass eigenstates ( $\nu_1, \nu_2$ ):

$$\begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$



The weak states are mixtures of the mass states:

$$|\nu_\mu\rangle = -\sin\theta|\nu_1\rangle + \cos\theta|\nu_2\rangle$$

$$|\nu_\mu(t)\rangle = -\sin\theta (|\nu_1\rangle e^{-iE_1 t}) + \cos\theta (|\nu_2\rangle e^{-iE_2 t})$$

The probability to find a  $\nu_e$  when you started with a  $\nu_\mu$  is:

$$P_{oscillation}(\nu_\mu \rightarrow \nu_e) = |\langle \nu_e | \nu_\mu(t) \rangle|^2$$

# Motivation: Neutrino Oscillations

In units that experimentalists like:

$$P_{oscillation}(\nu_{\mu} \rightarrow \nu_e) = \sin^2 2\theta \sin^2 \left( \frac{1.27 \Delta m^2 (eV^2) L (km)}{E_{\nu} (GeV)} \right)$$

*Oscillation probability between 2 flavour states depends on:*

## 1. fundamental parameters

$\Delta m^2 = m_1^2 - m_2^2$  = mass squared difference between states

$\sin^2 2\theta$  = mixing between  $\nu$  flavours

## 2. experimental parameters

L = distance from  $\nu$  source to detector

E =  $\nu$  energy

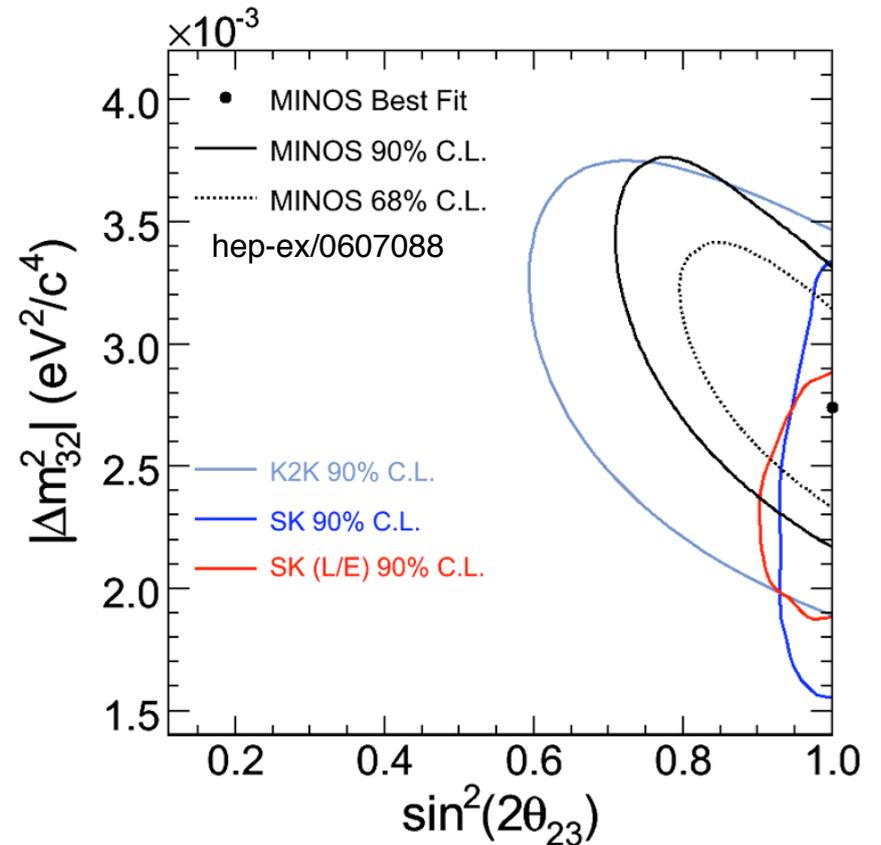
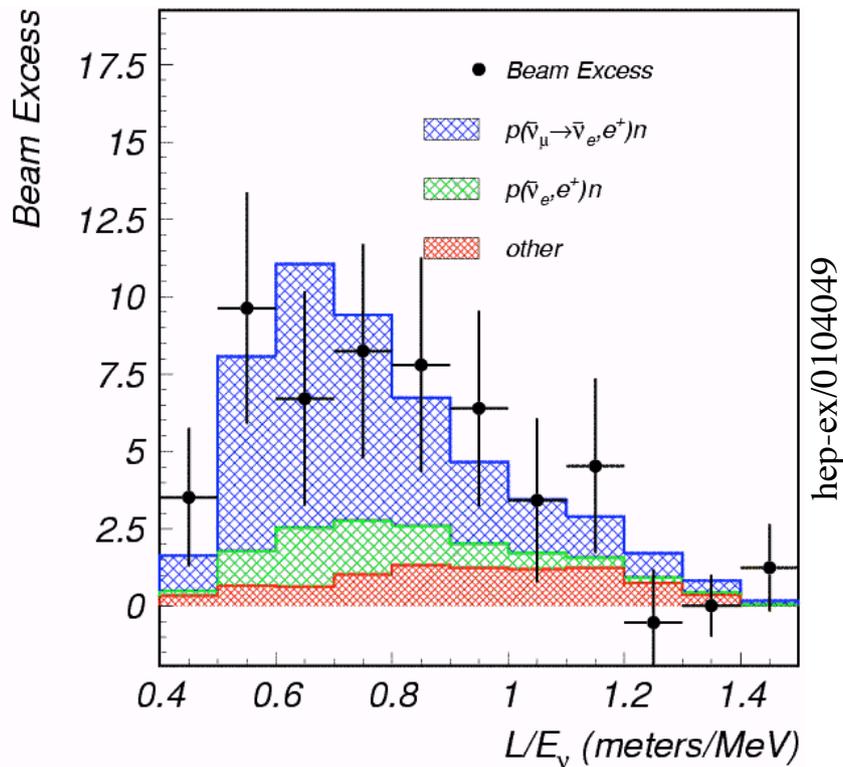
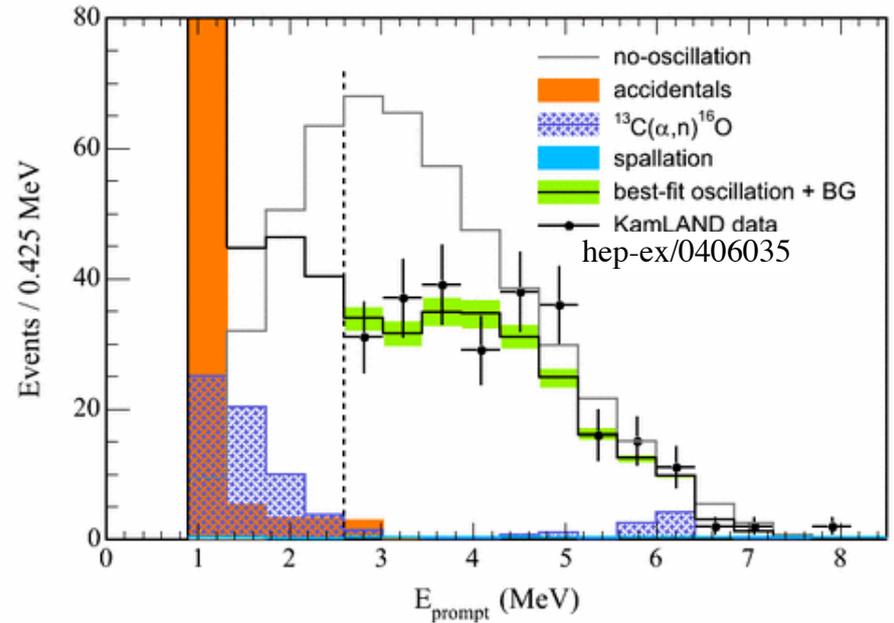


# Motivation: Oscillation Signals

*Solar*  $\nu$ : measured by Homestake, ..., SNO  
confirmed by KamLAND

*Atmospheric*  $\nu$ : measured by K-II, ..., Super-K  
confirmed by Soudan2, MACRO, K2K, MINOS

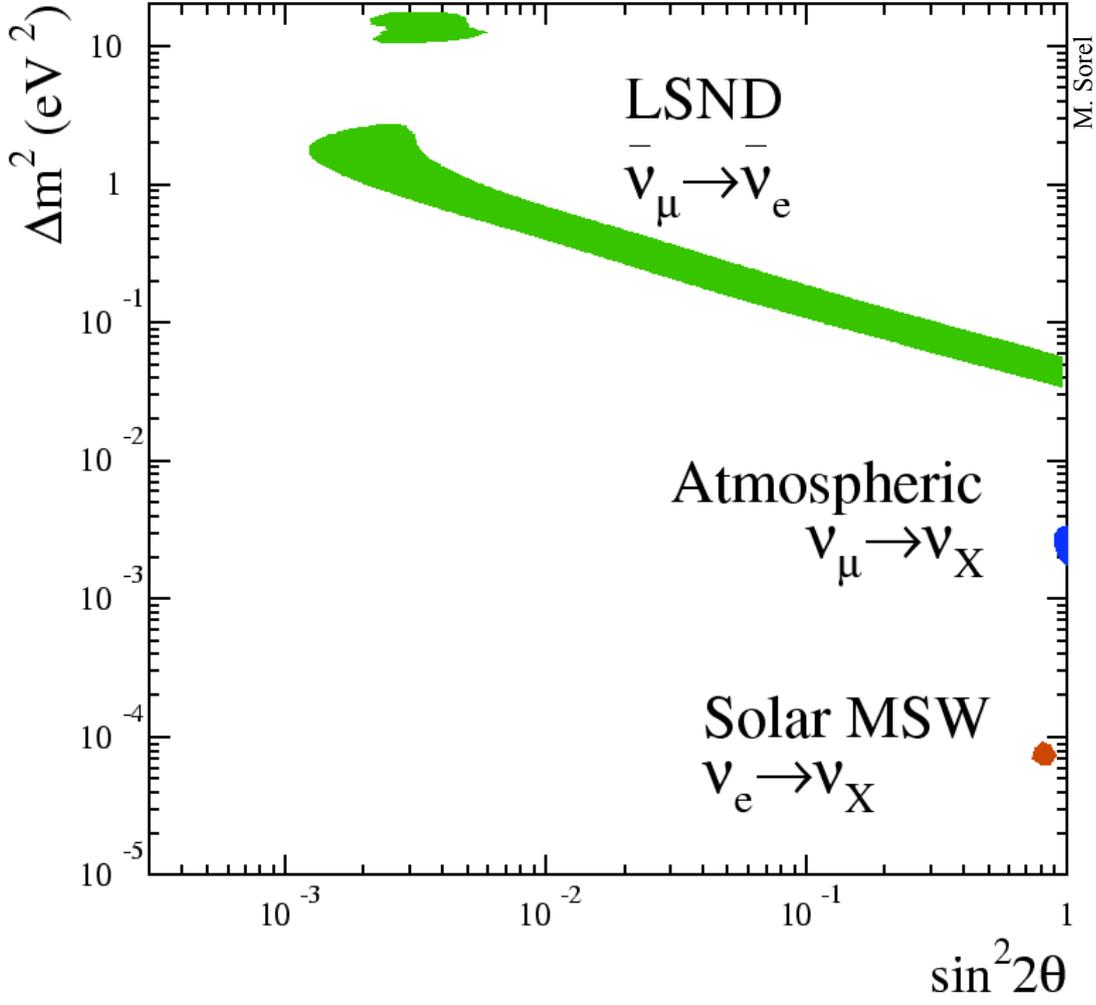
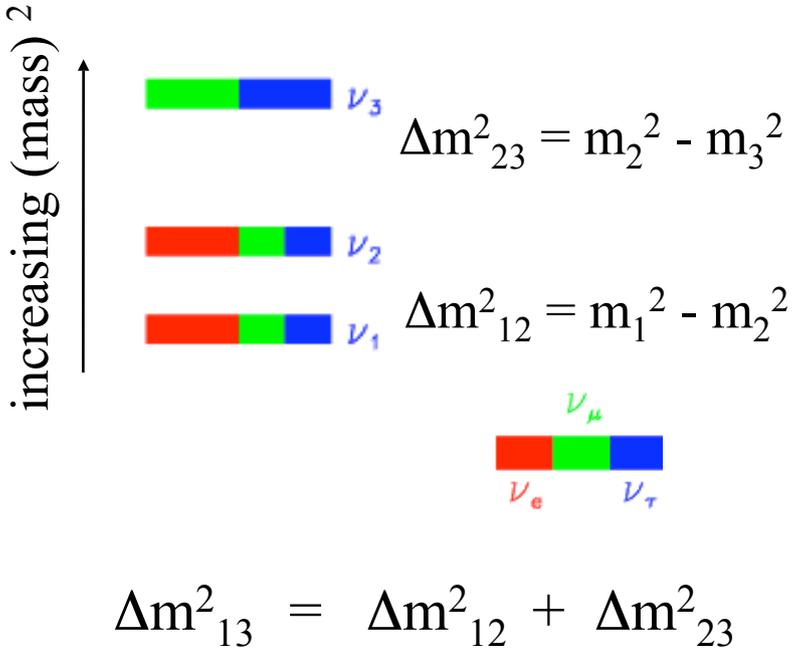
*Accelerator*  $\nu$ : measured by LSND  
**unconfirmed**



# Motivation: The Problem

$$P_{oscillation}(\nu_\mu \rightarrow \nu_e) = \sin^2 2\theta \sin^2 \left( \frac{1.27 \Delta m^2 (eV^2) L (km)}{E_\nu (GeV)} \right)$$

A standard 3 neutrino picture:

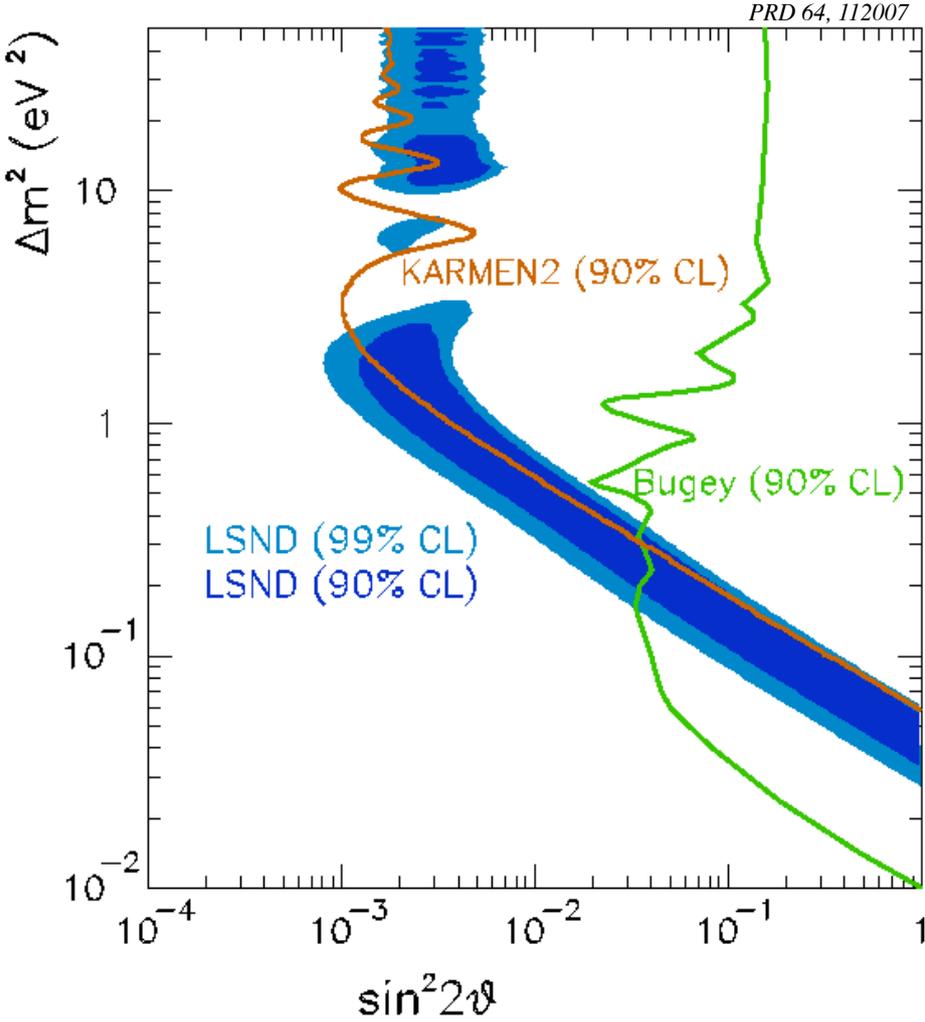


*The oscillation signals cannot be reconciled without introducing physics (even farther) beyond the Standard Model.*

# Motivation: LSND

*MiniBooNE was proposed in 1997 to address the LSND result.*

LSND observed a  $4\sigma$  excess of  $\bar{\nu}_e$  events in a  $\bar{\nu}_\mu$  beam:  $87.9 \pm 22.4 \pm 6.0$   
 interpreted as 2-neutrino oscillations,  $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) = 0.26\%$



$$P = \sin^2 2\theta \sin^2 \left( \frac{1.27 \Delta m^2 (eV^2) L (km)}{E_\nu (GeV)} \right)$$

### MiniBooNE strategy:

Keep  $(L/E_\nu)$  same as LSND but change systematics, including event signature:

- Order of magnitude higher  $E_\nu$  than LSND
- Order of magnitude longer baseline  $L$  than LSND
- Search for excess of  $\nu_e$  events above background

*Simple  $\nu_\mu \rightarrow \nu_e$  oscillation*

# The MiniBooNE Collaboration

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Fermilab Visual Media Services

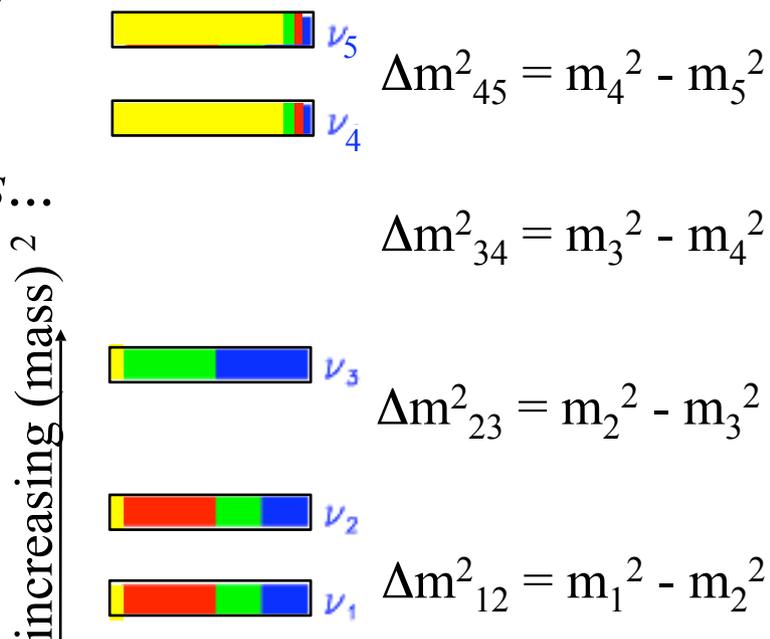
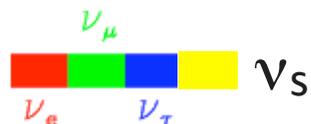
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# Motivation: MiniBooNE and LSND

*If MiniBooNE observes LSND-type  $\nu$  oscillations...*

The simplest explanation is to add more  $\nu$ s,  
to allow more independent  $\Delta m^2$  values.

The new  $\nu$ s would have to be **sterile**, otherwise  
they would have been seen already.



*If MiniBooNE does not observe LSND-type oscillations...*

The Standard Model wins again!

Today: MiniBooNE's initial results on testing the LSND anomaly

- A generic search for a  $\nu_e$  excess in our  $\nu_\mu$  beam,
- An analysis of the data within a  $\nu_\mu \rightarrow \nu_e$  appearance-only context

# Outline

1. Motivation & Introduction
2. Description of the Experiment
  - Beam
  - Detector
3. Analysis Overview
4. Two Independent Oscillation Searches
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# MiniBooNE Overview: Beam and Detector

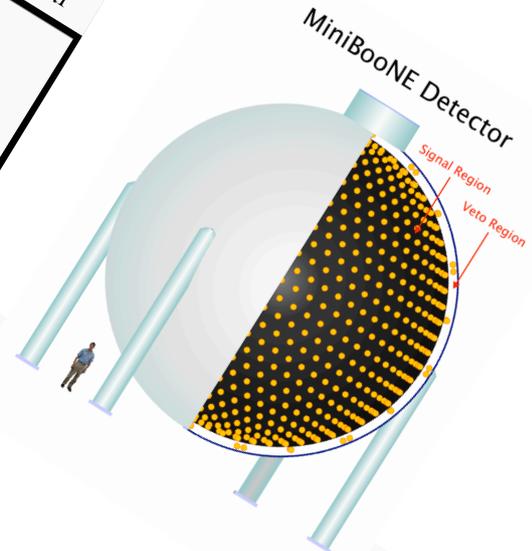
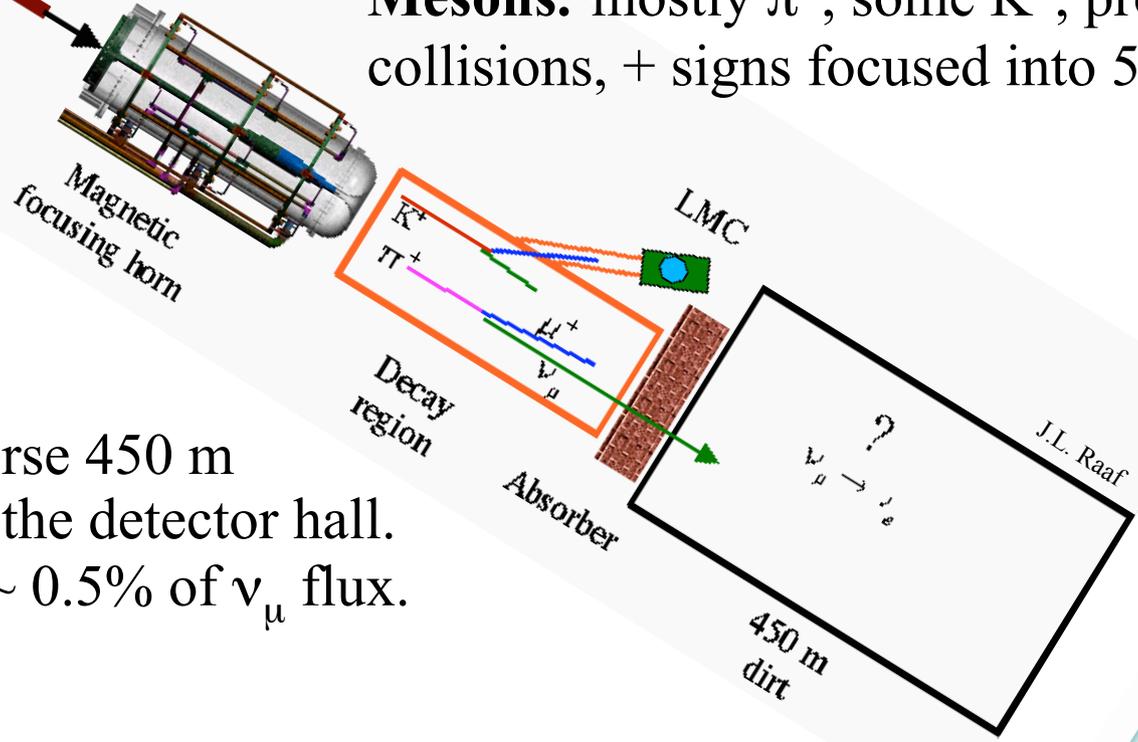
**Protons:**  $4 \times 10^{12}$  protons per  $1.6 \mu\text{s}$  pulse, at 3 - 4 Hz from Fermilab Booster accelerator, with  $E_{\text{proton}} = 8.9 \text{ GeV}$ .  
*First result uses  $(5.58 \pm 0.12) \times 10^{20}$  protons on target.*



**Mesons:** mostly  $\pi^+$ , some  $K^+$ , produced in p-Be collisions, + signs focused into 50 m decay region.

**Neutrinos:** traverse 450 m soil berm before the detector hall.  
 Intrinsic  $\nu_e$  flux  $\sim 0.5\%$  of  $\nu_\mu$  flux.

**Detector:** 6 m radius, 250,000 gallons of mineral oil ( $\text{CH}_2$ ), which emits Cherenkov and scintillation light.  
 1280 inner PMTs, 240 PMTs in outer veto region

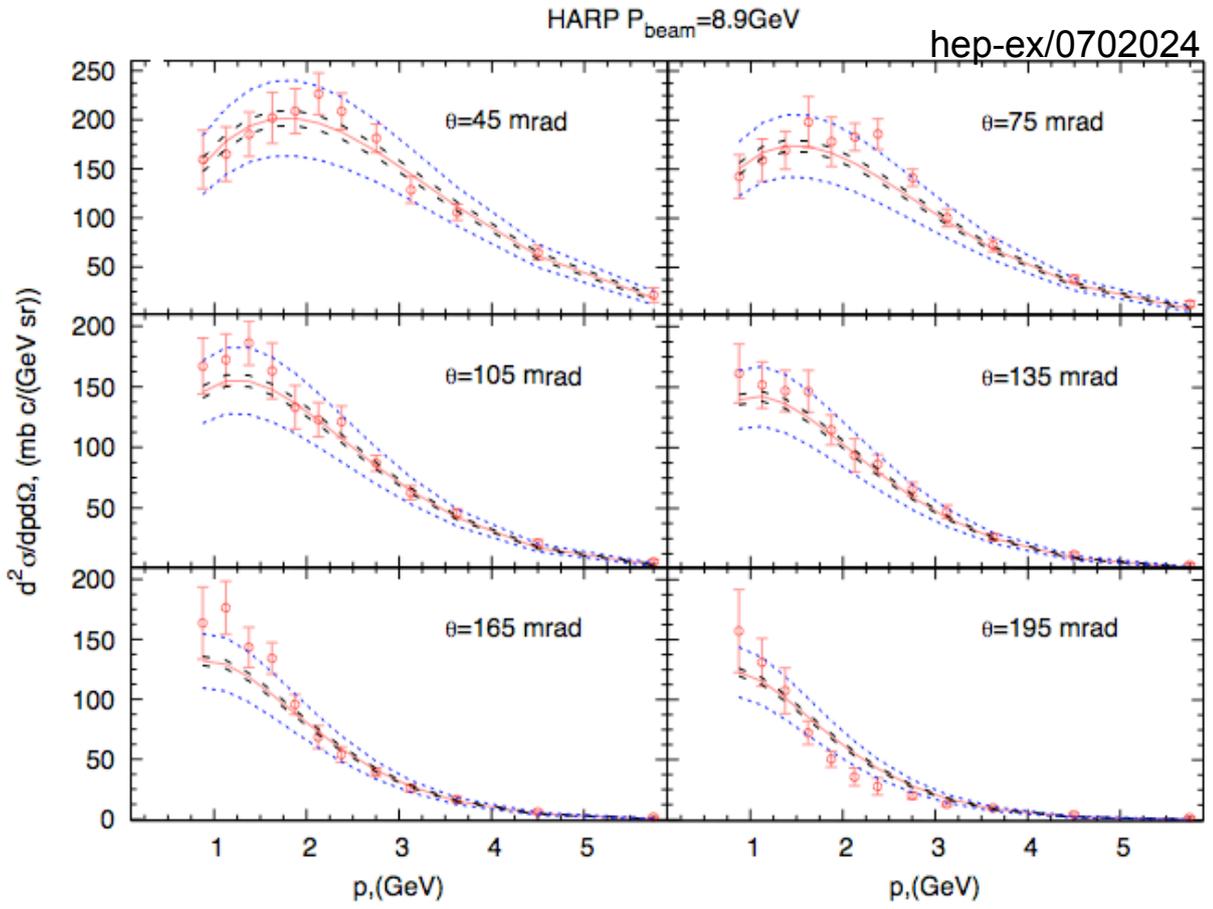
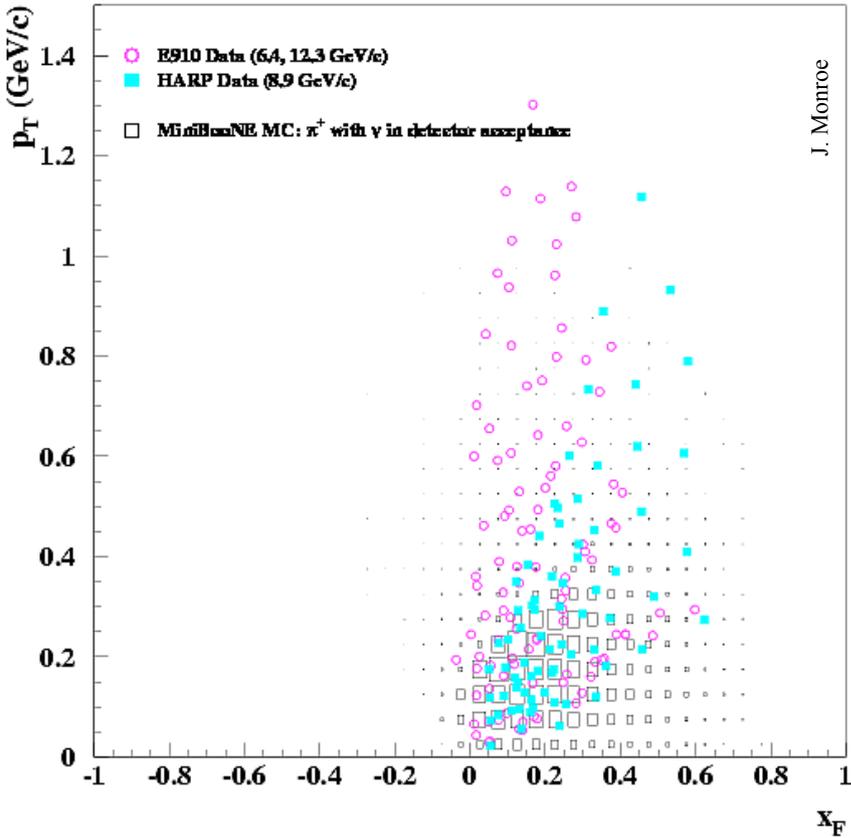


# Booster Neutrino Beam: Modelling Meson Production

*Prediction from a fit to  $p Be \rightarrow \pi^+ X$  production data from E910 and HARP experiments ( $p_p = 6-12 \text{ GeV}/c$ ,  $\Theta_\pi = 0 - 330 \text{ mrad}$ .)*

**Fit** (shown at right) uses Sanford-Wang parametrisation

HARP has excellent phase space coverage for MiniBooNE



$\pi^-$  similarly parametrised  
 Kaons flux predictions use a Feynman Scaling  
 parametrisation (no HARP data yet)

# Booster Neutrino Beam: Neutrino Flux

*MiniBooNE is searching for an excess of  $\nu_e$  in a  $\nu_\mu$  beam*

Modelled with a Geant4 Monte Carlo

“Intrinsic”  $\nu_e + \bar{\nu}_e$  content: 0.5%

$\nu_e$  Sources:

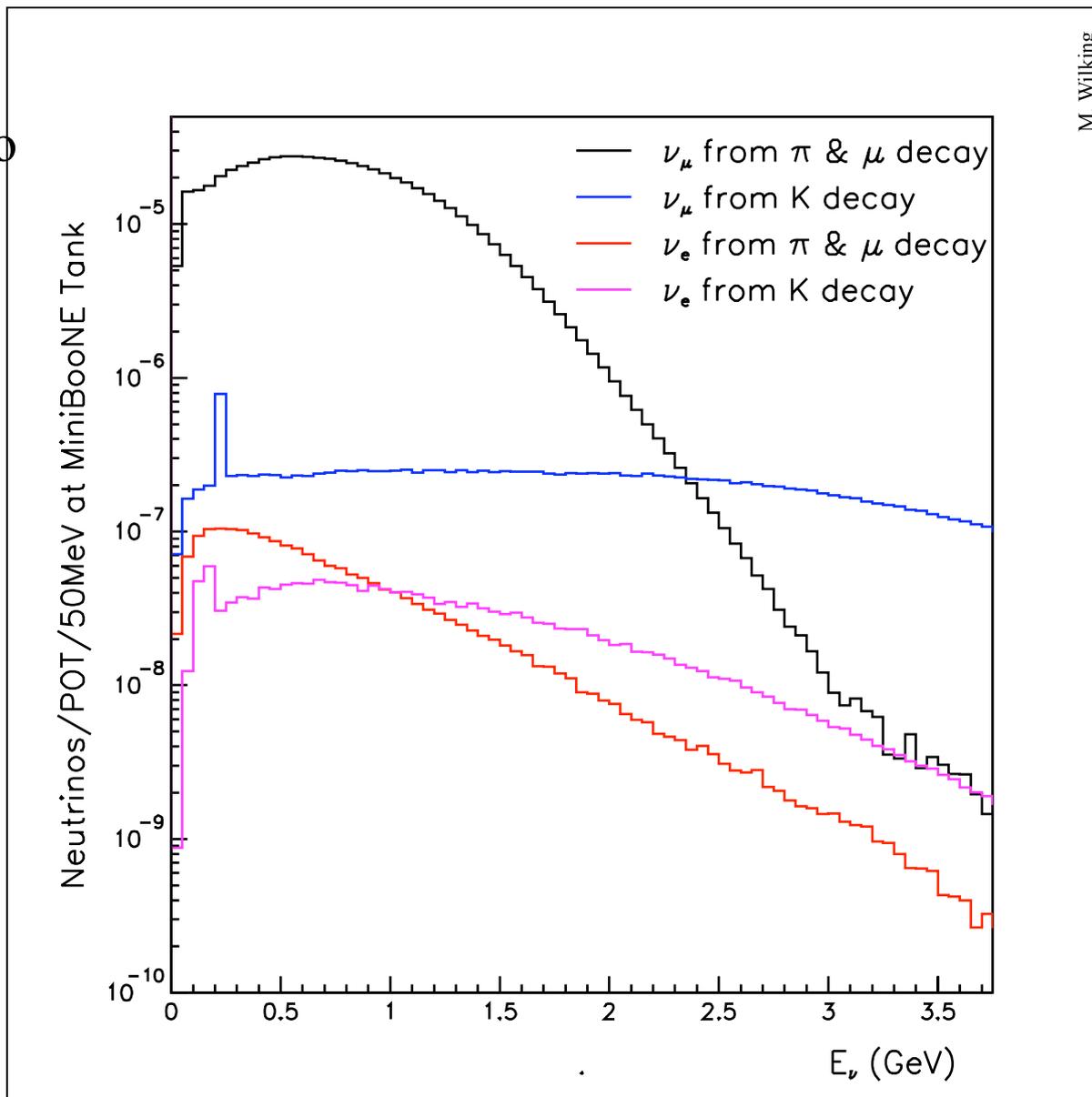
$$\mu^+ \rightarrow e^+ \bar{\nu}_\mu \nu_e \quad (42\%)$$

$$K^+ \rightarrow \pi^0 e^+ \nu_e \quad (28\%)$$

$$K^0 \rightarrow \pi^+ e^- \nu_e \quad (16\%)$$

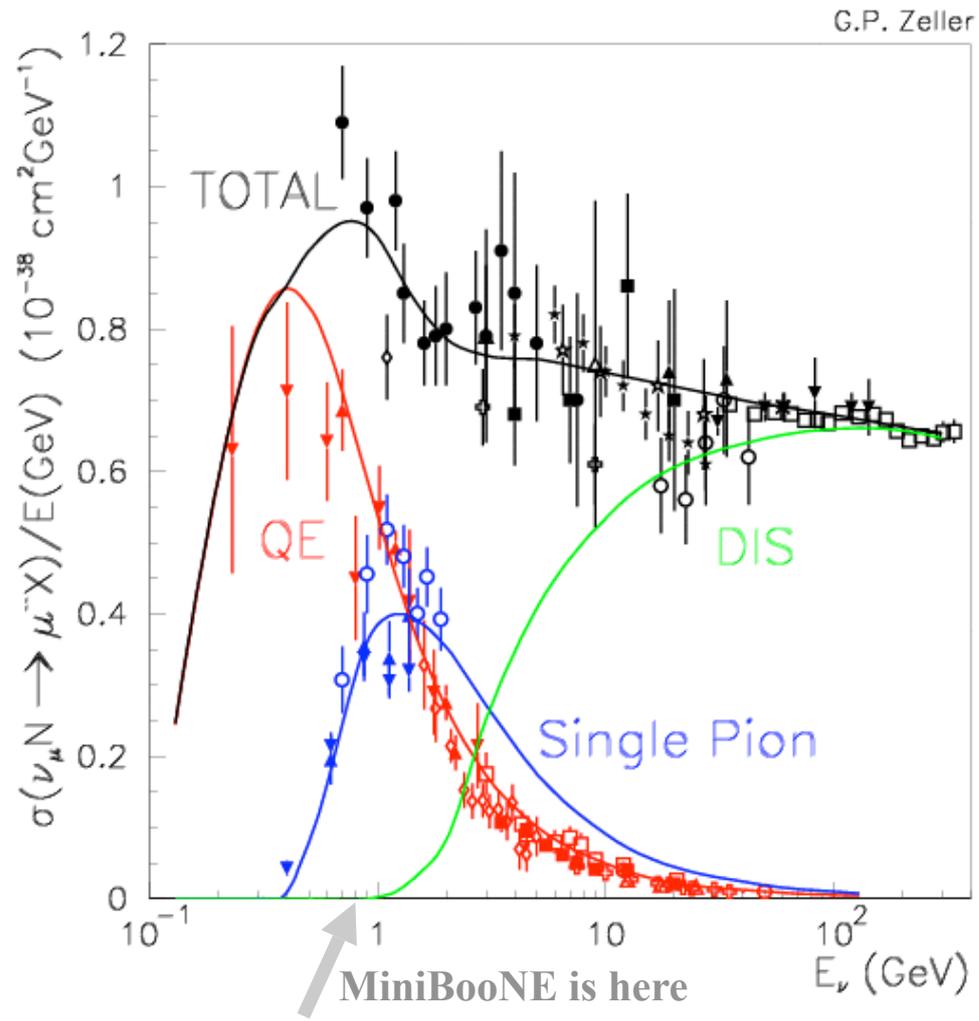
$$\pi^+ \rightarrow e^+ \nu_e \quad (4\%)$$

Antineutrino content: 6%

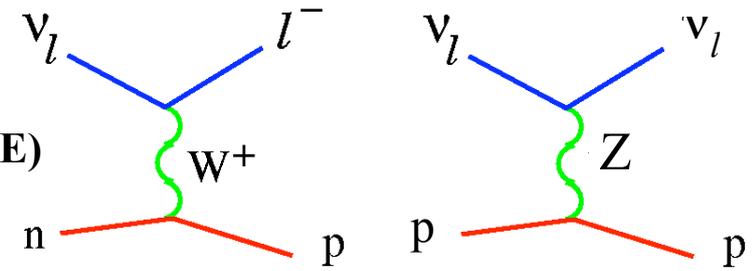


# MiniBooNE Detector: Neutrino Cross Sections

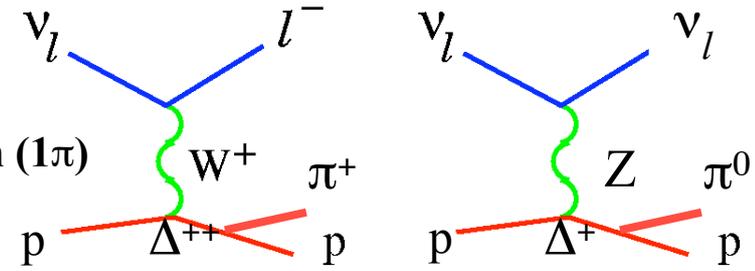
*Modelling what the neutrinos do in the detector*



**CC / NC  
quasi-elastic  
scattering (QE)**  
42% / 16%



**CC / NC  
resonance  
production (1π)**  
25% / 7%



Cross section predictions from  
NUANCE Monte Carlo

*Use CCQE events for oscillation analysis signal channel:*

$$E_{\nu}^{QE} = \frac{1}{2} \frac{2M_p E_{\ell} - m_{\ell}^2}{M_p - E_{\ell} + \sqrt{(E_{\ell}^2 - m_{\ell}^2) \cos \theta_{\ell}}}$$

Only need lepton direction  
and angle to find  $\nu$  energy!

# MiniBooNE Detector: Optics

*charged final state particles produce  $\Upsilon$ s*

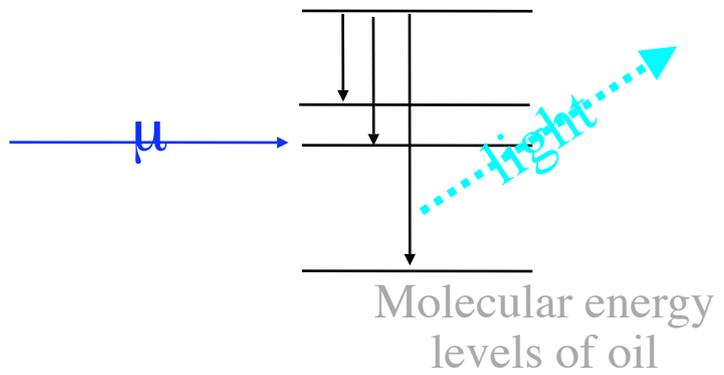
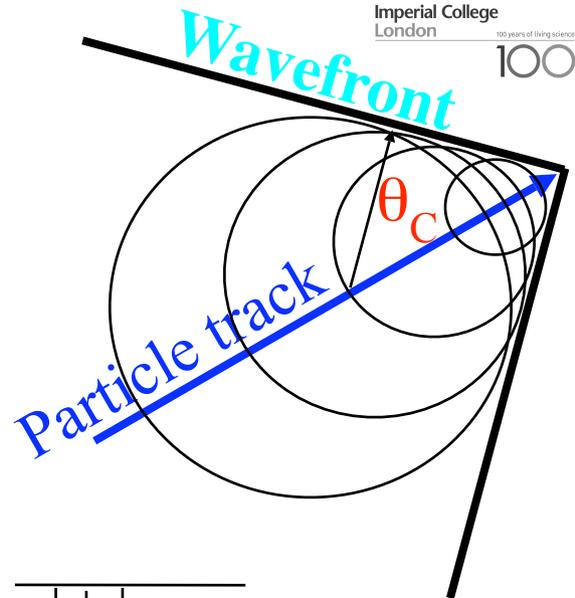
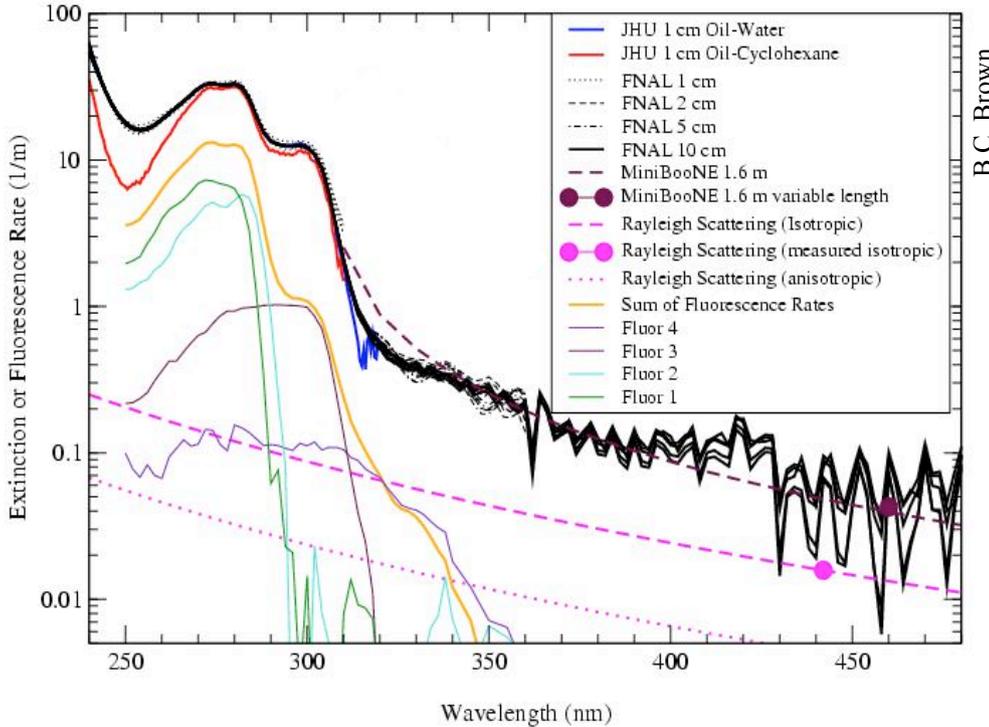
## Cherenkov radiation

- Light emitted by oil if particle  $v > c/n$
- forward and prompt in time

## Scintillation

- Excited molecules emit de-excitation  $\Upsilon$ s
- isotropic and late in time

Extinction Rate for MiniBooNE Marcol 7 Mineral Oil



*$\Upsilon$ s detected by PMTs after undergoing absorption reemission, scattering, fluorescence*

*“the optical model”*

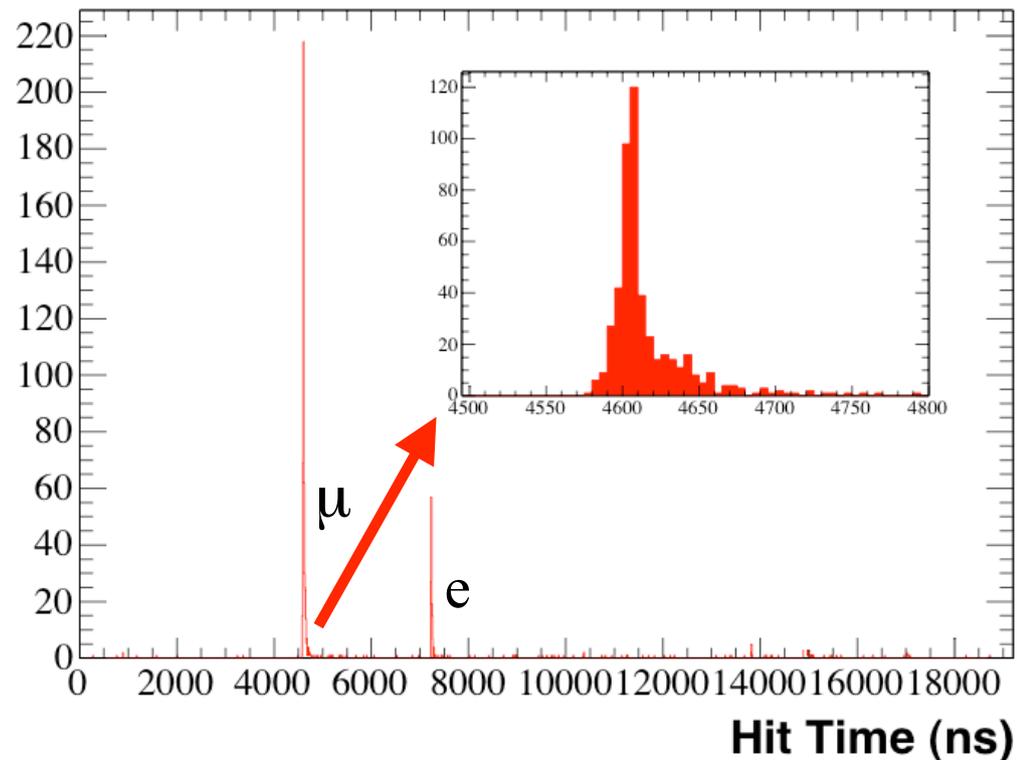
*→ PMT Hits*

# MiniBooNE Detector: Hits

*First set of cuts based on simple hit clusters in time: “sub-events.”*

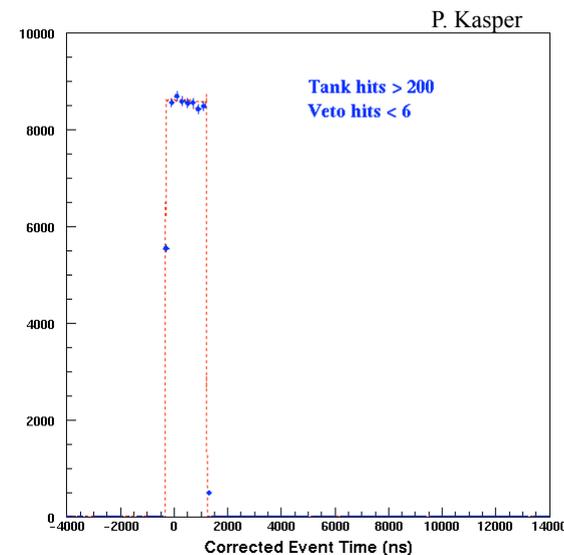
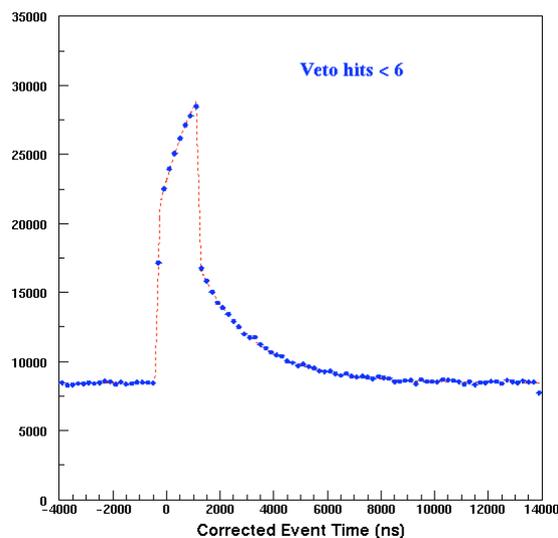
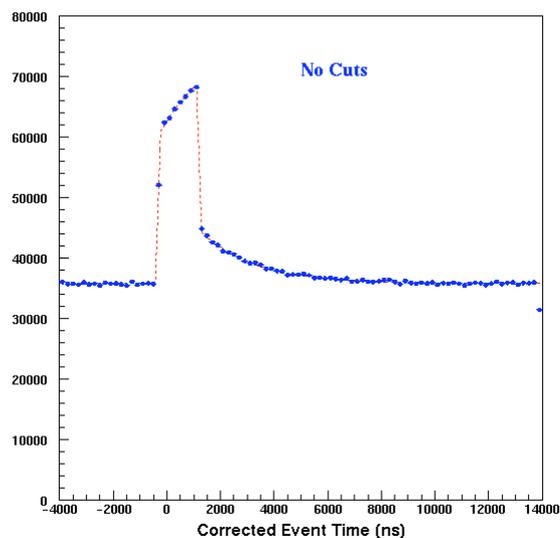
Most events are from  $\nu_\mu$  CC interactions, with characteristic two “sub-event” structure from stopped  $\mu$  decay.

$\nu_e$  CC interactions have 1 “sub-event”.



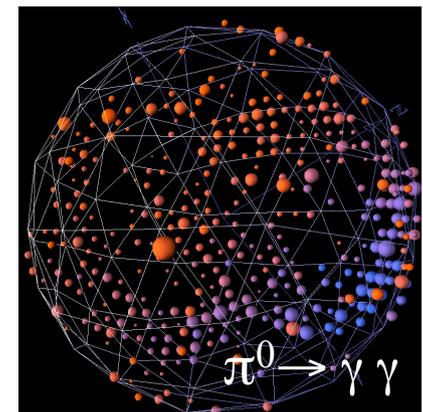
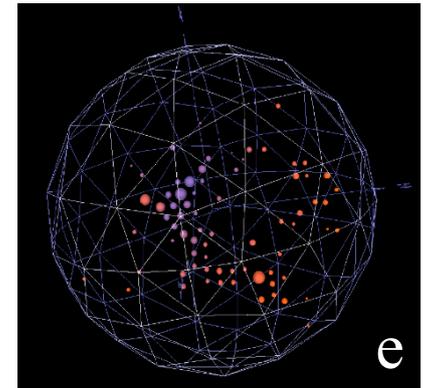
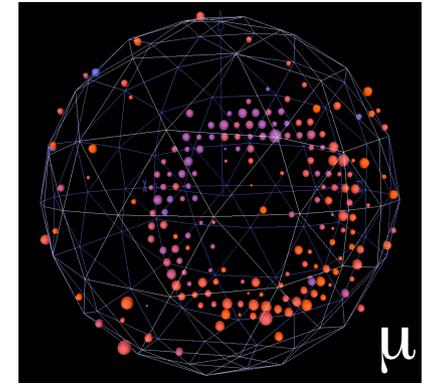
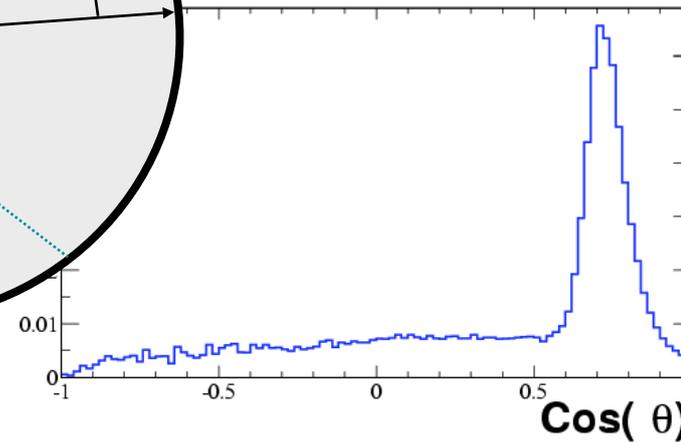
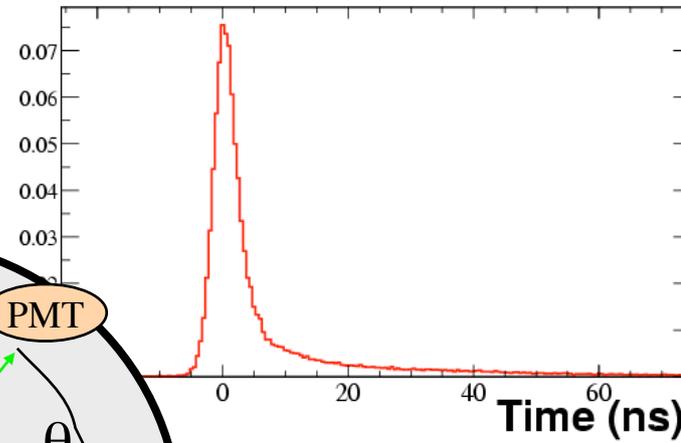
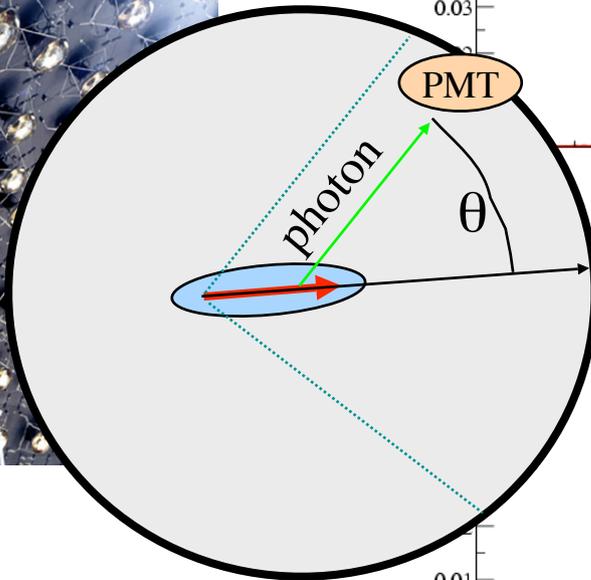
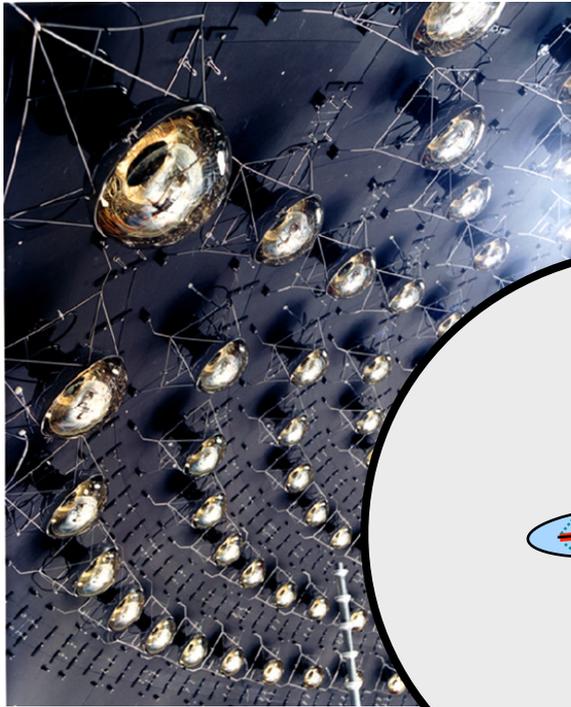
*Simple cuts eliminate cosmic ray events:*

1. Require  $< 6$  veto PMT hits,
2. Require  $> 200$  tank PMT hits.



P. Kasper

# MiniBooNE Detector: Reconstruction and Particle ID



## Reconstruction:

Fit time and angular distributions to find tracks

## Final State Particle Identification:

muons have sharp Cherenkov rings and long tracks

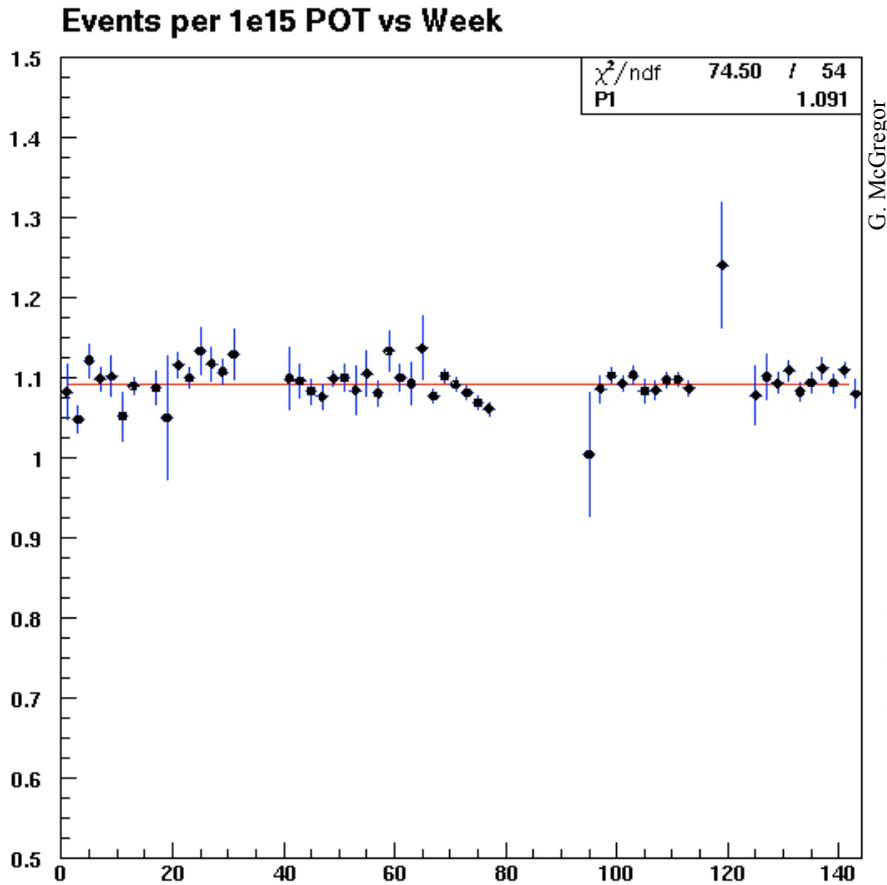
electrons have fuzzy rings, from multiple scattering, and short tracks

neutral pions decay to 2  $\gamma$ s, which convert and produce 2 fuzzy rings,

*easily misidentified as electrons if one ring gets lost!*

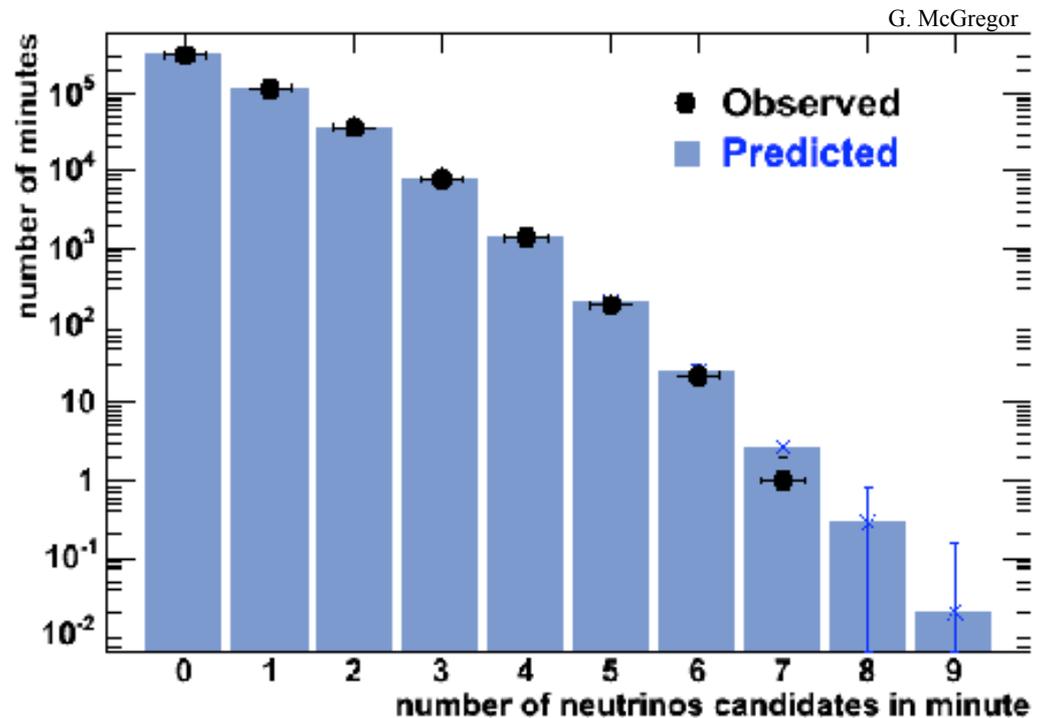
# MiniBooNE Beam & Detector: Stability

*Neutrinos per proton on target throughout the neutrino run:*



MiniBooNE observes  
~1 neutrino interaction  
per 1E15 protons.

*Observed and expected  
events per minute*



1. Motivation & Introduction
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  - Signal and Backgrounds
  - Strategy
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# Analysis Overview: Blind Analysis

*To avoid bias, MiniBooNE has done a blind analysis.*

**“Closed Box” Analysis**

To study the data, we defined specific event sets with  $< 1\sigma \nu_e$  signal for analysis.

## Initial Open Boxes

all non-beam-trigger data

0.25% random sample

$\nu_\mu$  CCQE

$\nu_\mu$  NC1 $\pi^0$

“dirt”

all events with  $E_\nu > 1.4$  GeV

$\nu_\mu$  CC1 $\pi^+$

$\nu_\mu$ -e elastic

## Use

calibration and MC tuning

an unbiased data set

measure flux,  $E_\nu^{QE}$ , oscillation fit

measure rate for MC

measure rate for MC

check MC rate

check MC rate

check MC rate

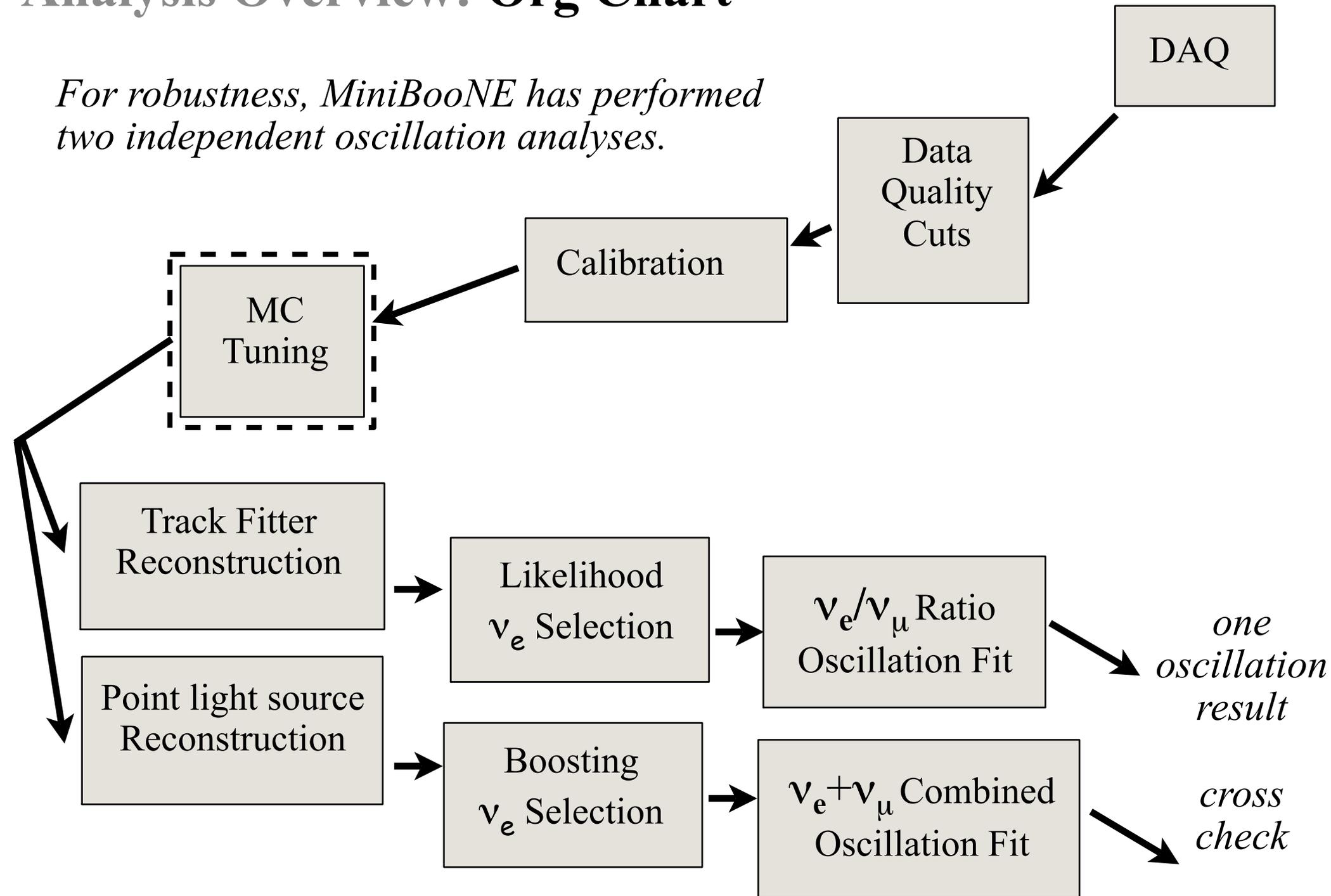
## Second Step:

One closed signal box

explicitly sequester the signal,  
99% of data open

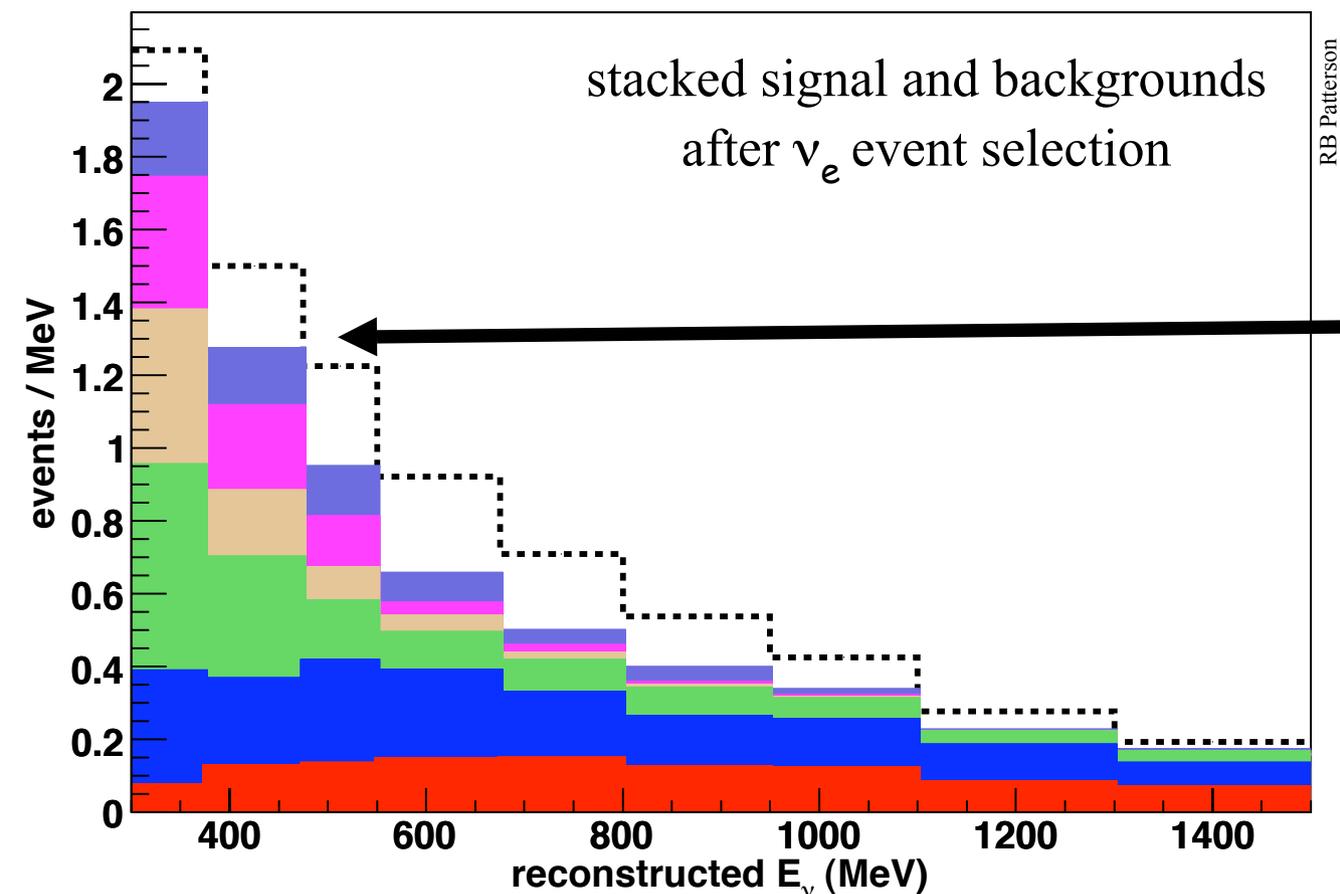
# Analysis Overview: Org Chart

*For robustness, MiniBooNE has performed two independent oscillation analyses.*



# Analysis Overview: Signal and Backgrounds

*what we predict for the full  $\nu$  data set ( $5.6E20$  protons on target):*



## Oscillation $\nu_e$

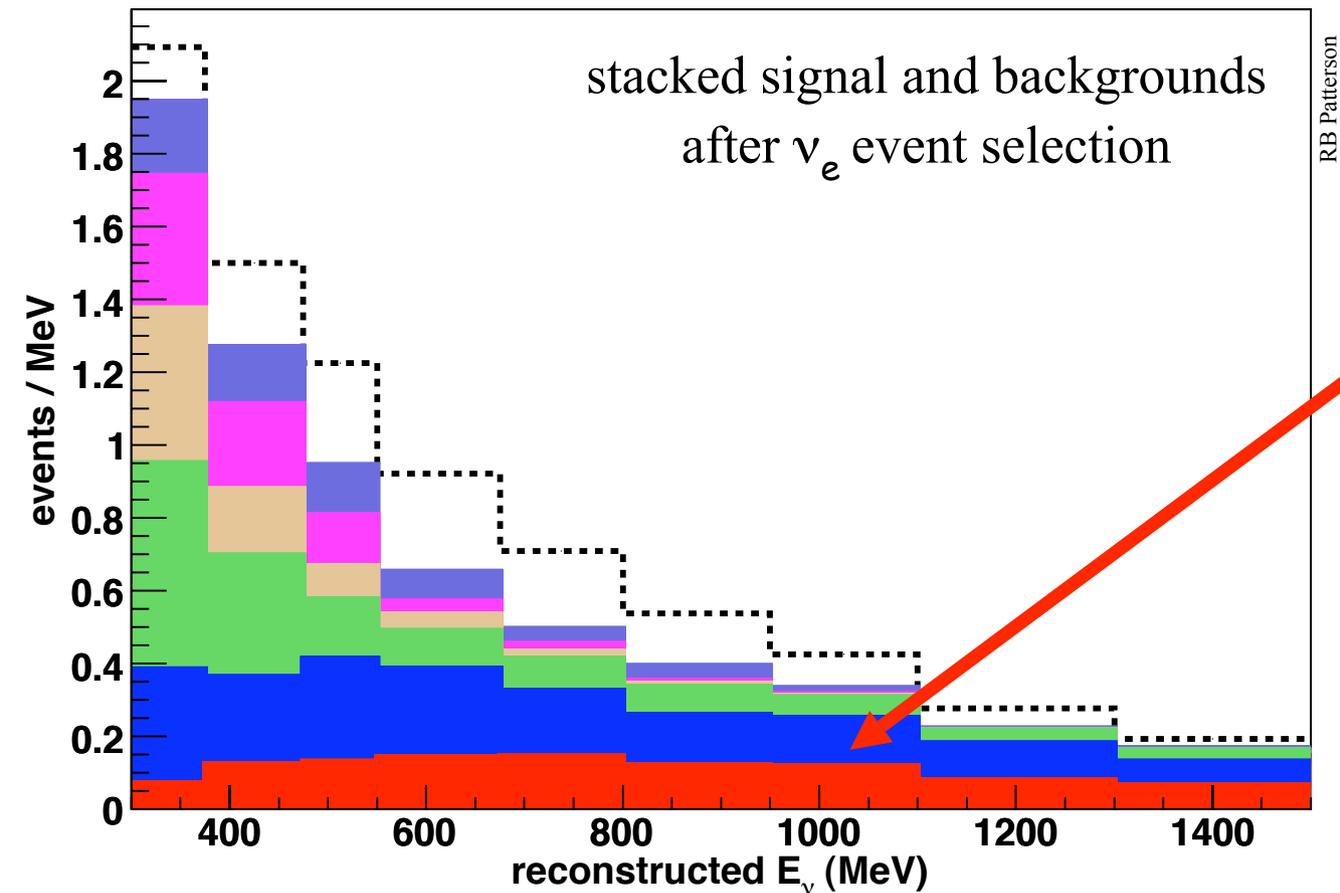
Example oscillation signal

- $\Delta m^2 = 1.2 \text{ eV}^2$
- $\text{SIN}^2 2\theta = 0.003$

Fit for excess as a function of  
reconstructed  $\nu_e$  energy

# Analysis Overview: Signal and Backgrounds

*what we predict for the full  $\nu$  data set ( $5.6E20$  protons on target):*



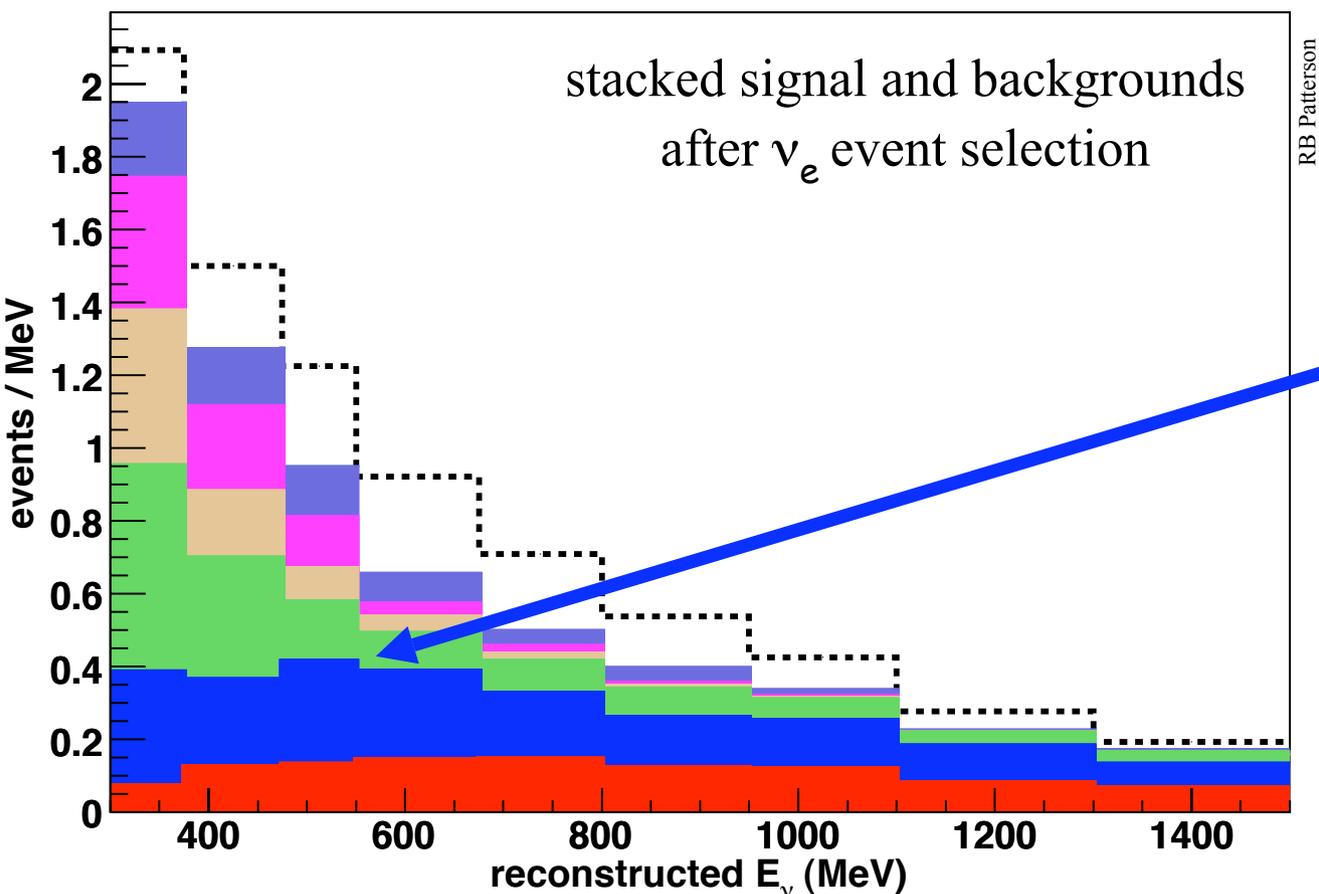
$\nu_e$  from  $K^+$  and  $K^0$

Use fit to kaon production data for shape

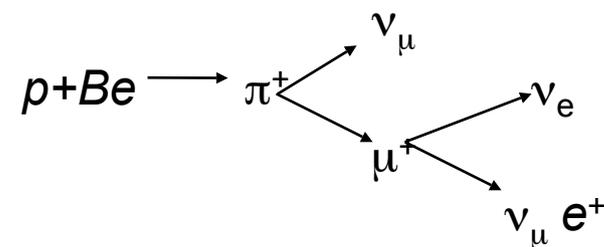
Use high energy  $\nu_e$  and  $\nu_\mu$  in-situ data for normalisation cross-check

# Analysis Overview: Signal and Backgrounds

what we predict for the full  $\nu$  data set ( $5.6E20$  protons on target):



$\nu_e$  from  $\mu^+$



Measured with in-situ  $\nu_\mu$   
CCQE sample

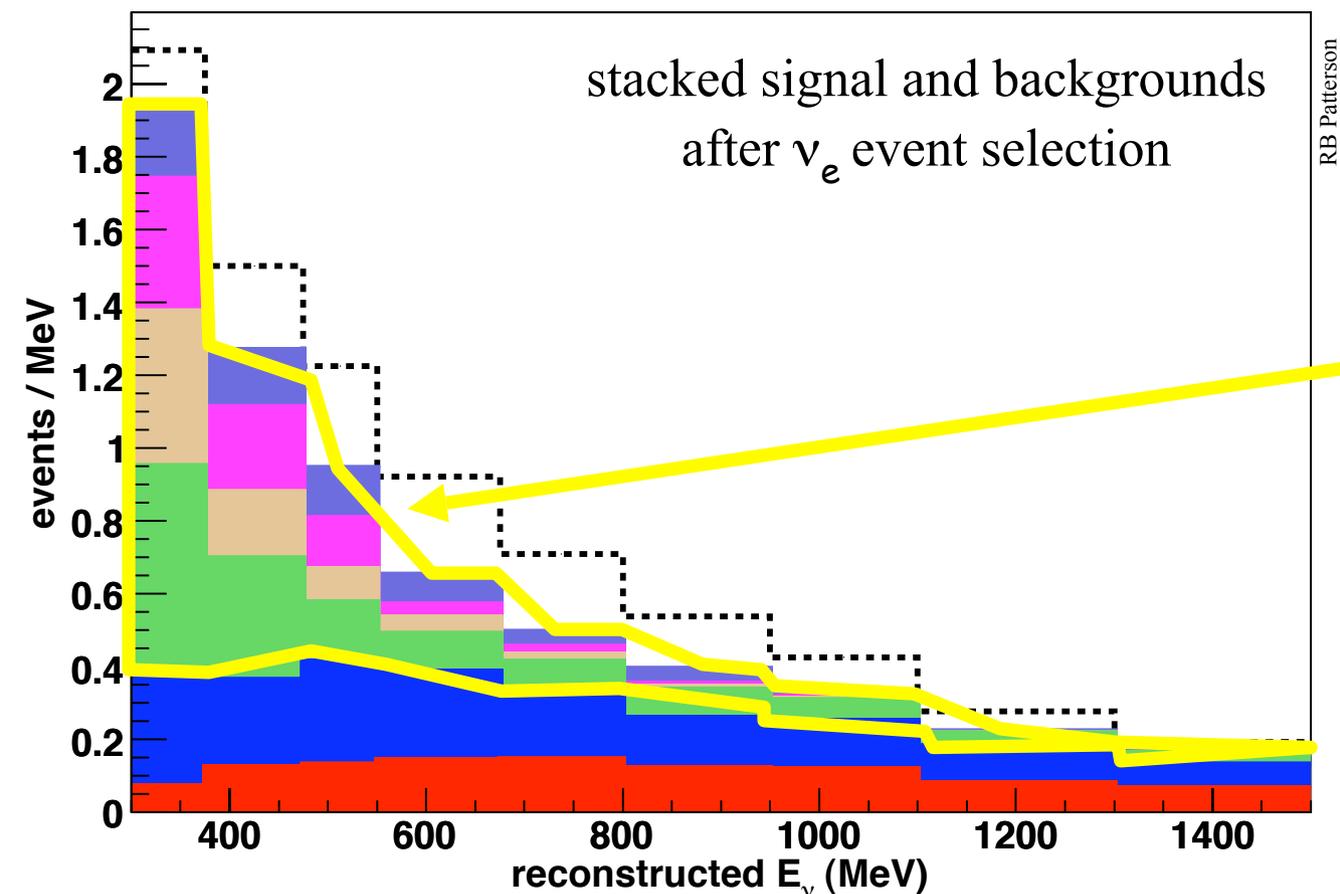
- Same ancestor  $\pi^+$  kinematics

Most important background

- Constrained to a few %

# Analysis Overview: Signal and Backgrounds

what we predict for the full  $\nu$  data set ( $5.6E20$  protons on target):



## MisID $\nu_\mu$

~46%  $\pi^0$

- Determined by clean  $\pi^0$  measurement

~14% “dirt”

- Measure rate to normalise and use MC for shape

~16%  $\Delta \gamma$  decay

- $\pi^0$  measurement constrains

~24% other

- Use  $\nu_\mu$  CCQE rate to normalise and MC for shape

# Analysis Overview: Strategy

*recurring theme: good data/MC agreement*

*in-situ data are incorporated wherever possible...*

(i) MC tuning with calibration data

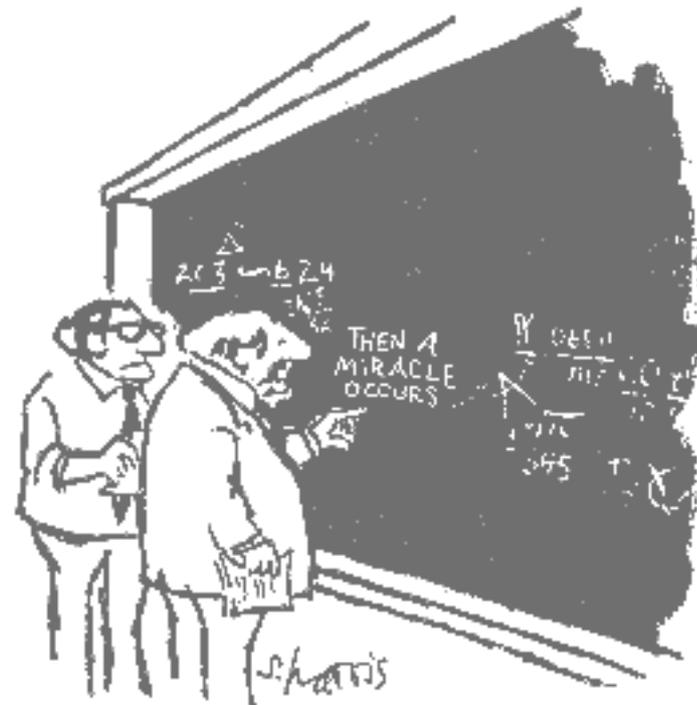
- energy scale
- PMT response
- optical model of light in the detector

(ii) MC fine-tuning with neutrino data

- cross section nuclear model parameters
- $\pi^0$  rate constraint

(iii) constraining systematic errors with neutrino data

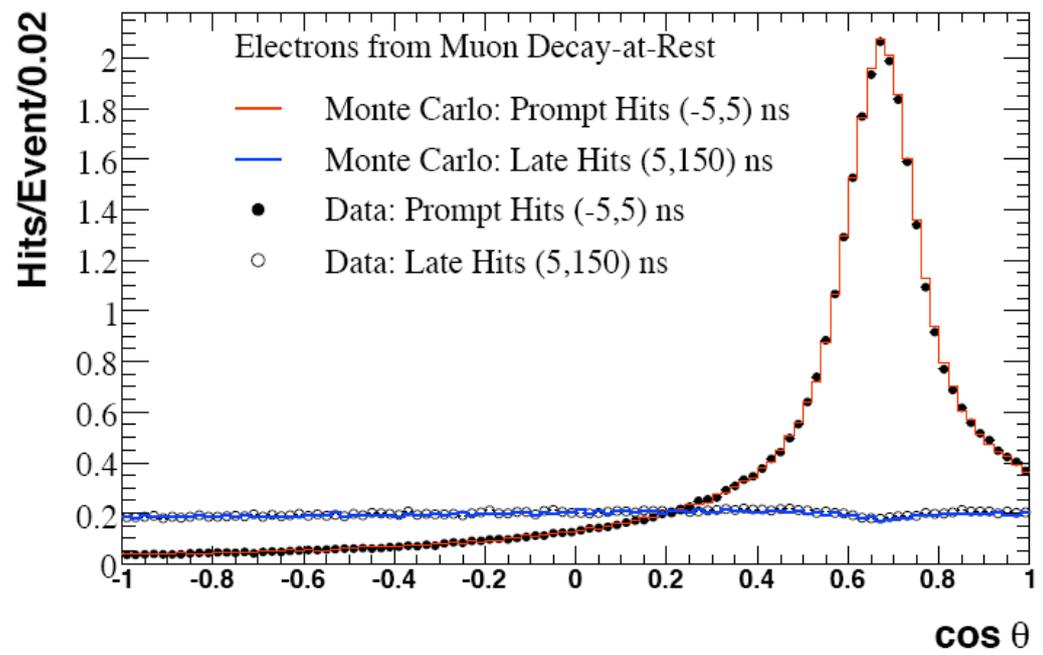
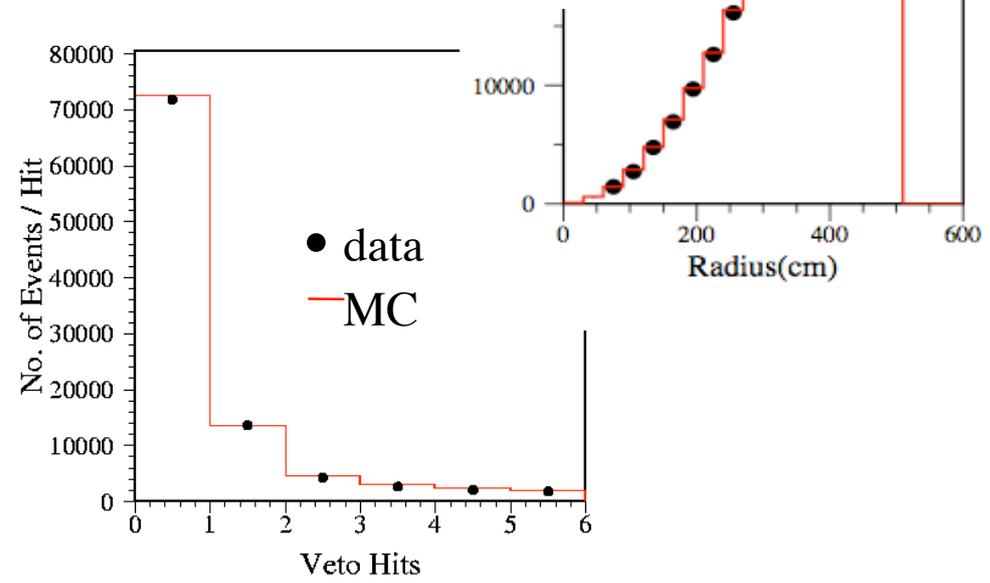
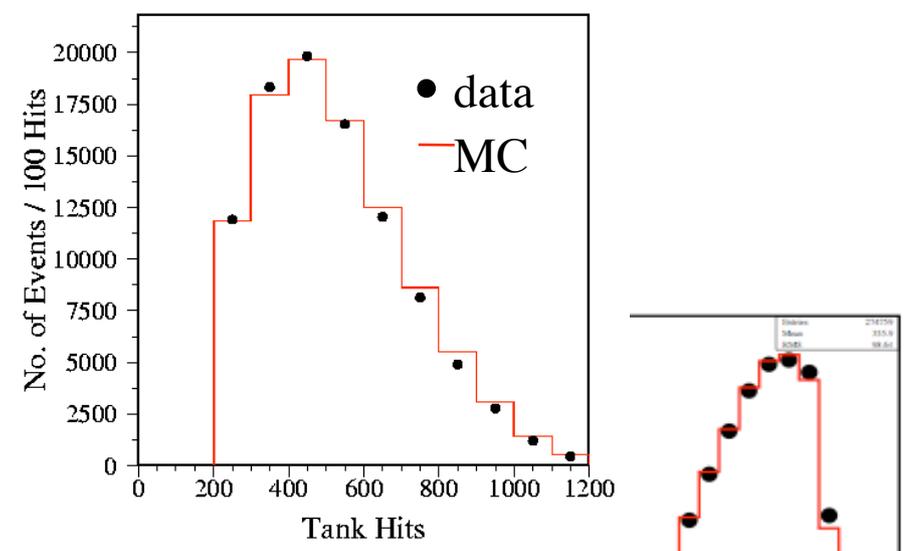
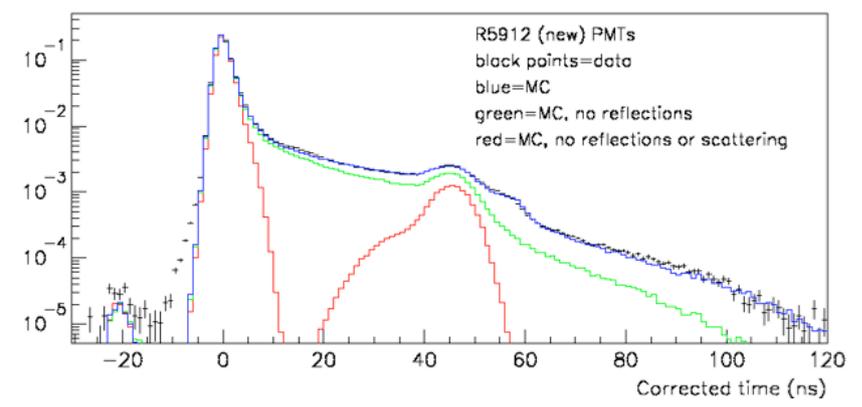
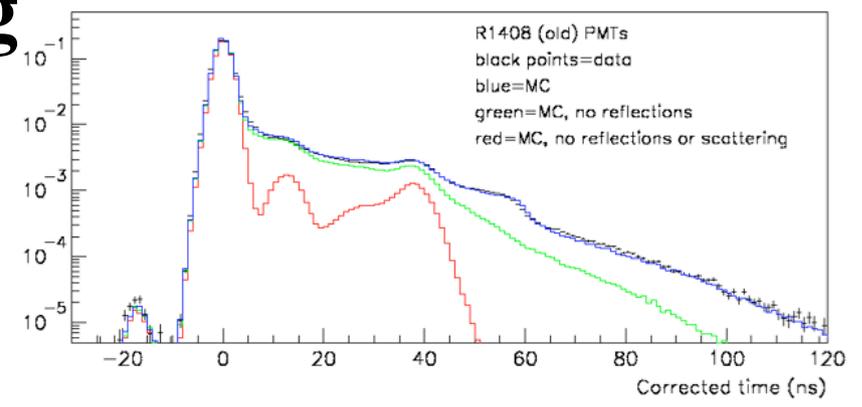
- ratio method example:  $\nu_e$  from  $\mu$  decay background
- combined oscillation fit to  $\nu_\mu$  and  $\nu_e$  data



"I think you should be more explicit here in step two."

# Analysis Overview: MC Tuning

## MC tuning with calibration data



# Analysis Overview: Strategy

*in-situ data are incorporated wherever possible...*

(i) MC tuning with calibration data

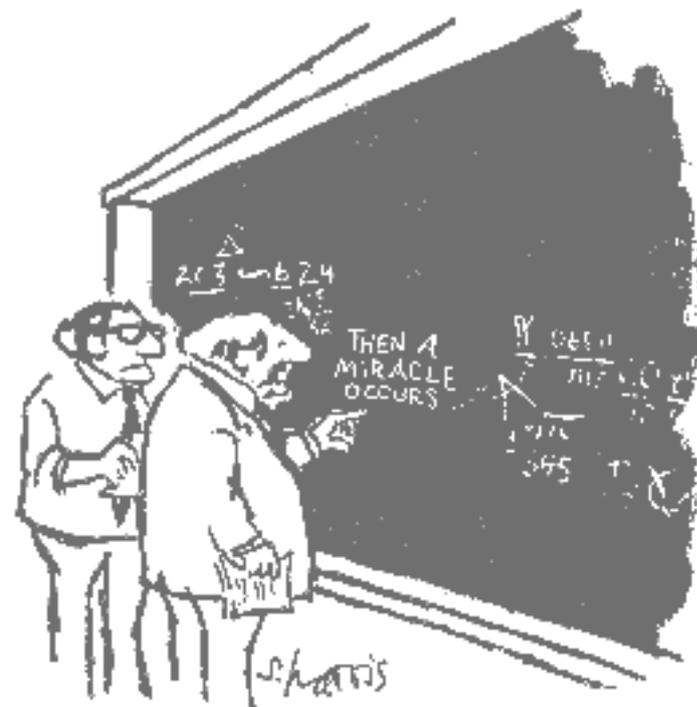
- energy scale
- PMT response
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- cross section nuclear model parameters
- $\pi^0$  rate constraint

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- ratio method example:  $\nu_e$  from  $\mu$  decay background
- combined oscillation fit to  $\nu_\mu$  and  $\nu_e$  data

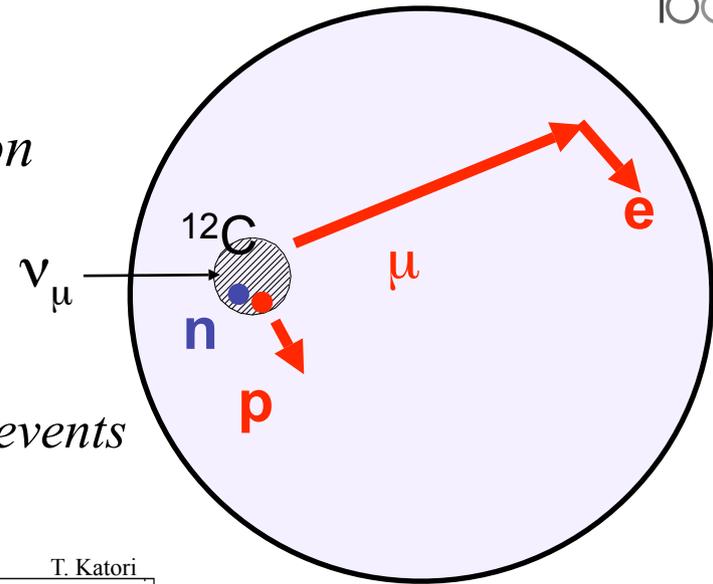


"I think you should be more explicit here in step two."

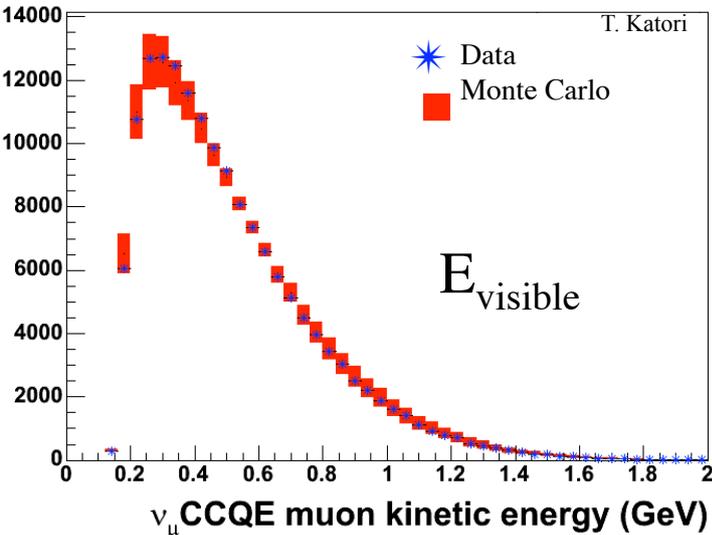
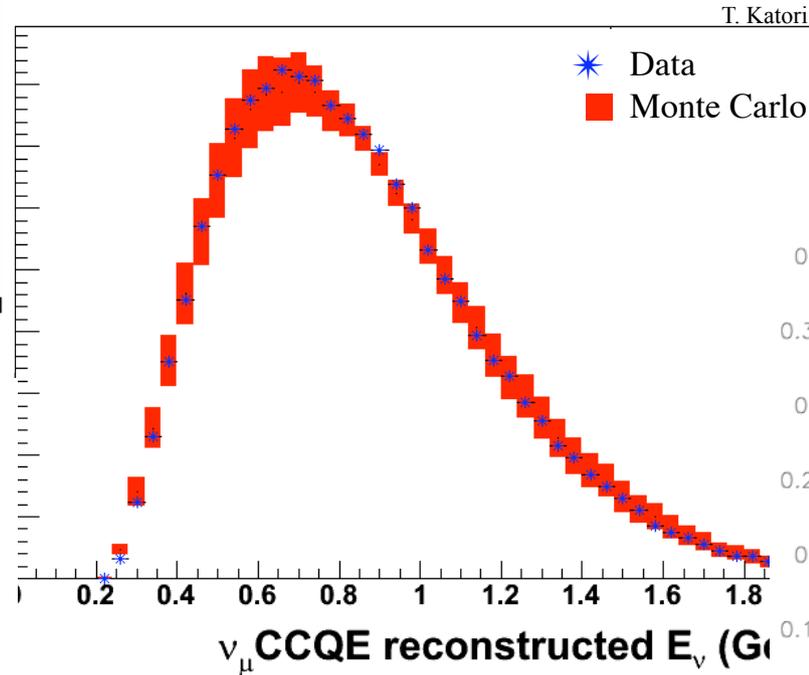
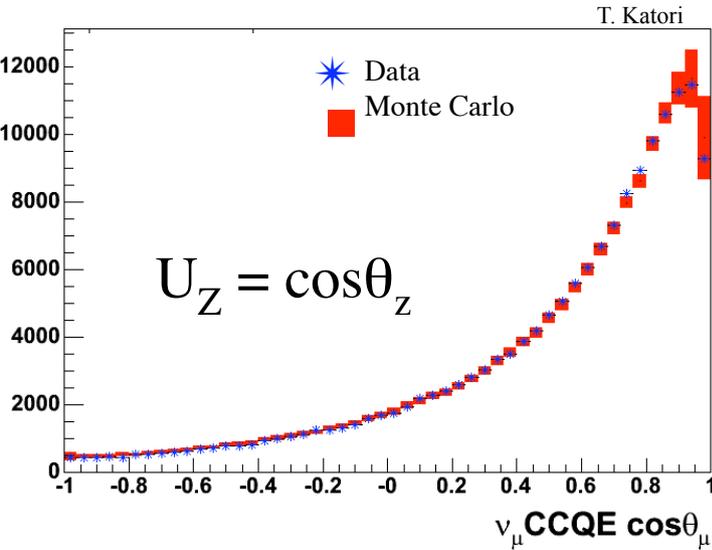
# Analysis Strategy: $\nu_\mu$ CCQE Events

used to measure the  $\nu_\mu$  flux and check  $E_\nu^{QE}$  reconstruction

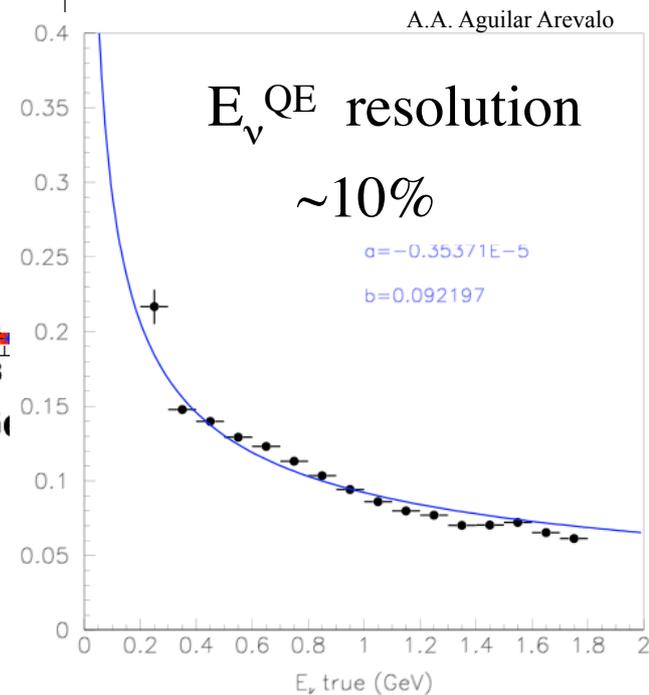
1. tag muons by requiring 2 sub-events in time
2. require reconstructed distance between sub-events  $< 1m$



$\sim 74\%$  CCQE purity,  $\sim 190k$  events

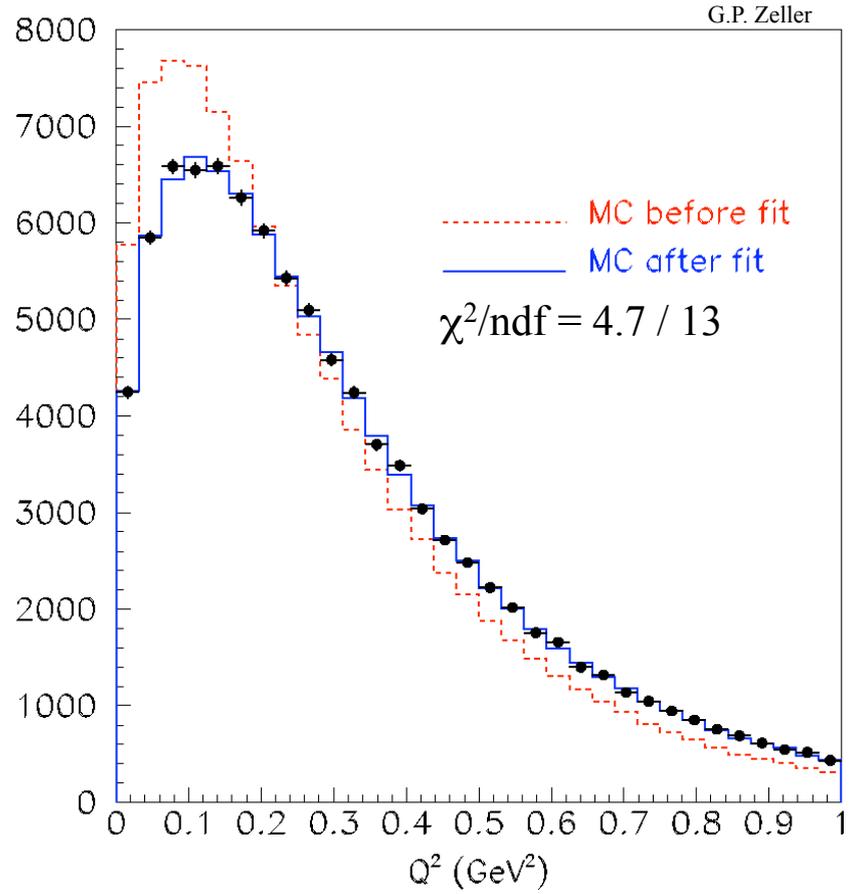


$$E_\nu^{QE} = \frac{1}{2} \frac{2M_p E_\mu - m_\mu^2}{M_p - E_\mu + \sqrt{(E_\mu^2 - m_\mu^2)} \cos\theta_\mu}$$

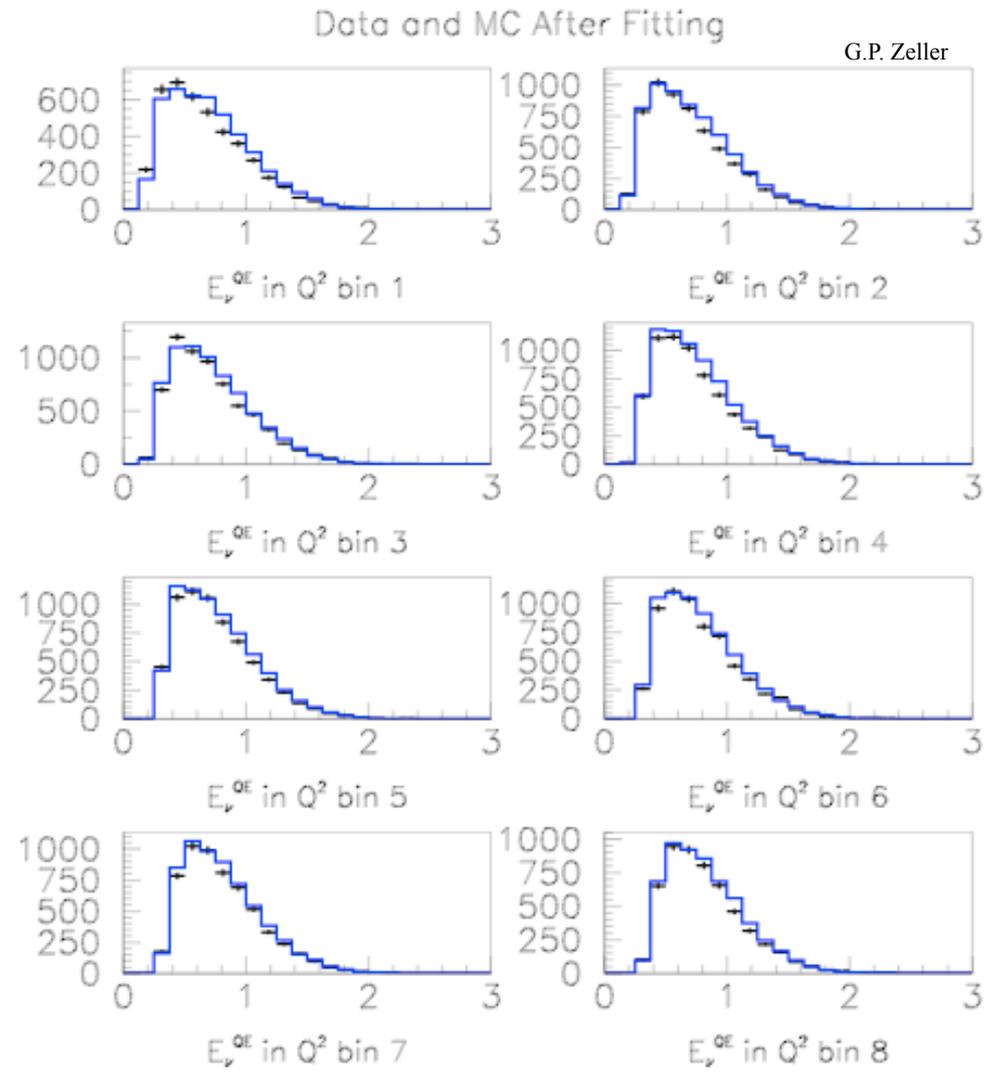


# Incorporating $\nu_\mu$ Data: CCQE Cross Section

The  $\nu_\mu$  CCQE data  $Q^2$  distribution is fit to tune empirical parameters of the nuclear model ( $^{12}\text{C}$  target)



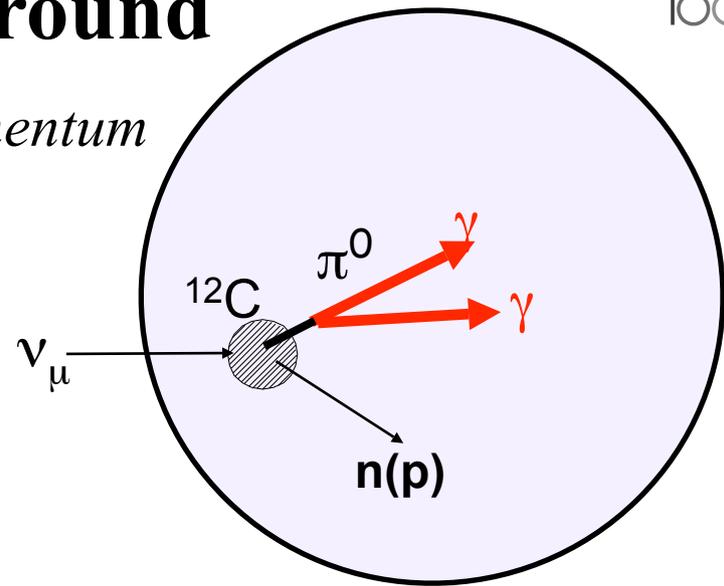
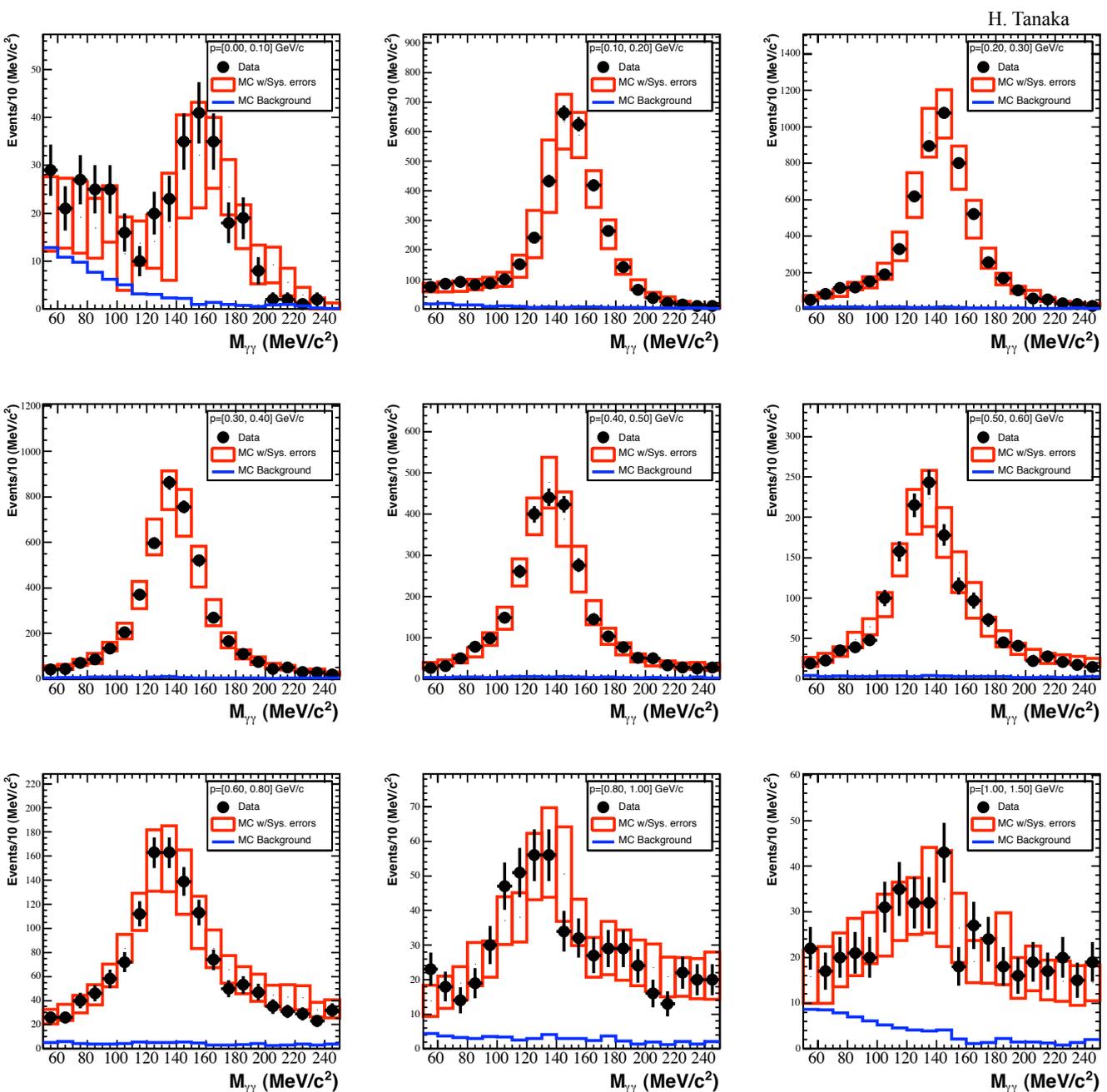
this results in good data-MC agreement for variables **not** used in tuning



the tuned model is used for both  $\nu_\mu$  and  $\nu_e$  CCQE

# Analysis Strategy: $\pi^0$ Mis-ID Background

*clean  $\pi^0$  events are used to tune the MC rate vs.  $\pi^0$  momentum*

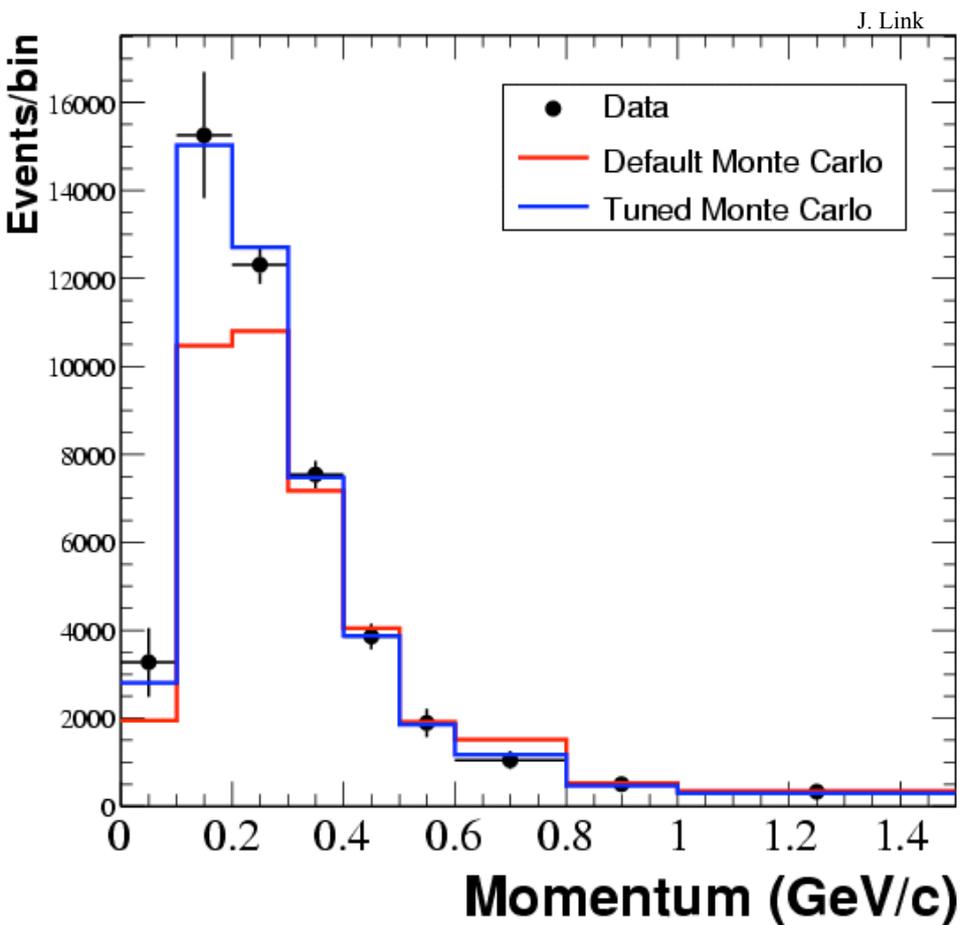


$\pi^0$  events can reconstruct outside of the mass peak when:

1. asymmetric decays fake 1 ring
2. 1 of the 2 photons exits the detector
3. high momentum  $\pi^0$  decays produce overlapping rings

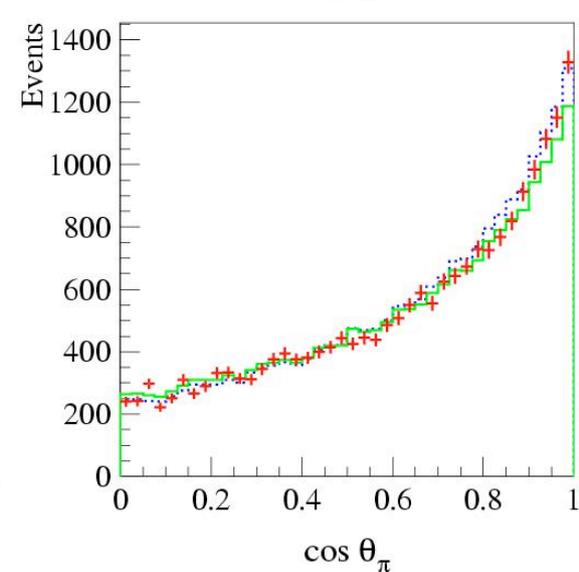
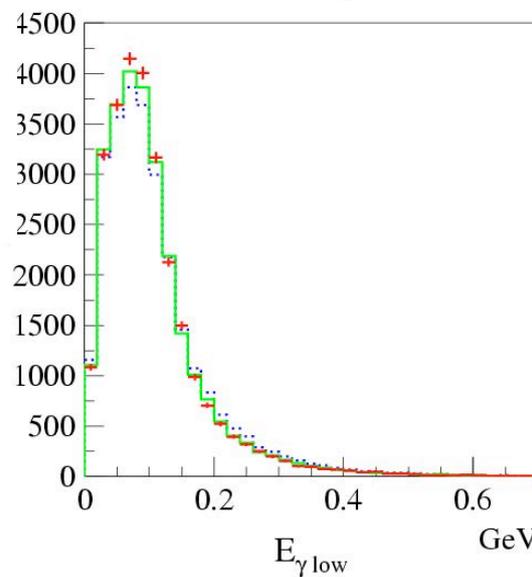
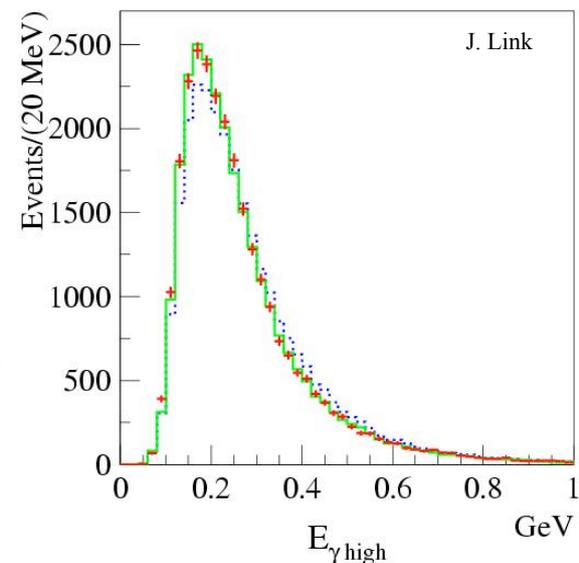
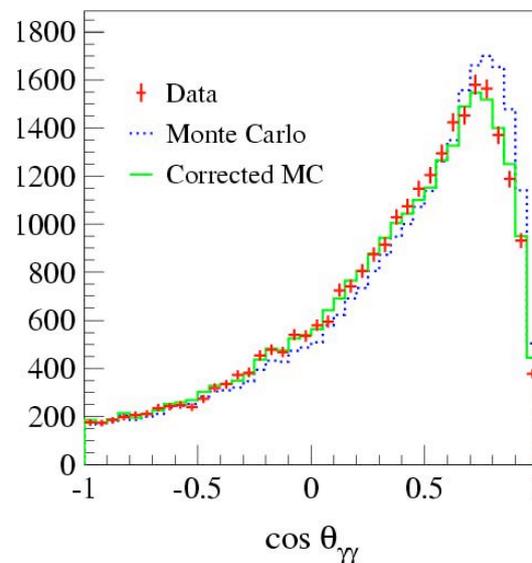
# Analysis Strategy: $\pi^0$ Mis-ID Background

The MC  $\pi^0$  rate (flux  $\times$  xsec) is re-weighted to match the measurement in  $p_\pi$  bins.



*Because this constrains the  $\Delta$  resonance rate, it also constrains the rate of  $\Delta \rightarrow N\gamma$  in MiniBooNE*

this procedure results in good data-MC agreement for variables **not** used in tuning



# Analysis Overview: Strategy

*in-situ data is incorporated wherever possible...*

(i) MC tuning with calibration data

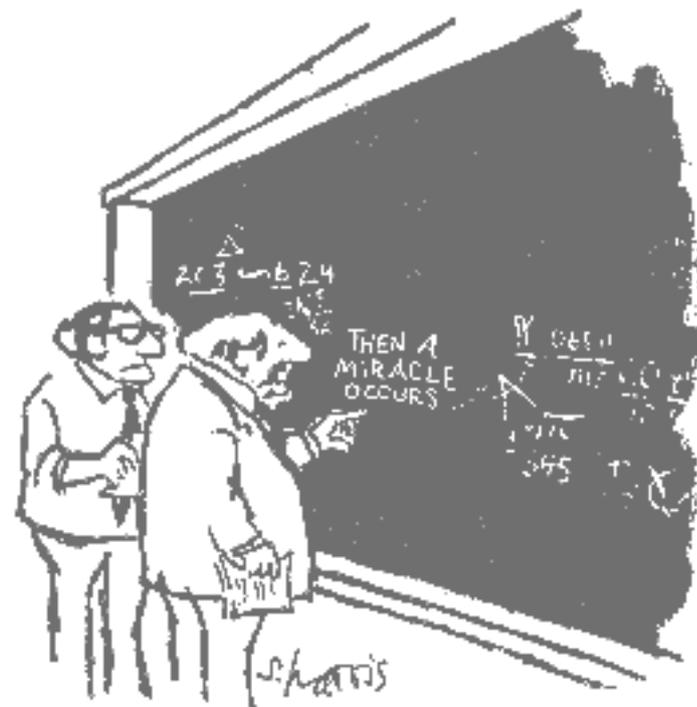
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- cross section nuclear model parameters
- $\pi^0$  rate constraint

(iii) constraining systematic errors with neutrino data

- ratio method example:  $\nu_e$  from  $\mu$  decay background
- combined oscillation fit to  $\nu_\mu$  and  $\nu_e$  data



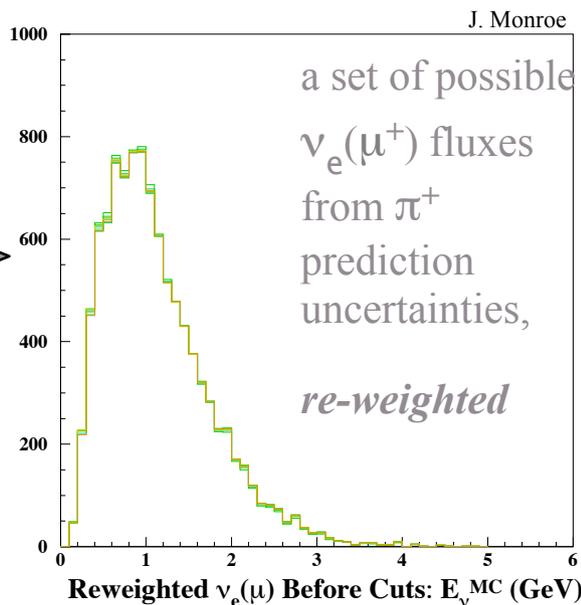
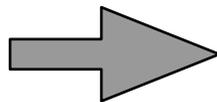
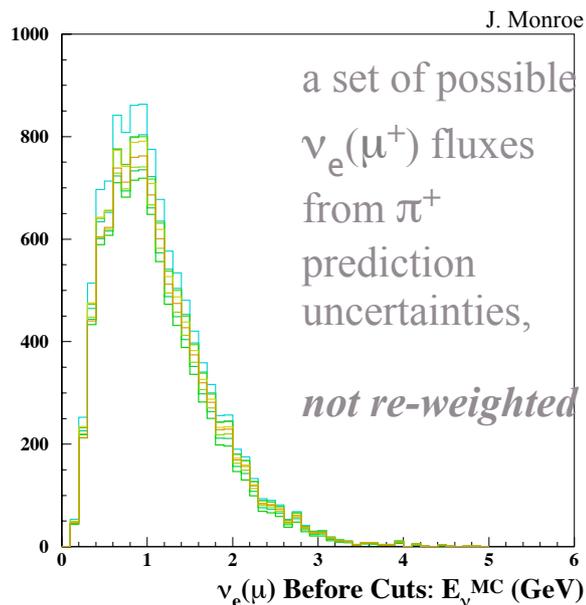
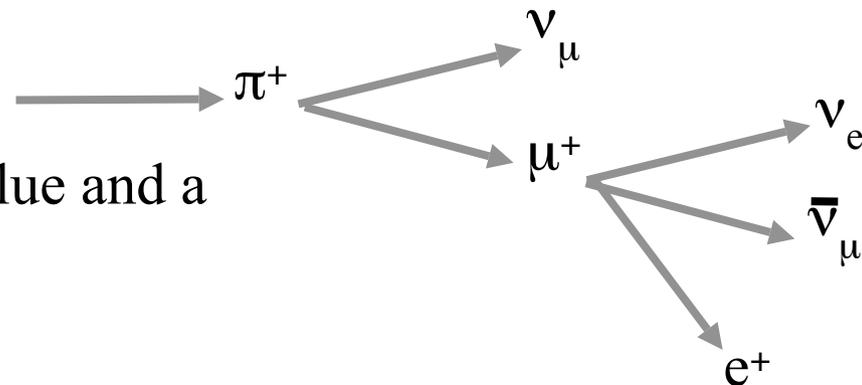
"I think you should be more explicit here in step two."

# Analysis Strategy 1: Ratio Method

Example:  $\nu_\mu$  CCQE events measure  $\pi^+$  spectrum, constrain  $\mu^+$ -decay  $\nu_e$  flux

## Ratio Method Constraint:

1. MC based on external data predicts a central value and a range of possible  $\nu_\mu(\pi)$  fluxes
2. make Data/MC ratio vs.  $E_\nu^{\text{QE}}$  for  $\nu_\mu$  CCQE data
3. re-weight each possible MC parent- $\pi^+$  flux by the ratio (2), including sister  $\mu^+$  & niece  $\nu_e$



reduction in the spread of possible fluxes translates directly into a reduction in the  $\mu^+$ -decay  $\nu_e$  background uncertainty

*Can use ratio method to constrain most BG sources*

# Analysis Strategy 2: Combined Fit

Fit the  $E_\nu^{QE}$  distributions of  $\nu_e$  and  $\nu_\mu$  events for oscillations, together

Raster scan in  $\Delta m^2$ , and  $\sin^2 2\theta_{\mu e}$  ( $\sin^2 2\theta_{\mu x} = 0$ ),  
calculate  $\chi^2$  value over  $\nu_e$  **and**  $\nu_\mu$  bins

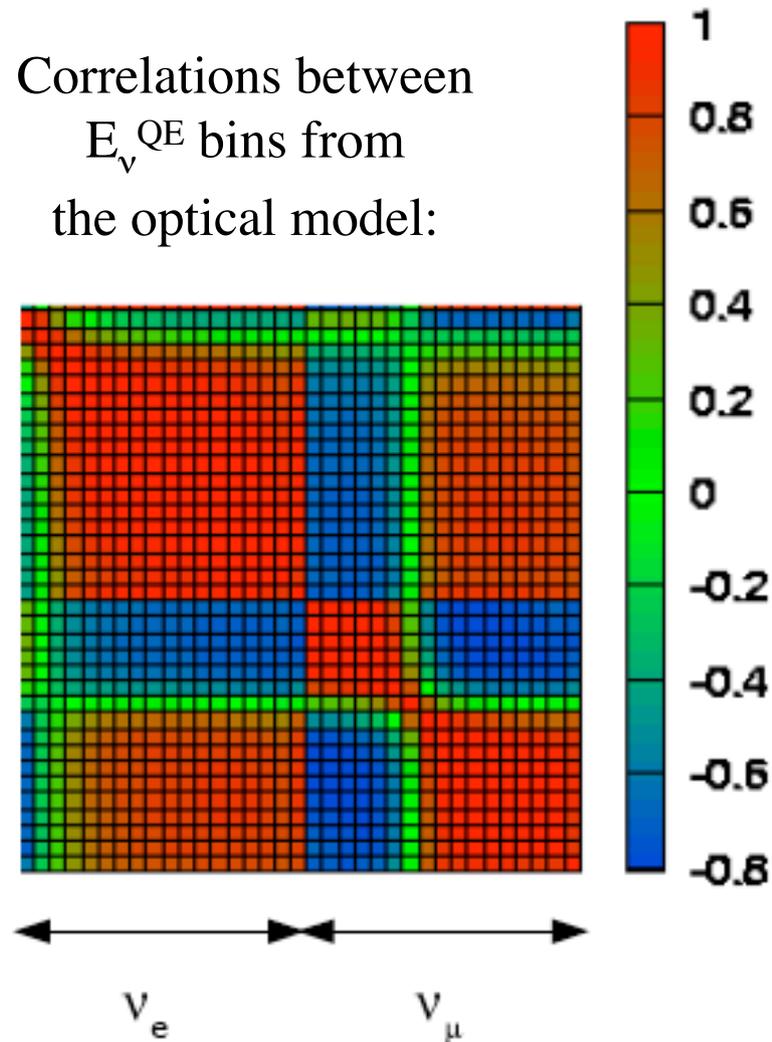
$$\chi^2 = \sum_{i=1}^{N_{bins}} \sum_{j=1}^{N_{bins}} (m_i - t_i) \mathcal{M}_{ij}^{-1} (m_j - t_j)$$

In this case, systematic error matrix  $\mathcal{M}_{ij}$  includes predicted uncertainties for  $\nu_e$  **and**  $\nu_\mu$  bins

$$M_{ij} = \begin{pmatrix} \nu_e & \nu_e \nu_\mu \\ \nu_\mu \nu_e & \nu_\mu \end{pmatrix}$$

Left: example,  $m_i$  = "fake data" = MC with no oscillations

*a combined fit constrains uncertainties in common*



# Analysis Strategy: Error Matrix

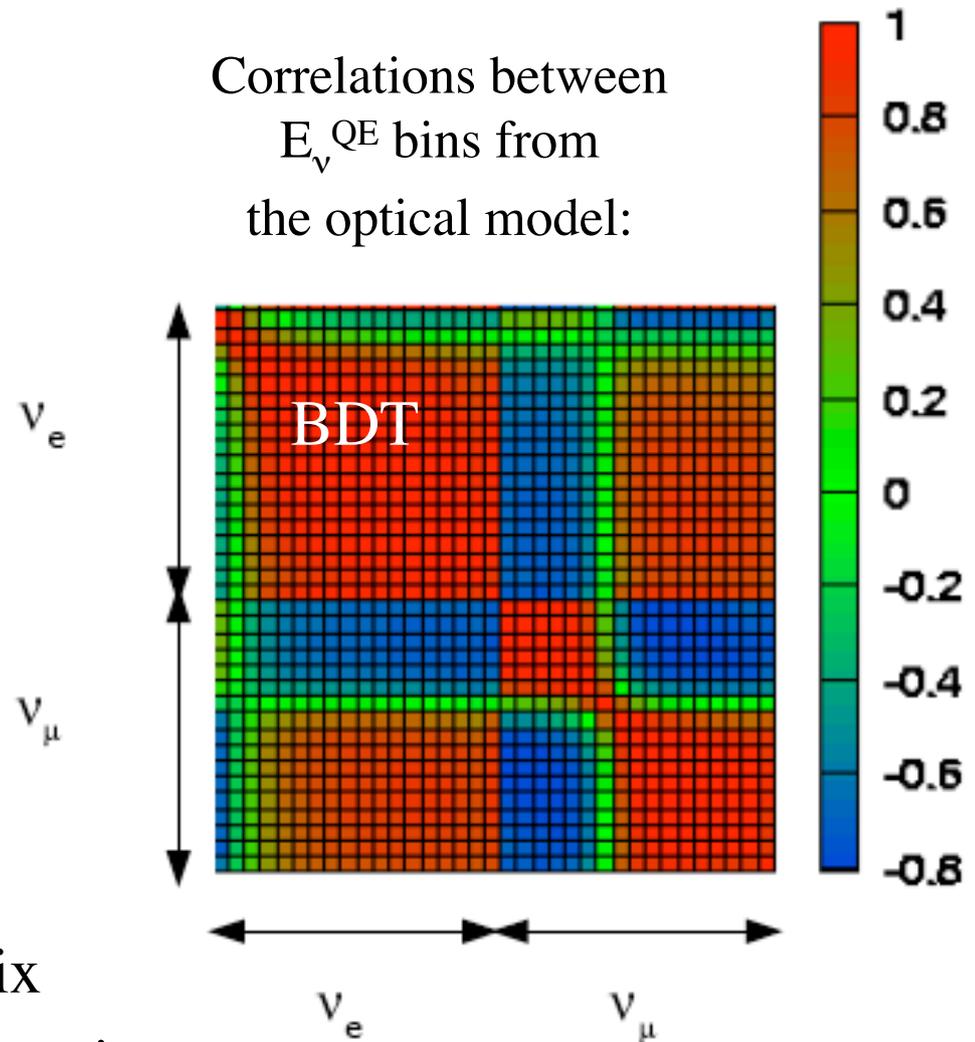
$$E_{ij} \approx \frac{1}{M} \sum_{\alpha=1}^M \left( N_i^\alpha - N_i^{MC} \right) \left( N_j^\alpha - N_j^{MC} \right)$$

- N is number of events passing cuts
- MC is standard Monte Carlo
- $\alpha$  represents a different MC draw  
(called a “multisim”)
- M is the total number of MC draws
- i,j are  $E_\nu^{QE}$  bins

Total error matrix  
is sum from each source.

Primary (TB):  $\nu_e$ -only total error matrix

Cross-check (BDT):  $\nu_\mu$ - $\nu_e$  total error matrix



# Analysis Overview: Systematic Errors

*A long list of systematic uncertainties are estimated using Monte Carlo:*

## neutrino flux predictions

- $\pi^+$ ,  $\pi^-$ ,  $K^+$ ,  $K^-$ ,  $K^0$ , n, and p total and differential cross sections
- secondary interactions of mesons
- focusing horn current
- target + horn system alignment



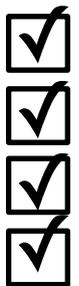
## neutrino interaction cross section predictions

- nuclear model
- rates and kinematics for relevant exclusive processes
- resonance width and branching fractions



## detector modelling

- optical model of light propagation in oil (39 parameters!)
- PMT charge and time response
- electronics response
- neutrino interactions in dirt surrounding detector hall



*✓ Most are constrained or checked using in-situ MiniBooNE data.*

1. Motivation & Introduction
2. Description of the Experiment
3. Analysis Overview
4. Two Independent Oscillation Searches
  - Reconstruction and Event Selection
  - Systematic Uncertainties
5. First Results

# Two Independent Oscillation Searches: Methods

## Method 1: Track-Based Analysis

Primary analysis

- *Use careful reconstruction of particle tracks*
- *Identify particle type by likelihood ratio*
- *Use ratio method to constrain backgrounds*

### Strengths:

Relatively insensitive to optical model

Simple cut-based approach with likelihoods

## Method 2: Boosted Decision Trees

Independent cross-check

- *Classify events using “boosted decision trees”*
- *Apply cuts on output variables to improve separation of event types*
- *Use combined fit to constrain backgrounds*

### Strengths:

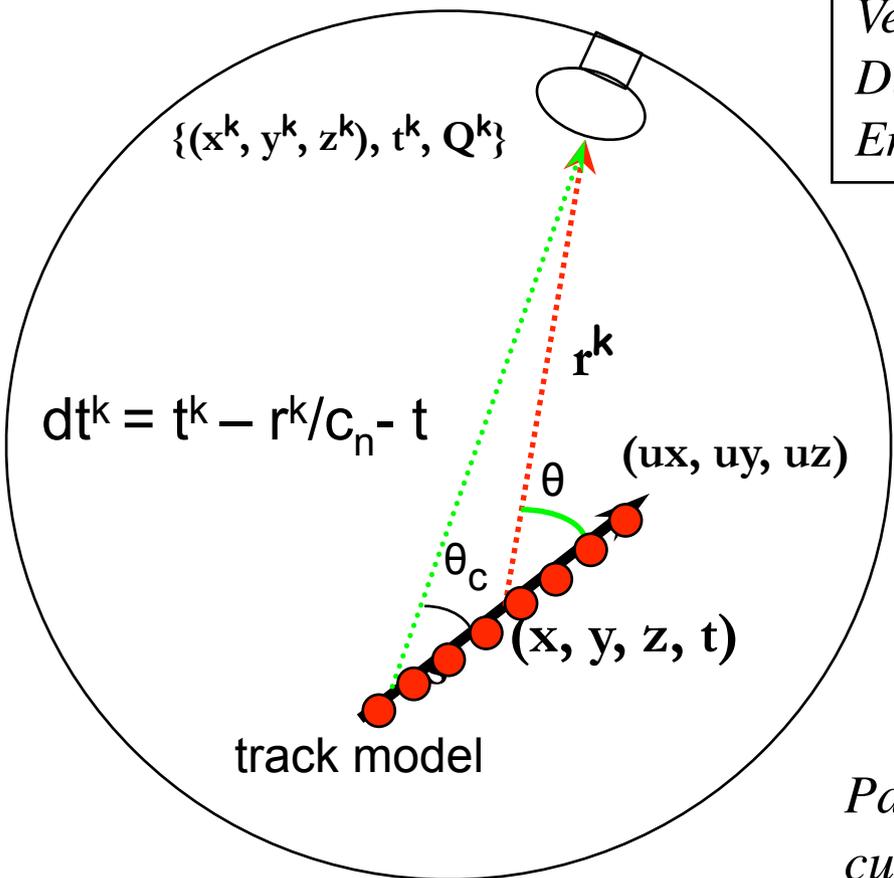
Combination of many weak variables form strong classifier

Better constraints on background events

# Method 1: Track-Based Analysis

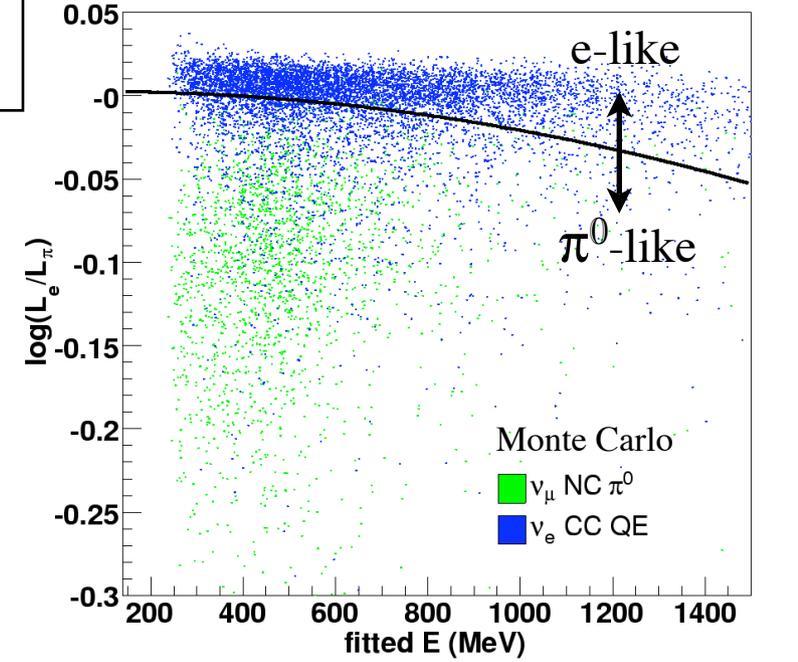
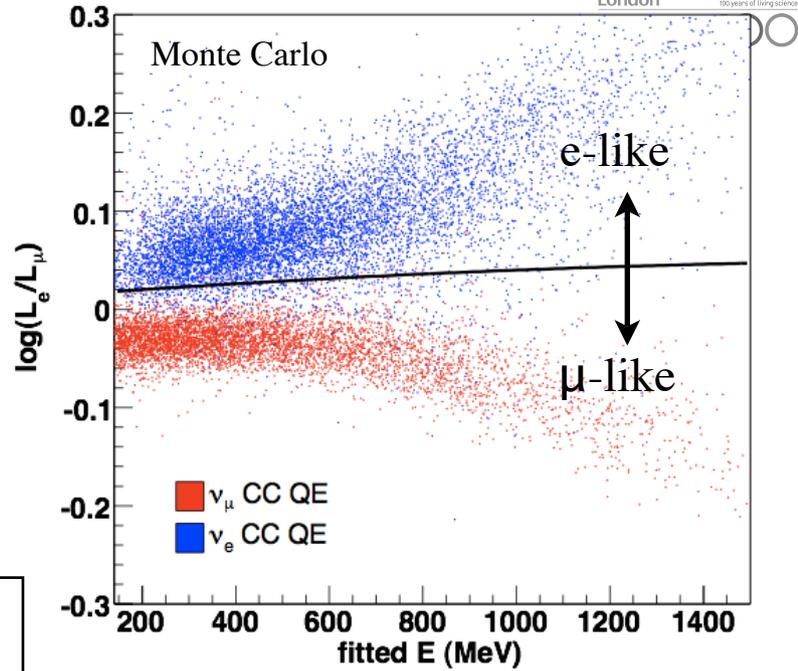
Reconstruction fits an extended light source with 7 parameters: vertex, direction ( $\theta, \phi$ ), time, energy

Fit events under 3 possible hypotheses:  $\mu$ -like, e-like, two track ( $\pi^0$ -like)



Fitter resolution

Vertex:	22 cm
Direction:	2.8°
Energy:	11%



Particle ID relies on likelihood ratio cuts to select  $\nu_e$ , cuts chosen to maximise sensitivity to  $\nu_\mu \rightarrow \nu_e$  oscillation

# Track-Based Analysis: $e/\mu$ Likelihood

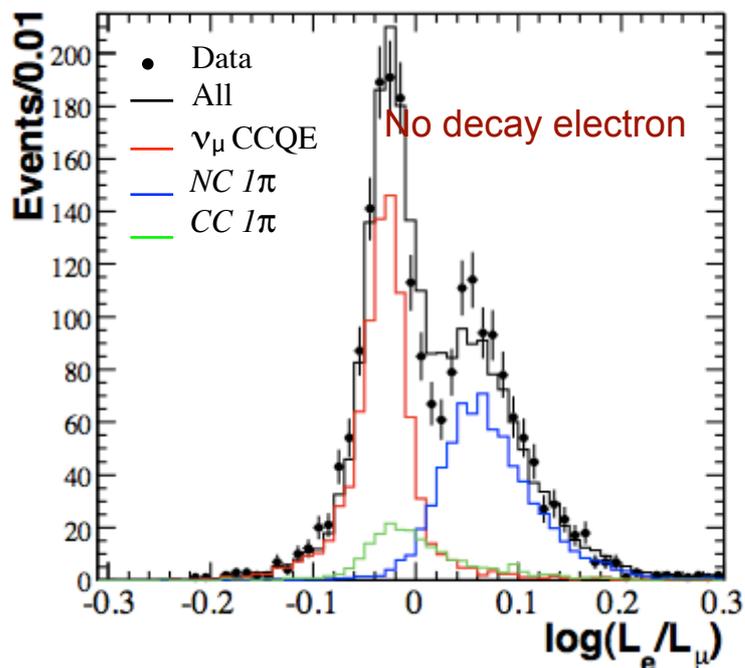
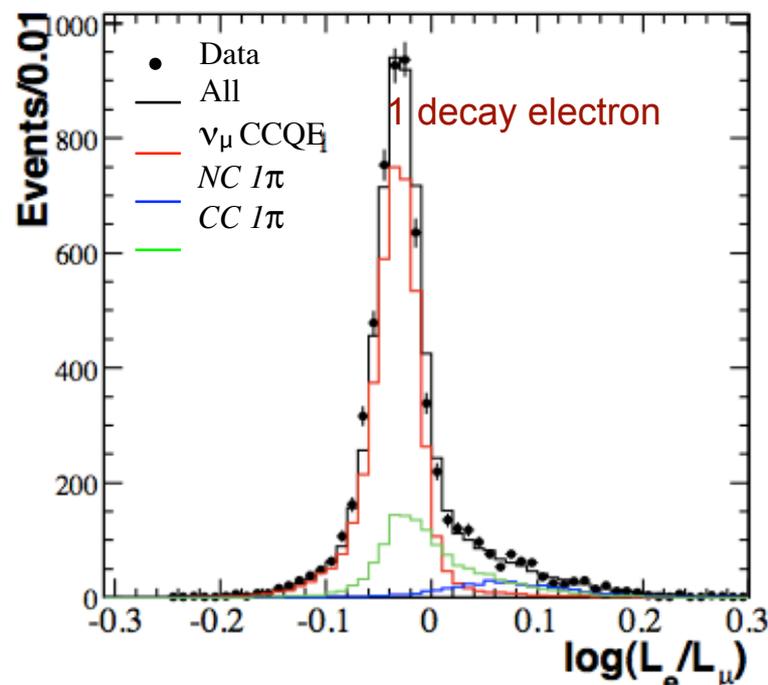
Test  $\mu$ - $e$  separation on data:

## $\nu_\mu$ CCQE data sample

Pre-selection cuts

Fiducial volume: ( $R < 500$  cm)

2 subevents: muon + decay electron



## “All-but-signal” data sample

Pre-selection cuts

Fiducial volume: ( $R < 500$  cm)

1 subevent: 8% of muons capture on  $^{12}\text{C}$

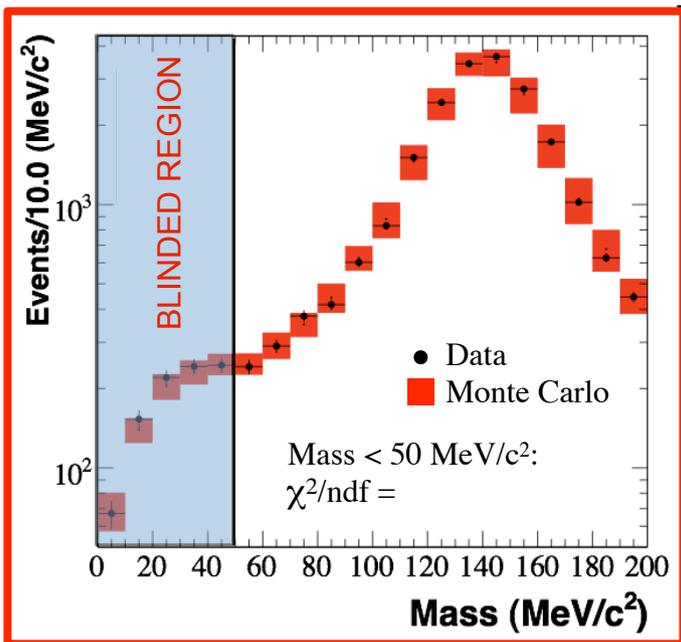
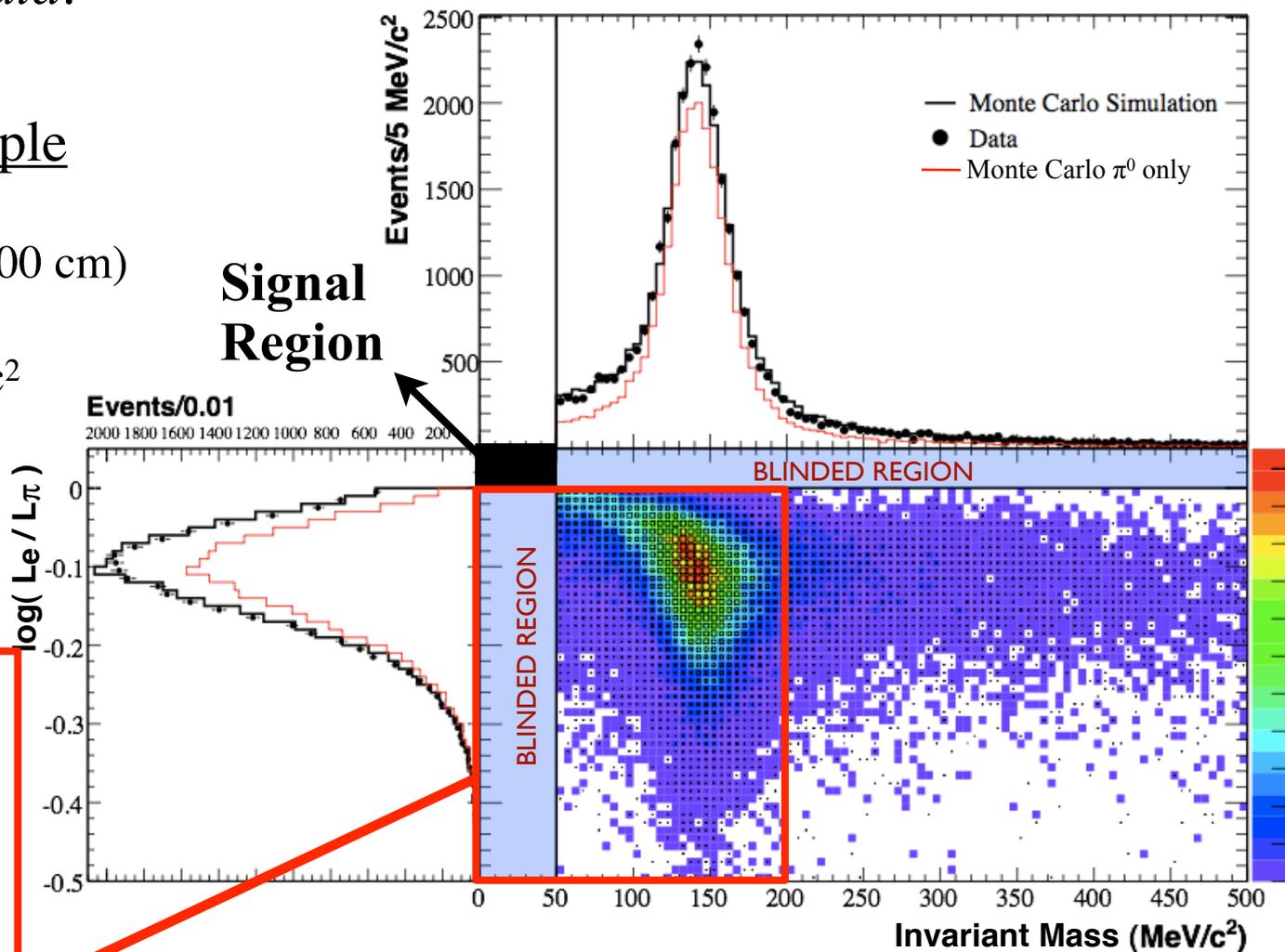
*Events with  $\log(L_e/L_\mu) > 0$  ( $e$ -like) undergo additional fit with two-track hypothesis.*

# Track-Based Analysis: $e/\pi^0$ Likelihood

Test  $e-\pi^0$  separation on data:

## “All-but-signal” data sample

- Pre-selection cuts
- Fiducial volume cut ( $R < 500$  cm)
- 1 subevent
- Invariant mass  $> 50$  MeV/c<sup>2</sup>
- $\log(L_e / L_\pi) < 0$  ( $\pi$ -like)



## Tighter selection cuts:

- Invariant mass  $< 200$  MeV/c<sup>2</sup>
- $\log(L_e / L_\mu) > 0$  (e-like)
- $\log(L_e / L_\pi) < 0$  ( $\pi$ -like)

# Method 2: Boosted Decision Trees

**Decision Trees:** A machine-learning technique which tries to recover signal events that would be eliminated in cut-based analyses.

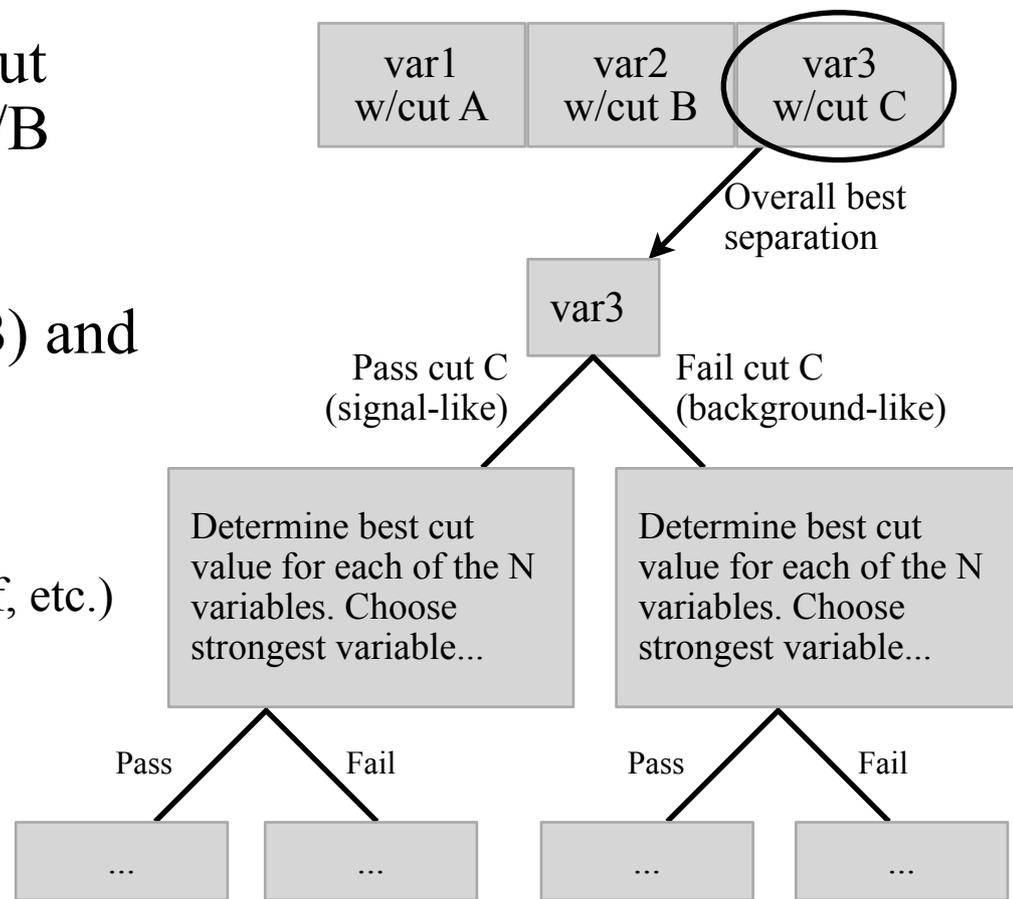
## *Training a decision tree:*

For a set of  $N$  variables, determine the cut value for each variable that gives best S/B separation.

Cut on the best variable (i.e. highest S/B) and repeat.

*Final score:* For each leaf,

- 1 for correct events (signal event on a signal leaf, etc.)
- +1 for incorrect events



**Boosting:** Increase weight of misclassified events.

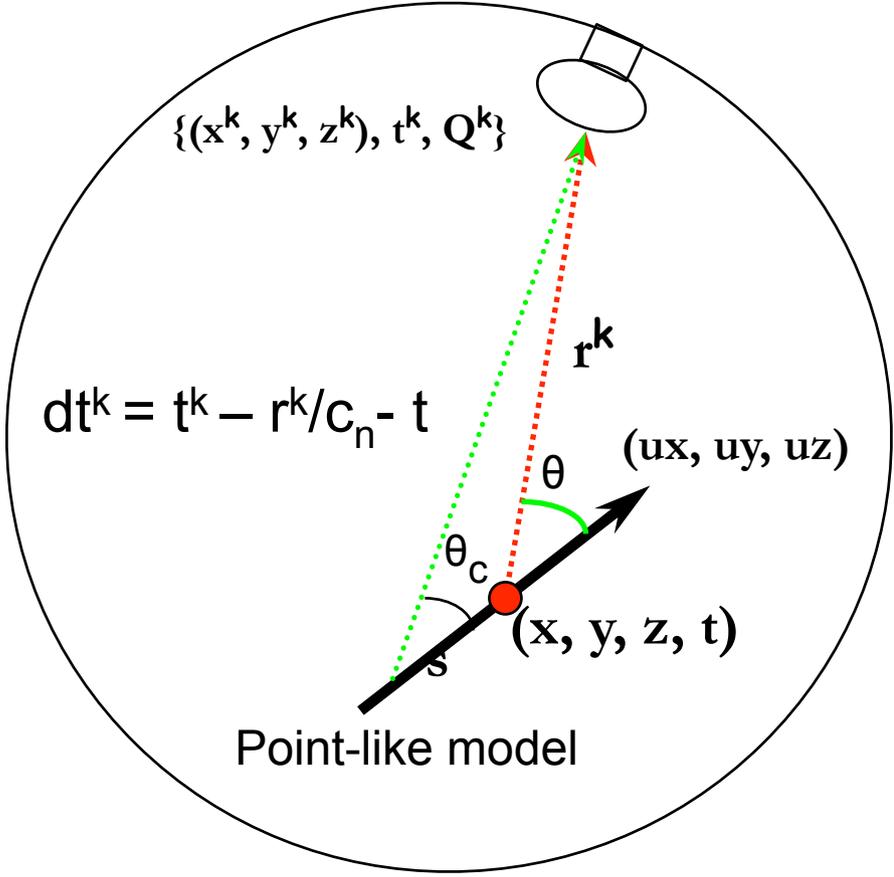
Re-training with newly weighted events improves performance.

# Boosted Decision Trees: Reconstruction and Particle ID

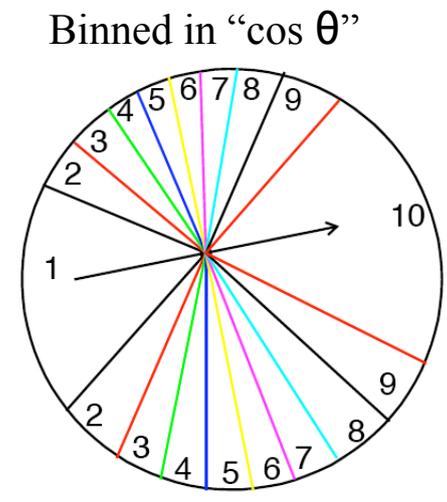
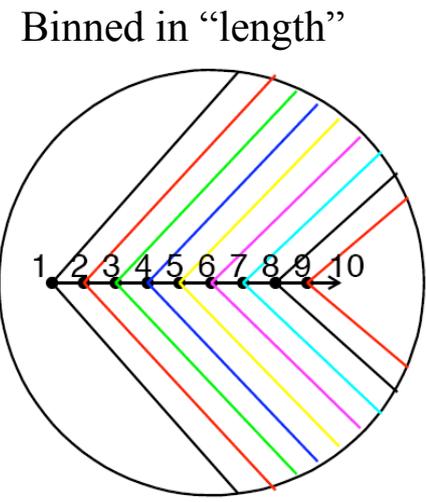
*Reconstruction fits a point-like light source:*

vertex, direction  $(\theta, \phi)$ , time, energy

Fitter resolution	
Vertex:	24 cm
Direction:	3.8°
Energy:	14%



Characterize topology of each event by dividing detector into “bins” relative to track:



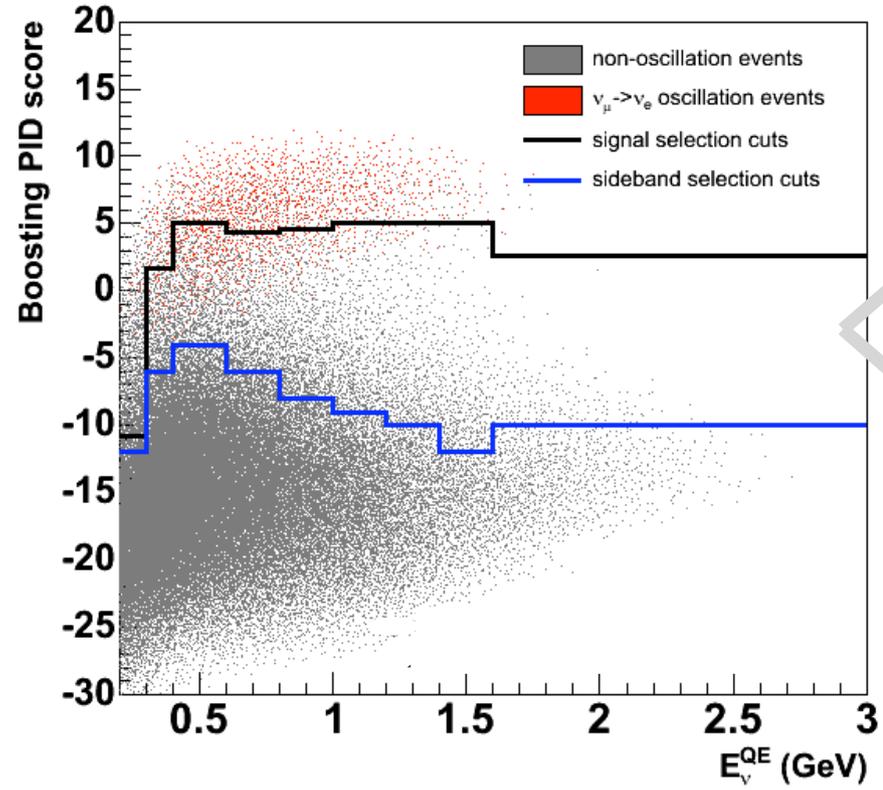
Particle ID “input variables” for the boosted decision trees are created from basic quantities in each bin: *e.g.*, charge, number of hits...

*To select events, a particle ID cut is made on the Boosting output score.*

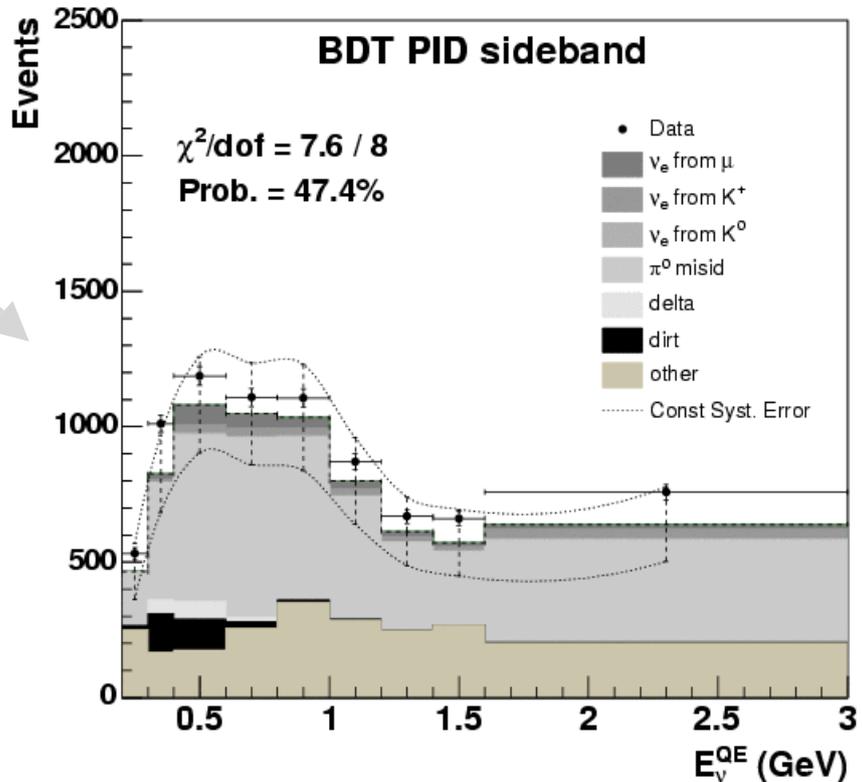
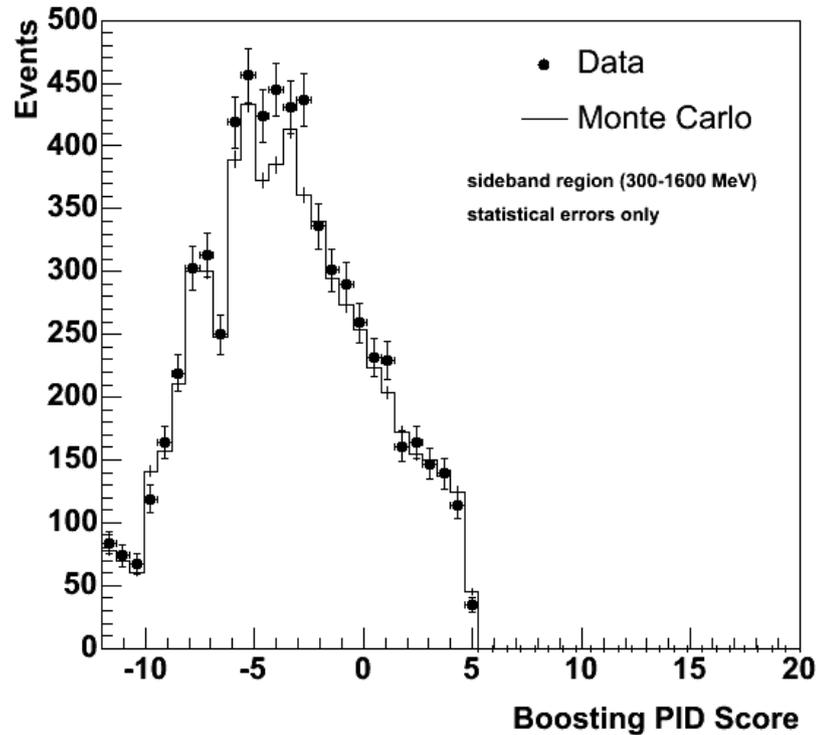
# Boosted Decision Trees: Particle ID

*A sideband region is selected to validate MC in region near signal.*

Sideband contains mostly mis-identified  $\pi^0$  background events.



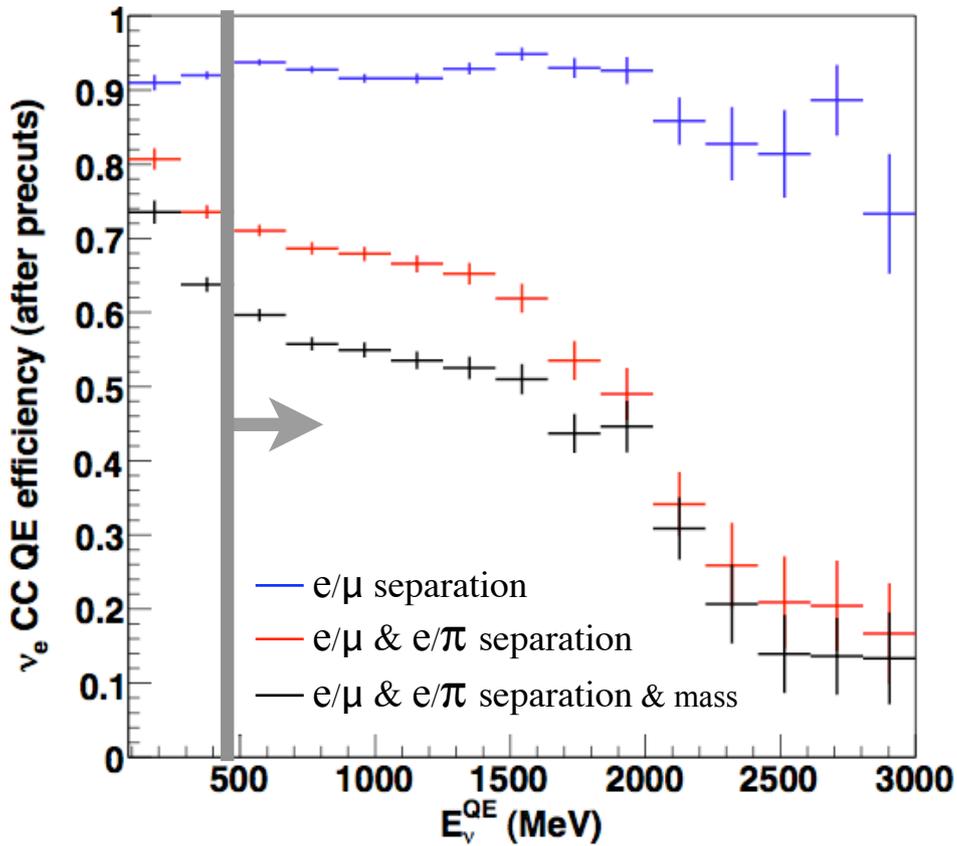
*A  $\chi^2$  is calculated using the full systematic error matrix, data and MC are consistent.*



# Comparison: Efficiencies

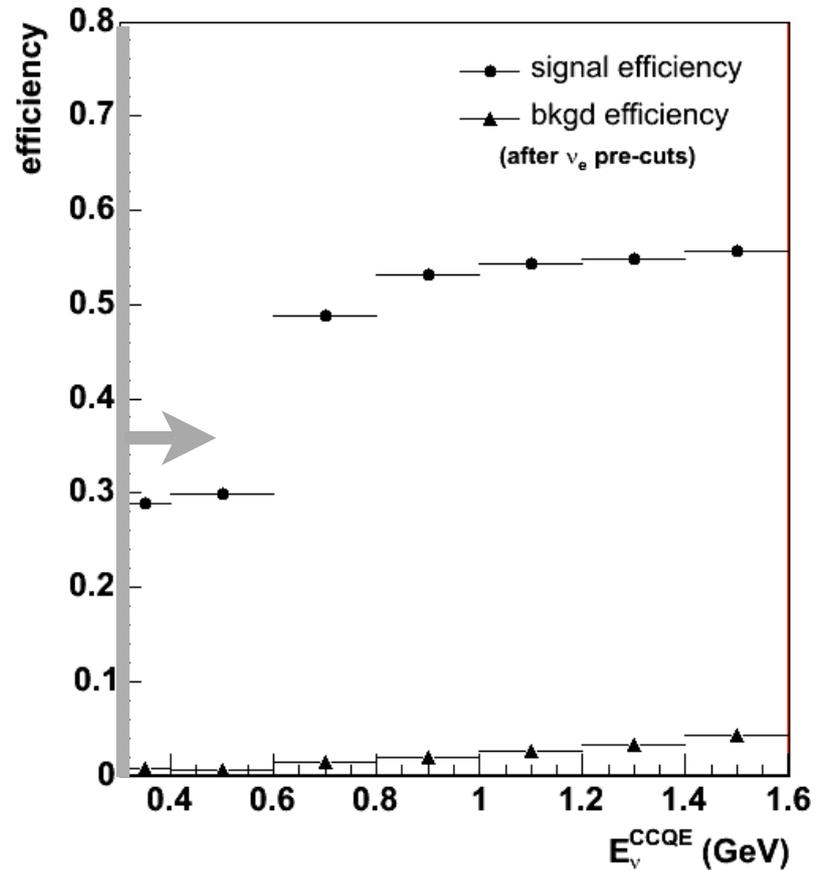
*The two analyses have different event selection efficiency vs. energy trends,*

Track-Based Analysis



$E_{\nu}^{QE} > 475$  MeV

Boosting Analysis



$E_{\nu}^{QE} > 300$  MeV

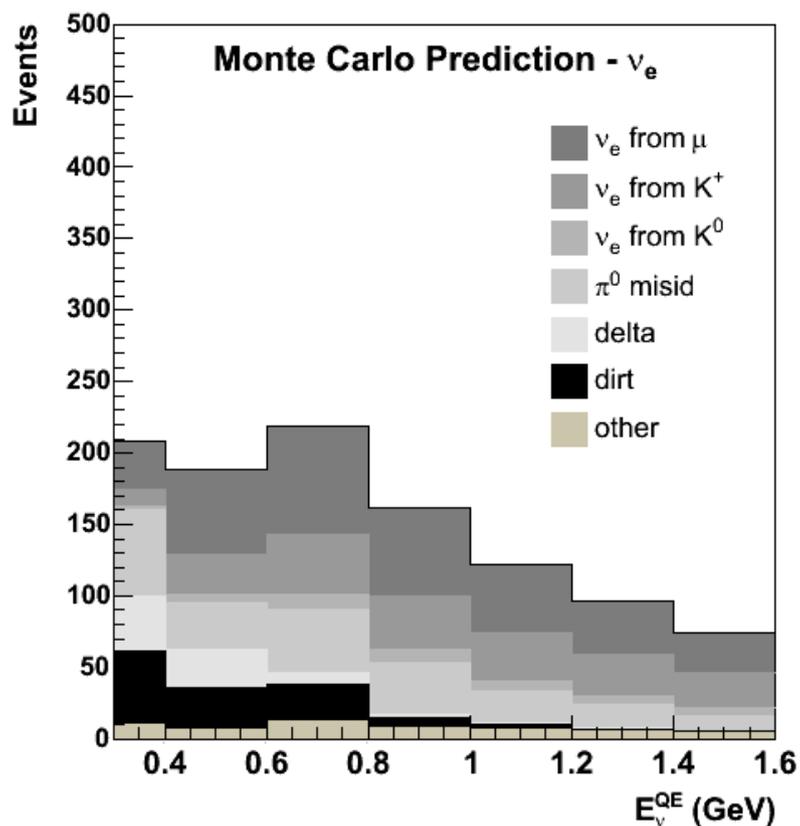
*and different reconstructed  $E_{\nu}$  regions for the oscillation analyses.*

# Comparison: Backgrounds

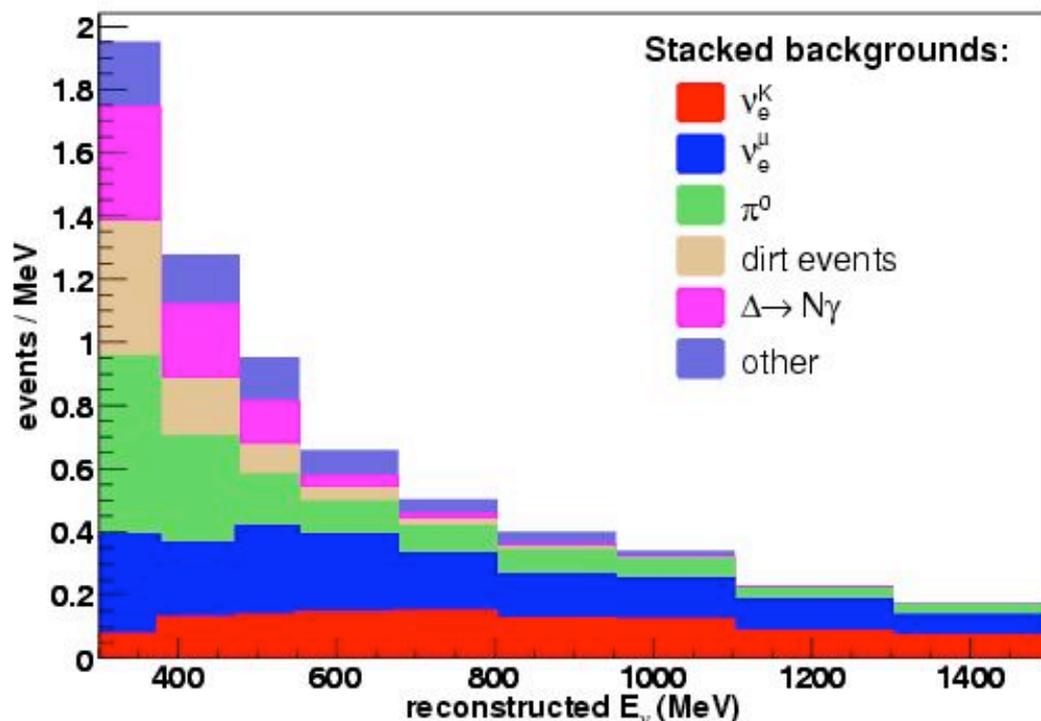
*The two analyses have somewhat different background compositions.*

Source	T-B	B
$\nu_e$ from $\mu$ decay	0.37	0.32
$\nu_e$ from $K$ decay	0.26	0.24
$\pi^0$ mis-ID	0.17	0.21
$\Delta \rightarrow N\gamma$	0.06	0.07
Dirt	0.05	0.11
Other	0.09	0.05

## Boosting Analysis



## Track-Based Analysis



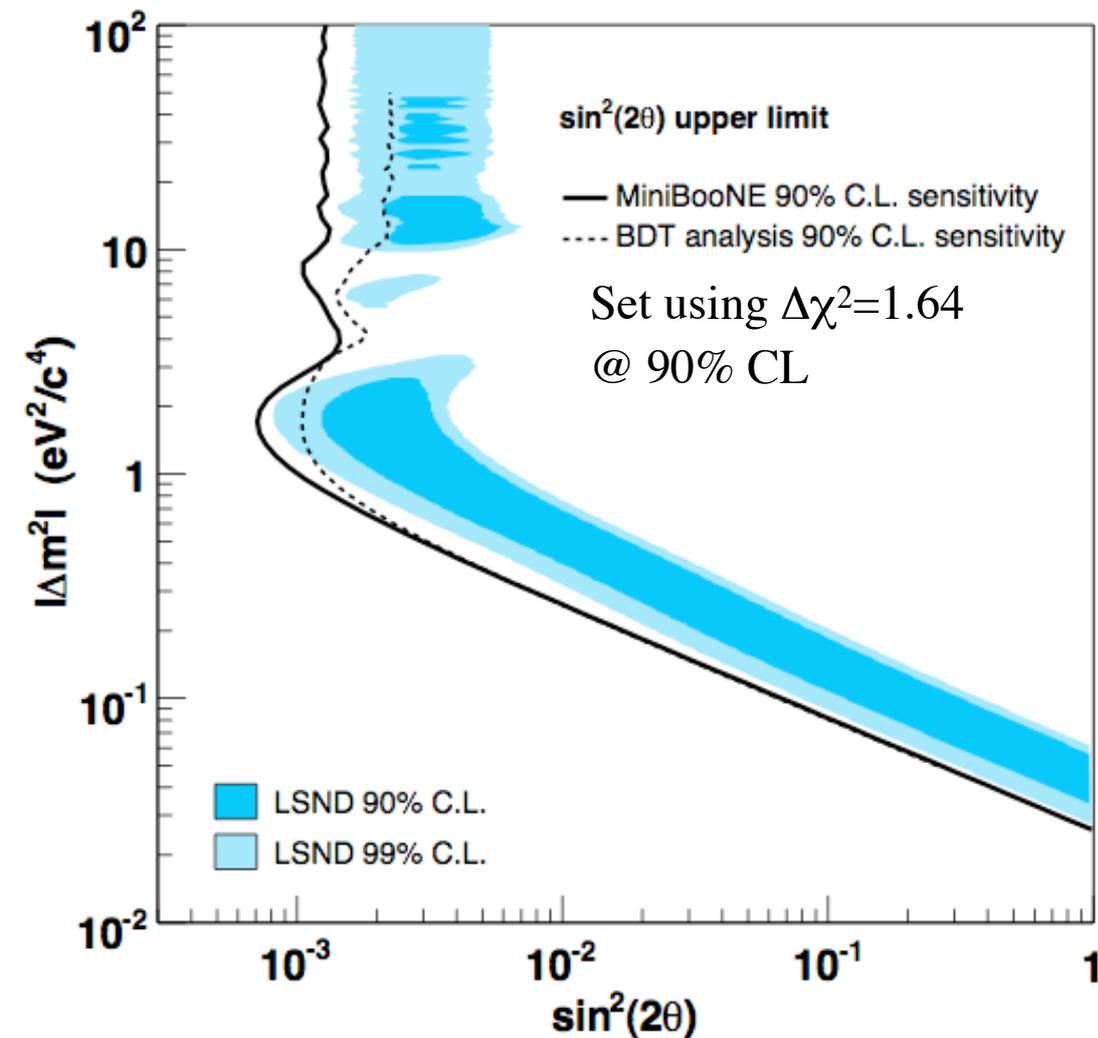
# Comparison: Systematic Errors

Both analyses construct error matrices for the oscillation fit, binned in  $E_\nu$ , to estimate the uncertainty on the expected number of  $\nu_e$  background events.

	<i>source</i>	<i>track-based (%)</i>	<i>boosting (%)</i>
✓	Flux from $\pi^+/\mu^+$ decay	6.2	4.3
✓	Flux from $K^+$ decay	3.3	1.0
✓	Flux from $K^0$ decay	1.5	0.4
	Target and beam models	2.8	1.3
✓	$\nu$ -cross section	12.3	10.5
	NC $\pi^0$ yield	1.8	1.5
	External interactions	0.8	3.4
✓	Optical model	6.1	10.5
	DAQ electronics model	7.5	10.8
	<i>constrained total</i>	9.6	14.5

Note:  
 “total” is **not**  
 the quadrature  
 sum-- errors are  
 further reduced  
 by fitting with  
 $\nu_\mu$  data ✓

# Comparison: Sensitivity



*Fit the Monte Carlo  $E_\nu^{QE}$  event distributions for oscillations*

Raster scan in  $\Delta m^2$ , and  $\sin^2 2\theta_{\mu e}$   
 (assume  $\sin^2 2\theta_{\mu x} = 0$ ),  
 calculate  $\chi^2$  value over  $E_\nu$  bins

$$\chi^2 = \sum_{i=1}^{N_{bins}} \sum_{j=1}^{N_{bins}} (m_i - t_i) \mathcal{M}_{ij}^{-1} (m_j - t_j)$$

$m_i$  = Number of measured data events in bin  $i$

$t_i$  = Number of predicted events in bin  $i$

( $t_i$  events are a function of  $\Delta m^2$ ,  $\sin^2 2\theta$ ,

$\mathcal{M}_{ij}^{-1}$  = Inverse of the covariance matrix

*Since the track-based analysis achieved better sensitivity than the boosted decision tree analysis, we decided (before opening the box) that it would be used for the primary result.*

1. Motivation & Introduction
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# Results: Opening the Box

After applying all analysis cuts:

*Step 1: Fit sequestered data to an oscillation hypothesis*

Fit does not return fit parameters

Unreported fit parameters applied to MC; diagnostic variables compared to data

Return only the  $\chi^2$  of the data/MC comparisons (for diagnostic variables only)

*Step 2: Open plots from Step 1 (Monte Carlo has unreported signal)*

Plots chosen to be useful diagnostics, without indicating if signal was added  
(reconstructed position, direction, visible energy...)

*Step 3: Report only the  $\chi^2$  for the fit to  $E_{\nu}^{QE}$*

No fit parameters returned

*Step 4: Compare  $E_{\nu}^{QE}$  for data and Monte Carlo,*

Fit parameters **are** returned

**This step breaks blindness**

*Step 5: Present results within two weeks*

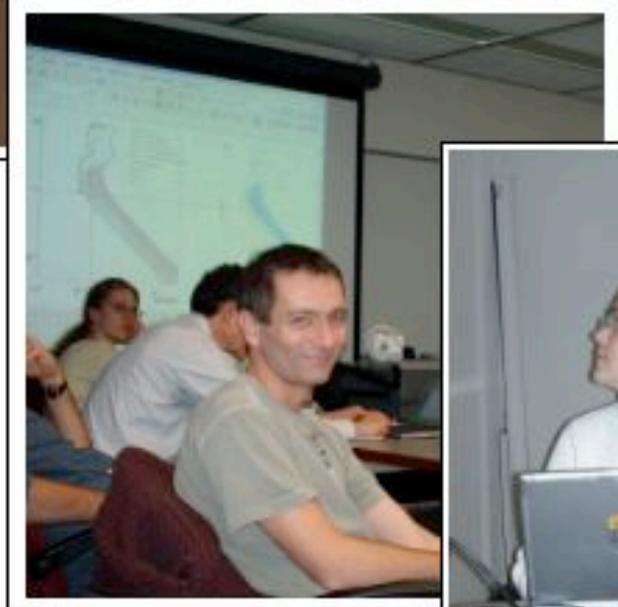
# Training for a blind search



MOW  
(blinded)  
c.2002

We opened the box on March 26, 2007

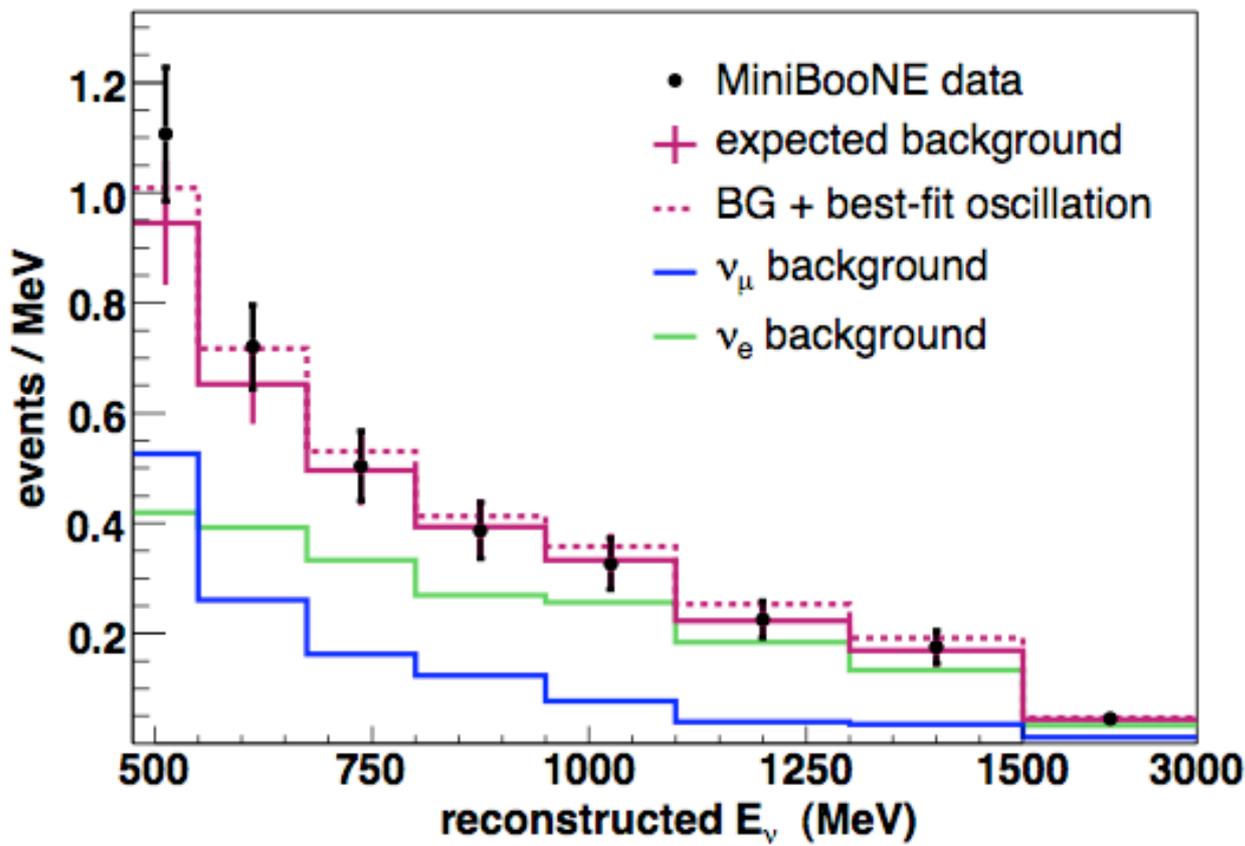
# The Box Opening



# Results: Track Based Analysis

Counting Experiment:  $475 < E_\nu^{QE} < 1250$  MeV  
 data: **380**  
 expectation:  **$358 \pm 19$  (stat)  $\pm 35$  (sys)**

**significance:**  
 **$0.55 \sigma$**



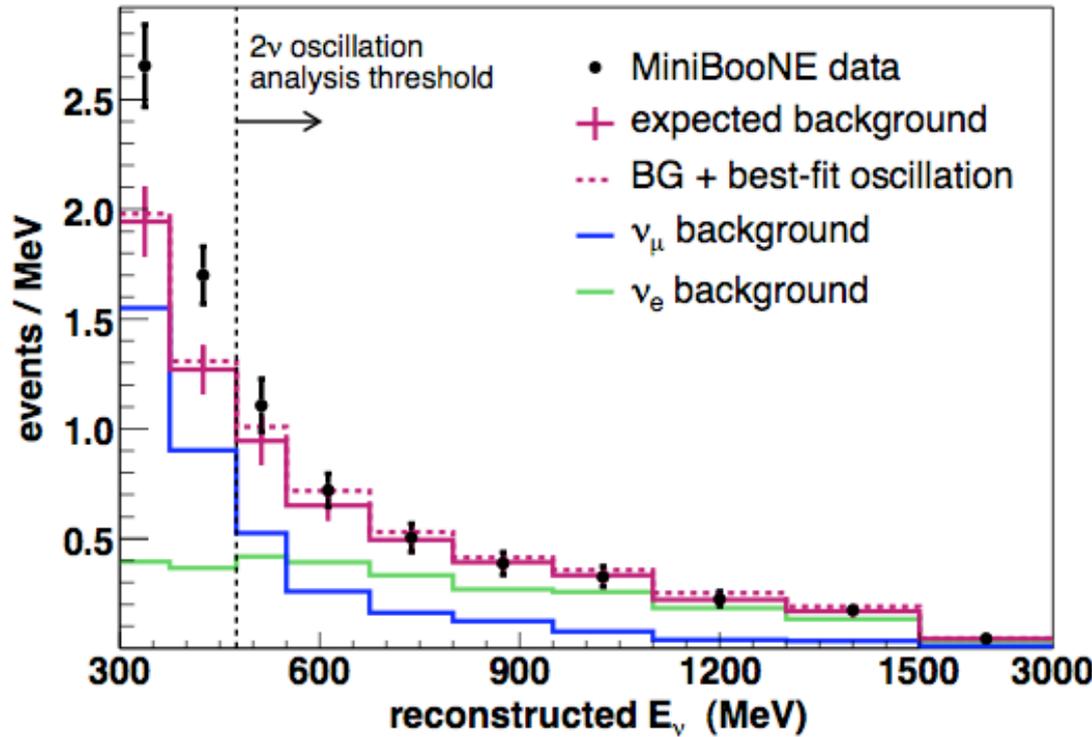
Best Fit (dashed):  
 $(\sin^2 2\theta, \Delta m^2) = (0.001, 4 \text{ eV}^2)$

$\chi^2$  prob. of best-fit point: 99%  
 $\chi^2$  prob. of null hypothesis: 93%

*We observe no significant evidence for an excess of  $\nu_e$  events in the energy range of the analysis.*

NB: Errors bars = diagonals of error matrix

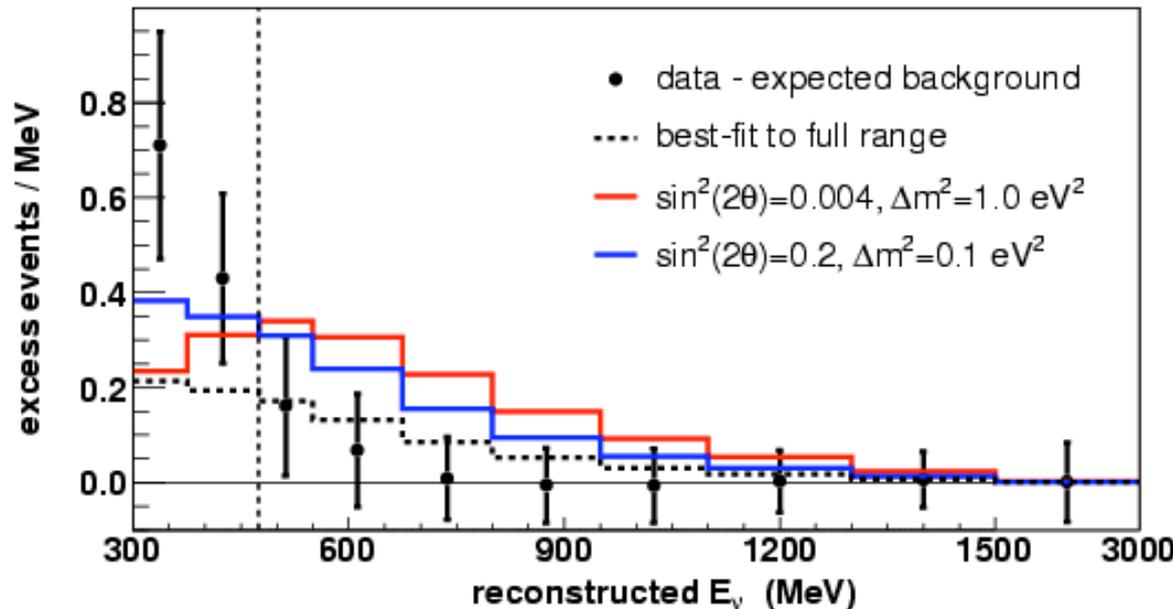
# Results: Track Based Analysis, Lower Energy Threshold



Extending down to energies below the analysis range:  $E_\nu^{QE} > 300$  MeV (we agreed to report this before box opening)

Data deviation for  $300 < E_\nu^{QE} < 475$  MeV:  $3.7\sigma$

Oscillation fit to  $E_\nu^{QE} > 300$  MeV:  
 Best Fit  $(\sin^2 2\theta, \Delta m^2) = (1.0, 0.03 \text{ eV}^2)$   
*Ruled out by Bugey*

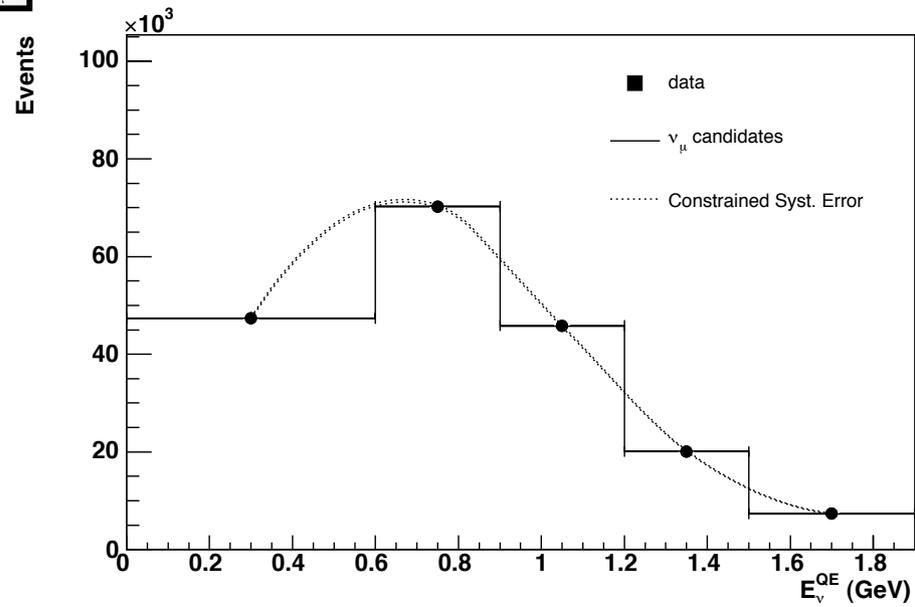


$\chi^2$  prob. at best-fit point: 18%  
 No closed contour for 90%CL

*Fit is inconsistent with  $\nu_\mu \rightarrow \nu_e$  oscillations.*

# Results: Boosted Decision Tree Analysis

$\nu_\mu$



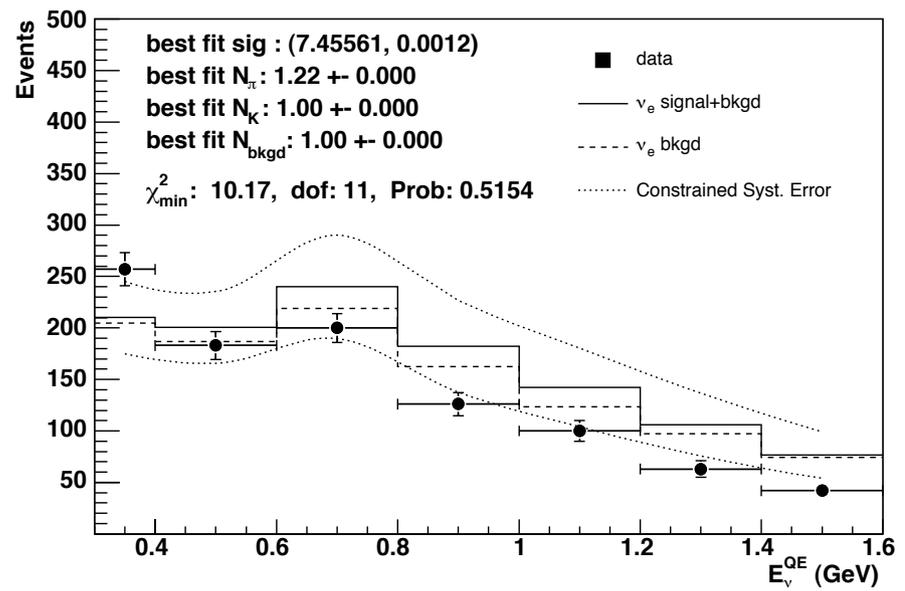
Counting Experiment:  
 $300 < E_\nu^{QE} < 1500 \text{ MeV}$

**significance:**  
 **$-0.38 \sigma$**

data: **971**  
 expectation:  **$1070 \pm 33 \text{ (stat)} \pm 225 \text{ (sys)}$**

Best Fit Point (dashed):  
 $(\sin^2 2\theta, \Delta m^2) = (0.001, 7 \text{ eV}^2)$

$\nu_e$

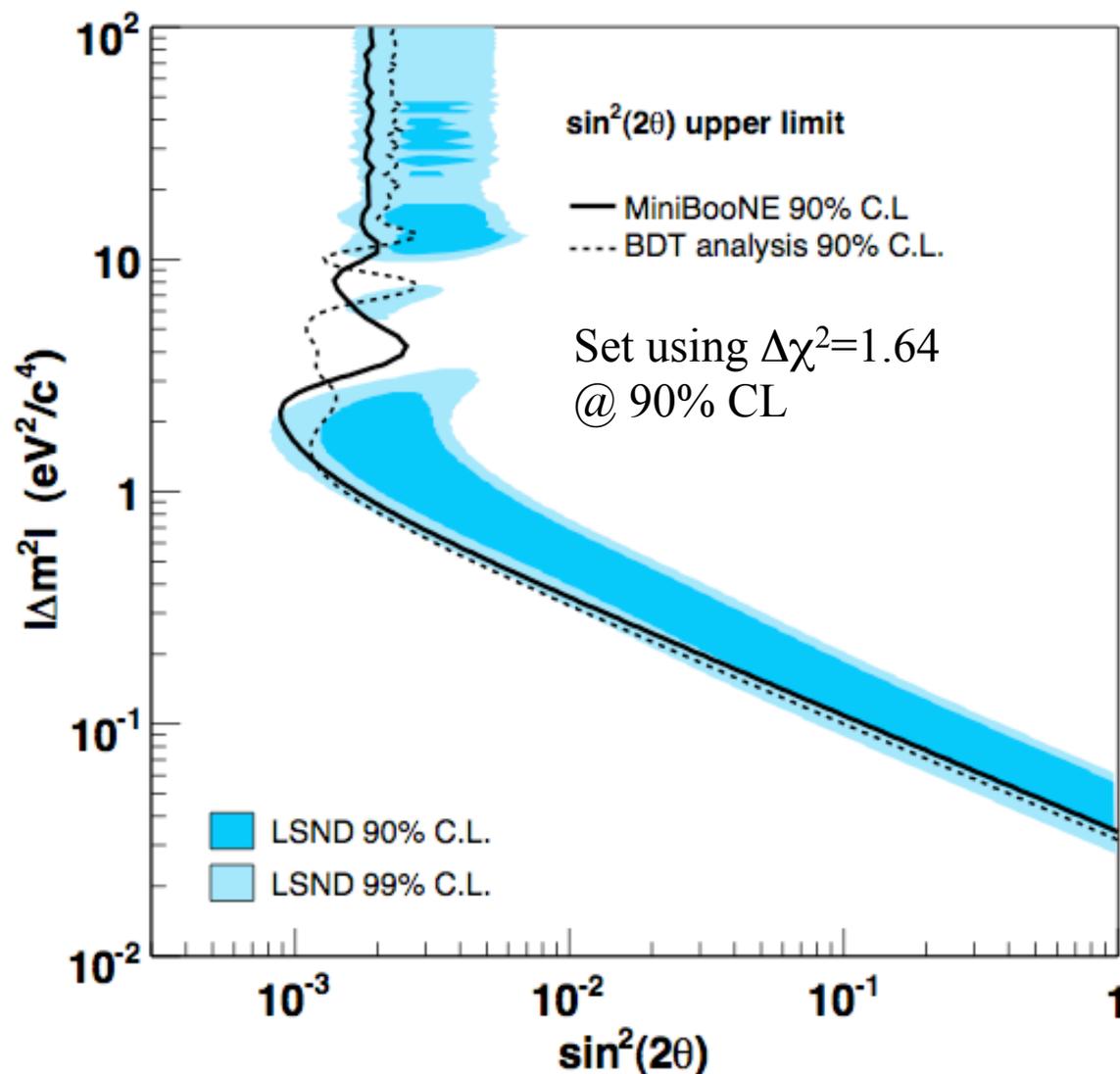


$\chi^2$  probability of best-fit point: 52%  
 $\chi^2$  probability of null hypothesis: 62%

*We observe no significant evidence for an excess of  $\nu_e$  events in the energy range of the analysis.*

# Results: Comparison

*MiniBooNE observes no evidence for  $\nu_\mu \rightarrow \nu_e$  appearance-only oscillations.*



The two independent oscillation analyses are in agreement.

solid: track-based  

$$\Delta\chi^2 = \chi^2_{best\ fit} - \chi^2_{null} = 0.94$$

dashed: boosting  

$$\Delta\chi^2 = \chi^2_{best\ fit} - \chi^2_{null} = 0.71$$

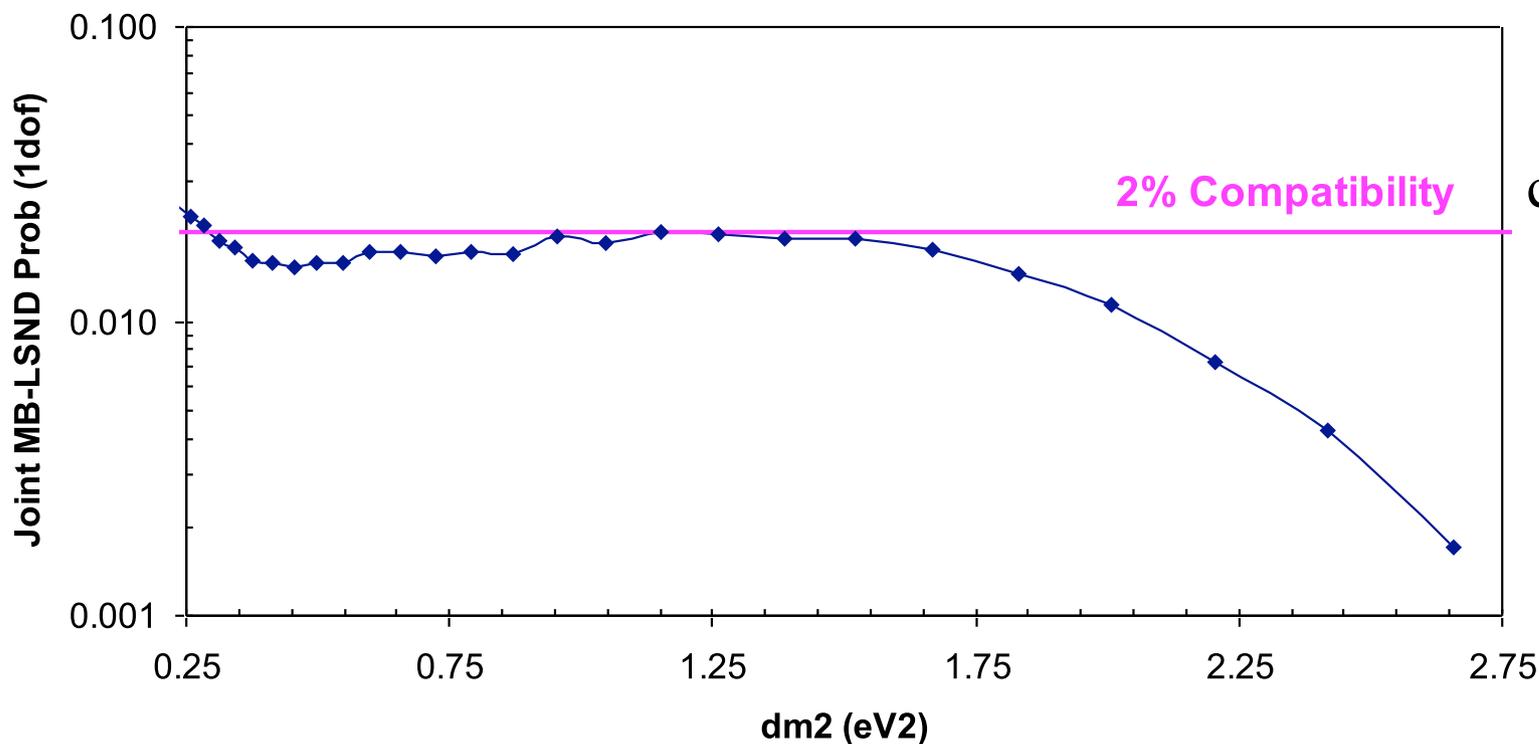
*Therefore, we set a limit.*

# Results: Compatibility with LSND

*A MiniBooNE-LSND Compatibility Test:*

$$\chi_0^2 = \frac{(z_{MB} - z_0)^2}{\sigma_{MB}^2} + \frac{(z_{LSND} - z_0)^2}{\sigma_{LSND}^2}$$

- For each  $\Delta m^2$ , form  $\chi^2$  between MB and LSND measurement
- Find  $z_0$  ( $\sin^2 2\theta$ ) that minimises  $\chi^2$  (weighted average of 2 measurements), this gives  $\chi_{\min}^2$
- Find probability of  $\chi_{\min}^2$  for 1 dof = joint compatibility probability for this  $\Delta m^2$



cf. LSND-KARMEN:  
64% compatibility

*MiniBooNE is incompatible with a  $\nu_{\mu} \rightarrow \nu_e$   
appearance-only interpretation of LSND at **98% CL***

# Results: Plans

*A paper on this analysis is posted to the archive.*

Many more papers supporting this analysis will follow, in the very near future:

$\nu_\mu$  CCQE production  
 $\pi^0$  production

We are pursuing further analyses of the neutrino data, including:  
an analysis which combines TB and BDT,  
less simplistic models for the LSND effect.

MiniBooNE is presently taking data in antineutrino mode.

*SciBooNE will start taking data in June!*

Will improve constraints on  $\nu_e$  backgrounds  
(intrinsic  $\nu_e$ s, improved  $\pi^0$  kinematics)

Will provide important constraints on “wrong-sign” BGs for  
antineutrino oscillation analysis

# Conclusions

- 1. Within the energy range of the analysis, MiniBooNE observes no statistically significant excess of  $\nu_e$  events above background.**
- 2. In two independent oscillation analyses, the observed  $E_\nu$  distribution is inconsistent with a  $\nu_\mu \rightarrow \nu_e$  appearance-only model.**
- 3. Therefore, we set a limit on  $\nu_\mu \rightarrow \nu_e$  oscillations at  $\Delta m^2 \sim 1 \text{ eV}^2$ .  
The MiniBooNE - LSND joint probability is 2%.**





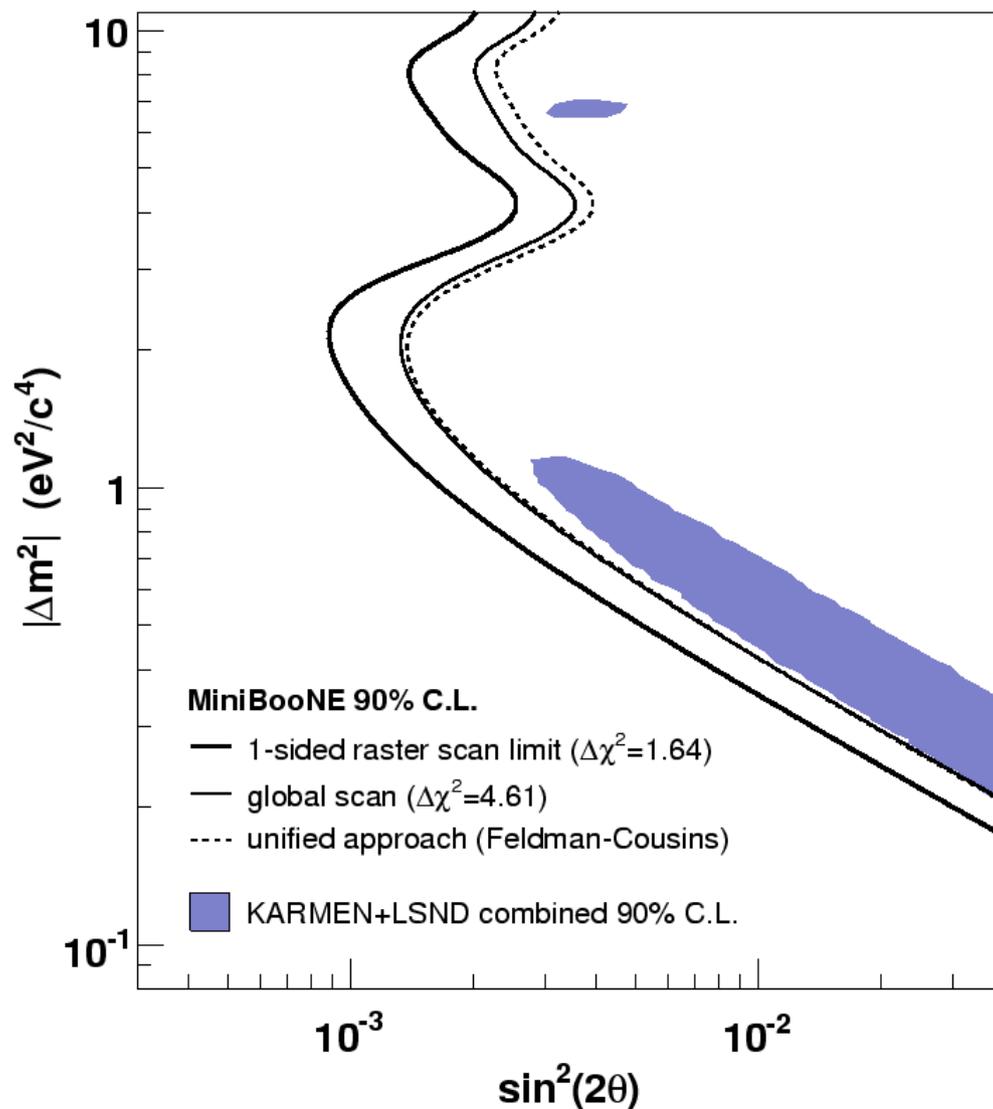
# Results: Interpreting Our Limit

*There are various ways to present limits:*

- Single sided raster scan  
(historically used, presented here)
- Global scan
- Unified approach  
(most recent method)

This result must be folded into an LSND-Karmen joint analysis.

*Church, et al., PRD 66, 013001*



*We will present a full joint analysis soon.*

# Results: Event Overlap

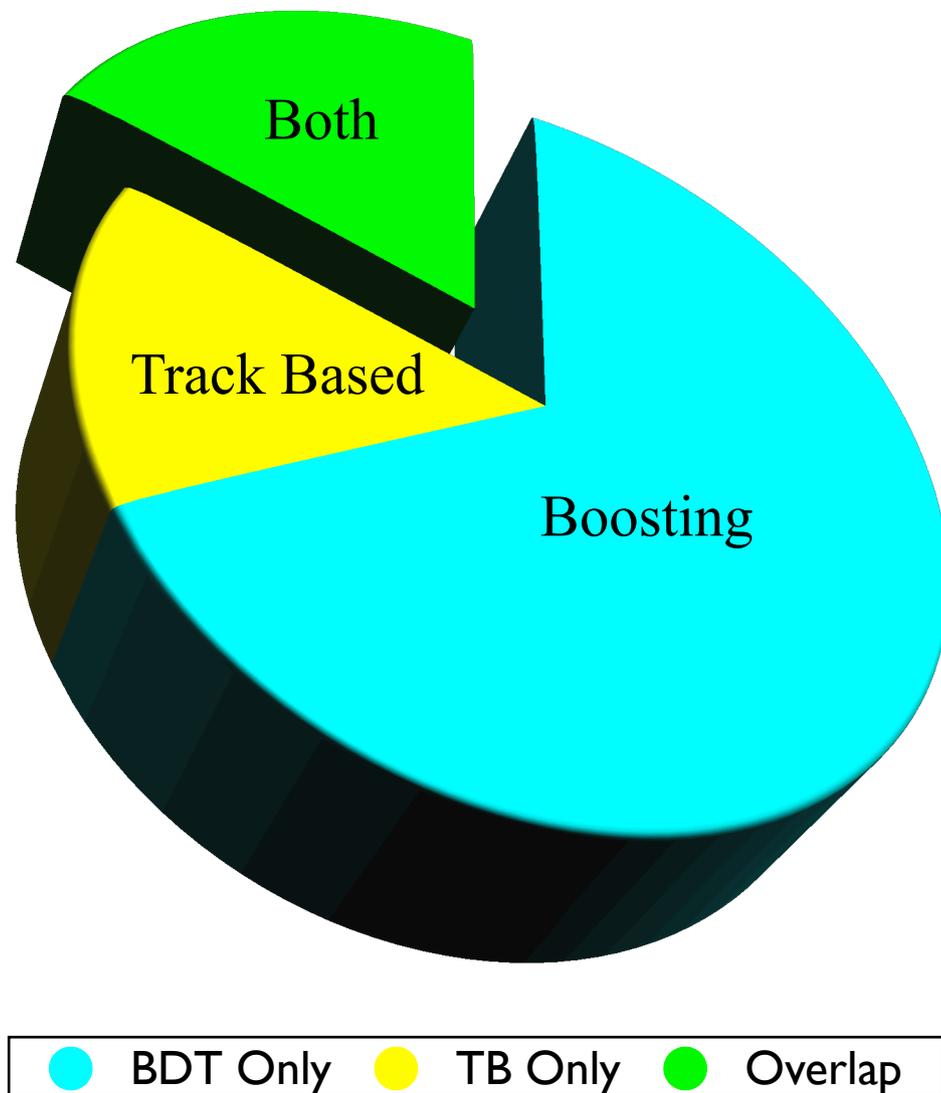
Counting experiment numbers:

Track Based Algorithm finds 380 events

Boosting Algorithm finds 971 events

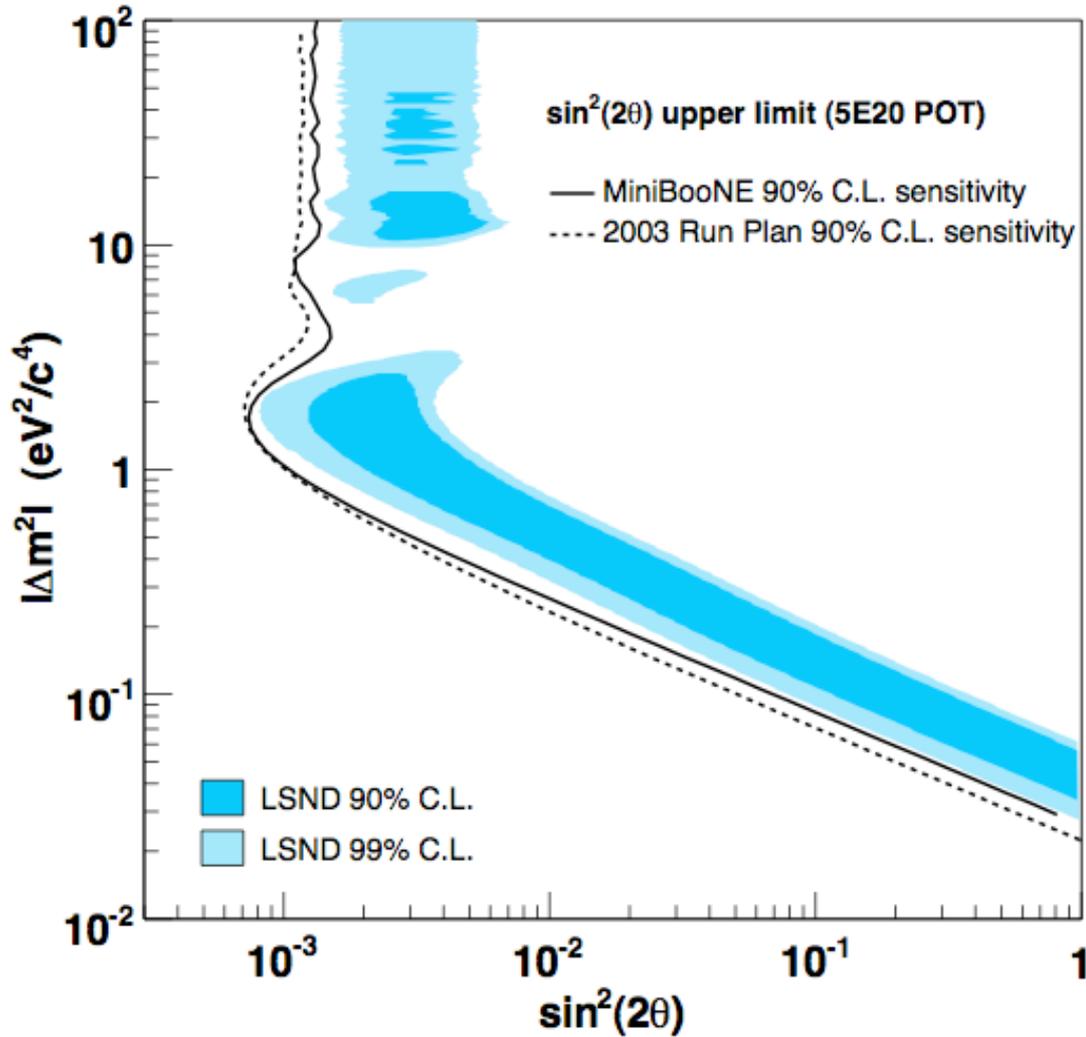
However, only 1131 events total,  
because 220 overlap

- chosen by both algorithms!



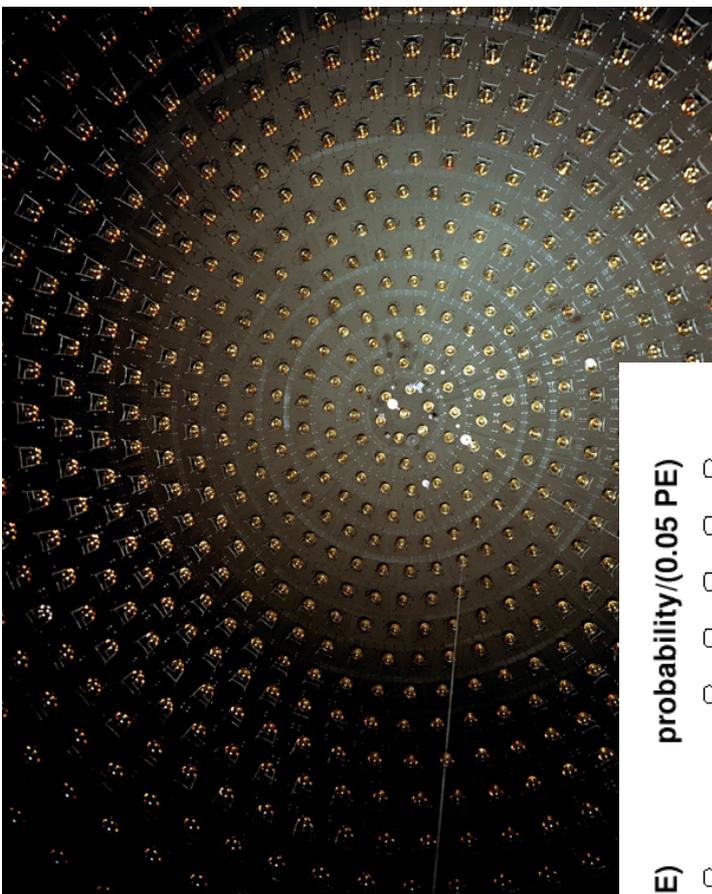
# Results: Sensitivity Goal

*Compared to our sensitivity goal for 5E20 protons on target from 2003 Run Plan*



Set using  $\Delta\chi^2=1.64$  @ 90% CL

# MiniBooNE Detector: PMT Calibration



10% photo-cathode coverage

*PMTs are calibrated with a laser + 4 flask system*

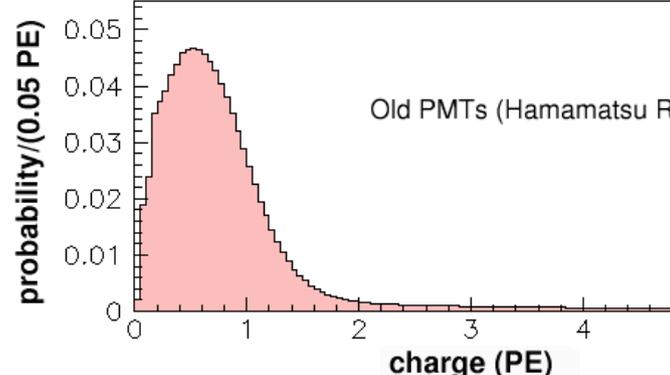
PMT Charge Resolution: 1.4 PE, 0.5 PE

PMT Time Resolution: 1.7 ns, 1.1 ns

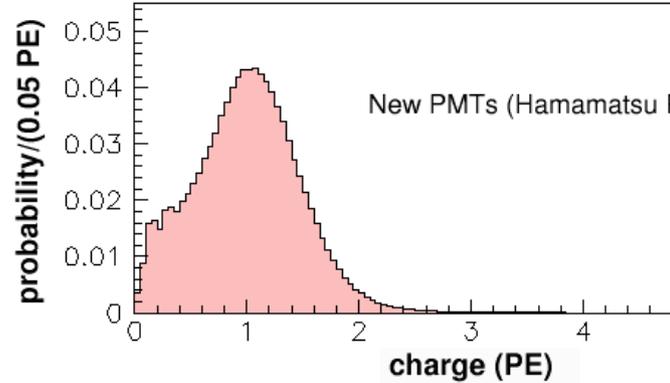
Two types of 8”  
Hamamatsu Tubes:  
R1408, R5912

*Laser data are acquired at 3.3 Hz to continuously  
calibrate PMT gain and timing constants*

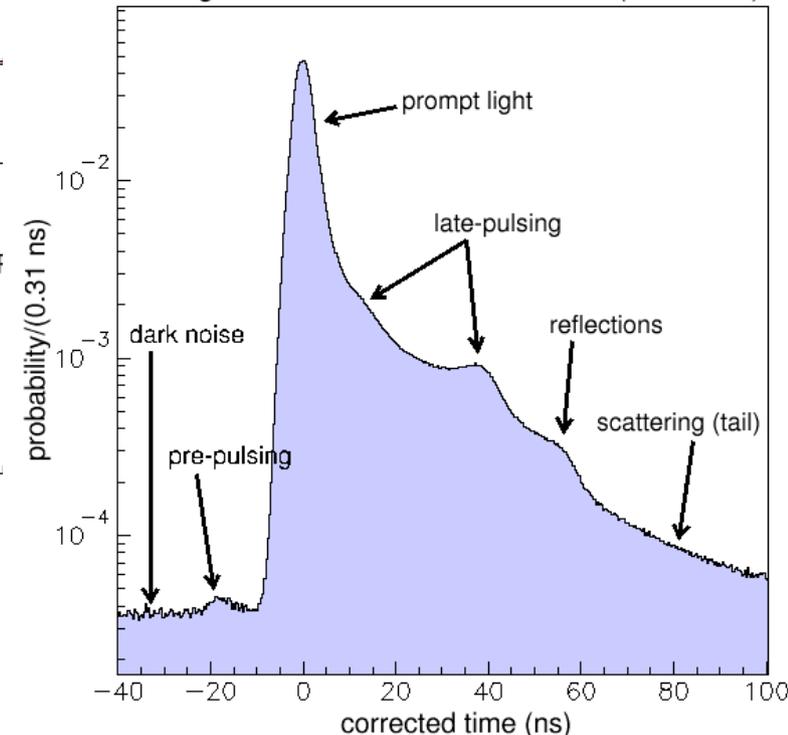
Single Photoelectron Response in MiniBooNE



R. B. Patterson

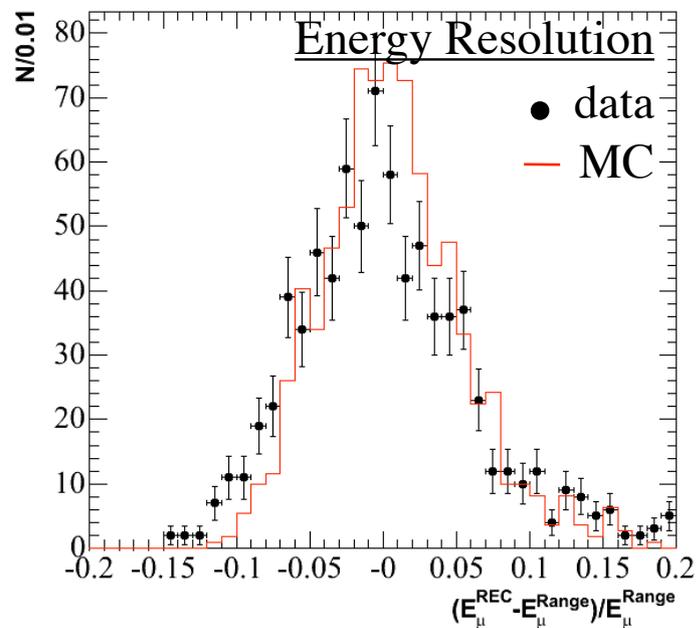
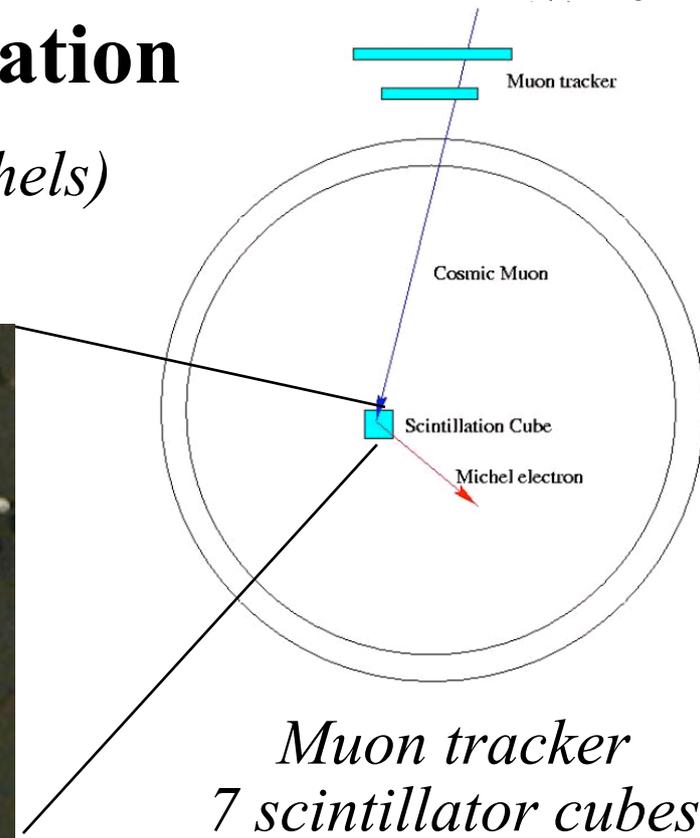
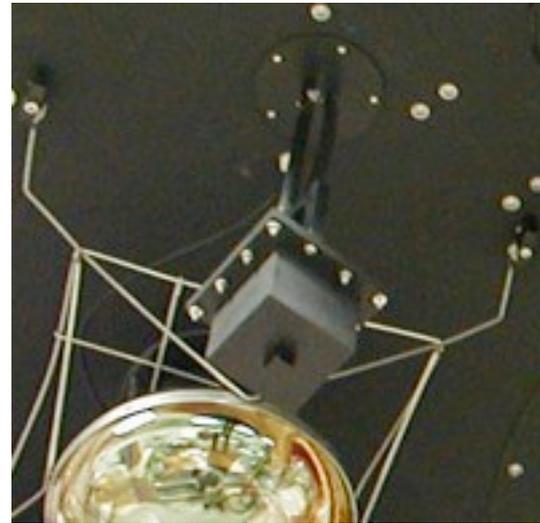
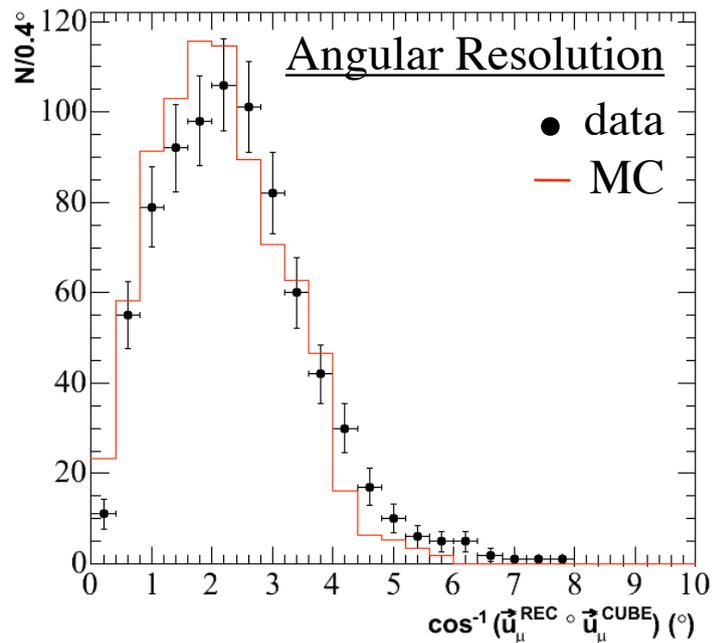


Timing Distribution for Laser Events (old tubes)



# MiniBooNE Detector: Cosmic Calibration

*use cosmic muons and their decay electrons (Michels)*



- Cosmic muons which stop in cubes:
- test energy scale extrapolation up to 800 MeV
  - measure energy, angle resolution
  - compare data and MC

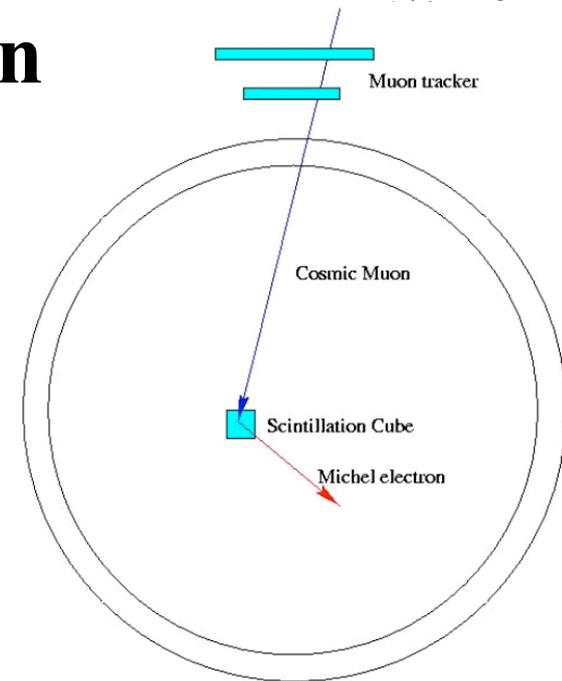
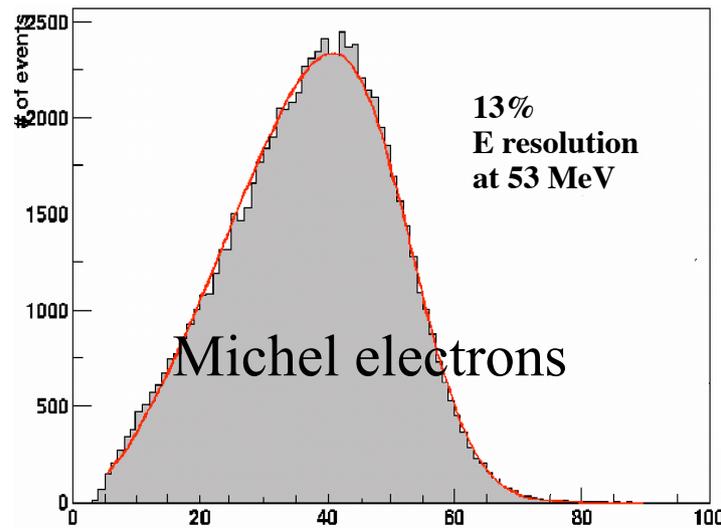
*Muon tracker + cube calibration  
data continuously acquired at 1 Hz*

# MiniBooNE Detector: Cosmic Calibration

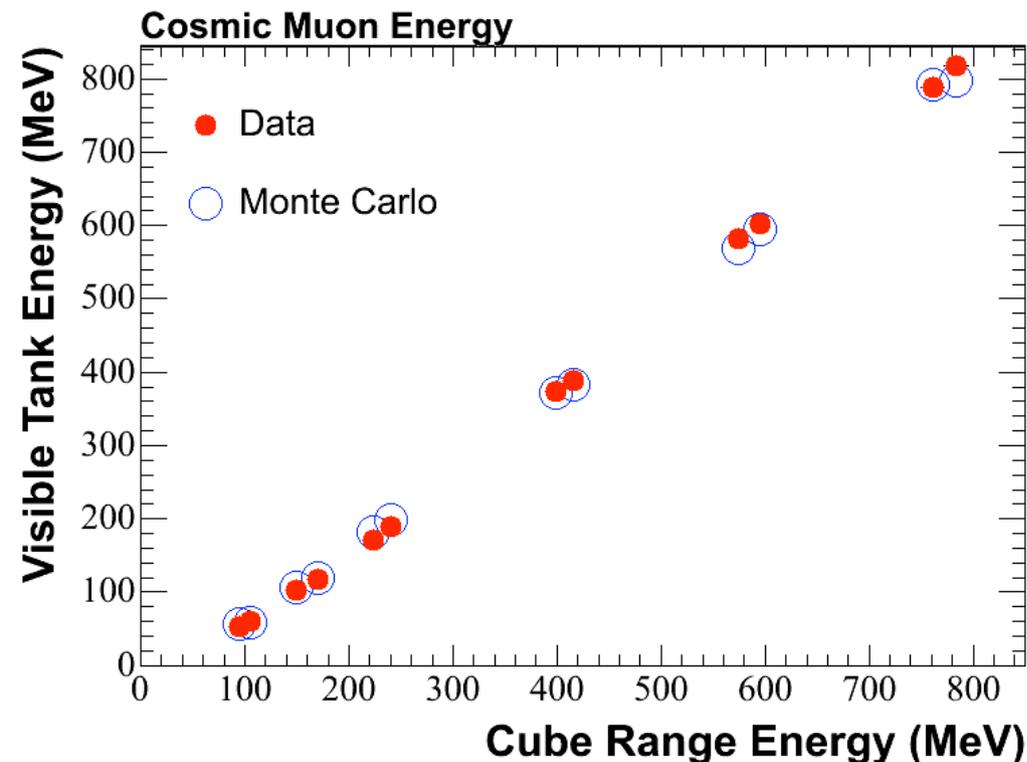
*use cosmic muons and their decay electrons (Michels)*

## Michel electrons:

- set absolute energy scale and resolution at 53 MeV endpoint
- optical model tuning



*Muon tracker  
7 scintillator cubes*

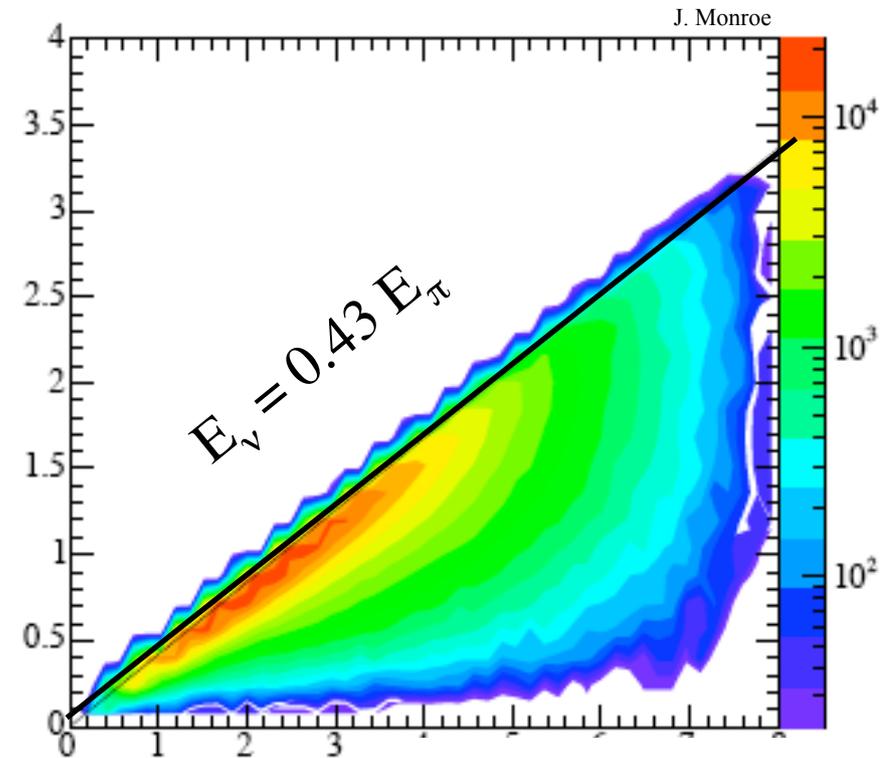
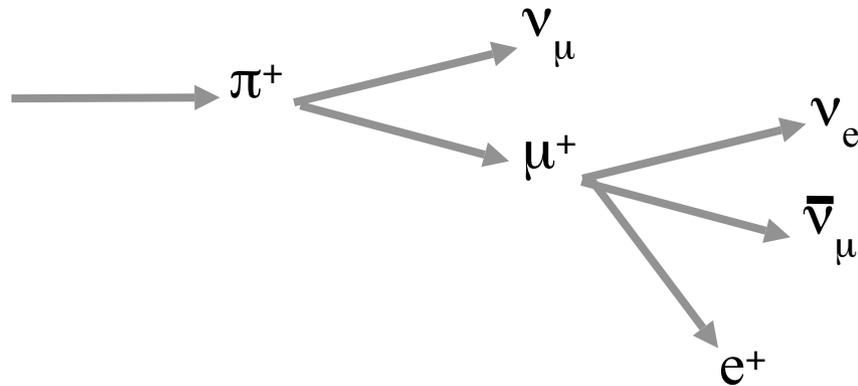


- Cosmic muons which stop in cubes:
- test energy scale extrapolation up to 800 MeV
  - measure energy, angle resolution
  - compare data and MC

*Muon tracker + cube calibration  
data continuously acquired at 1 Hz*

# Incorporating $\nu_\mu$ Data: $\mu^+$ -Decay $\nu_e$ Background

$\nu_\mu$  CCQE events measure the  $\pi^+$  spectrum, this constrains the  $\mu^+$ -decay  $\nu_e$  flux



*this works well because the  $\nu_\mu$  energy is highly correlated with the  $\pi^+$  energy*

## Ratio Method Constraint:

1. MC based on external data predicts a central value and a range of possible  $\nu_\mu(\pi)$  fluxes
2. make Data/MC ratio vs.  $E_{\nu}^{QE}$  for  $\nu_\mu$  CCQE data
3. re-weight each possible MC flux by the ratio (2) including the  $\nu_\mu$ , its parent  $\pi^+$ , sister  $\mu^+$ , and niece  $\nu_e$

# Analysis Strategy: Delta Background

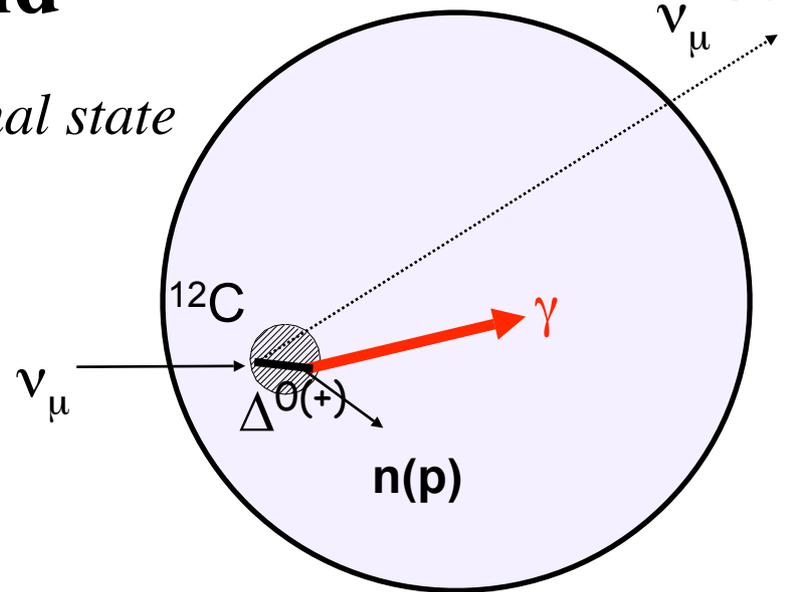
*$\nu$  induced interactions that produce single  $\nu$ s in the final state*

## Radiative Delta Decay (NC)

- (i) Use  $\pi^0$  events to measure rate of NC  $\Delta$  production
- (ii) Use PDG branching ratio for radiative decay  
- 15% uncertainty on branching ratio

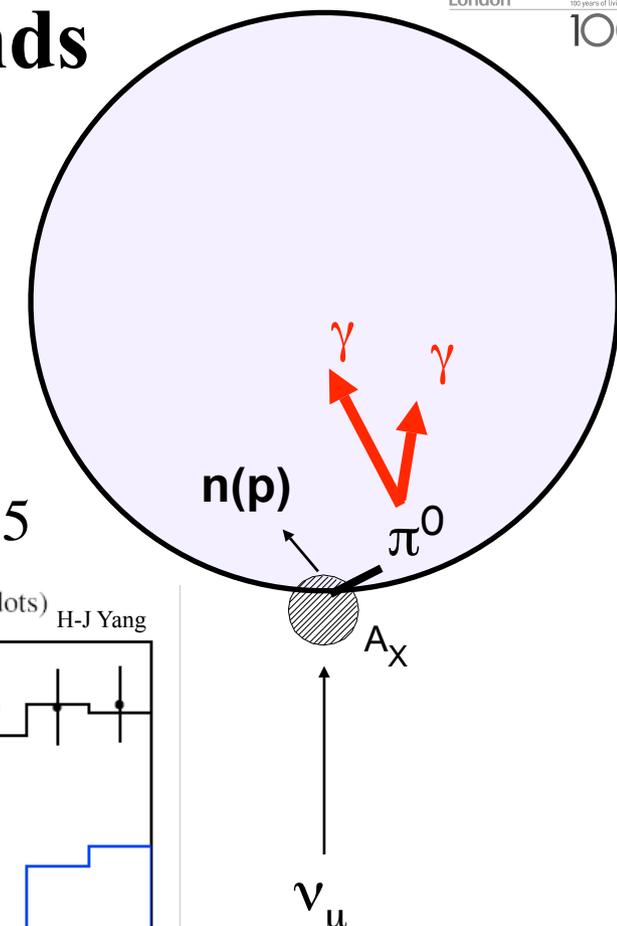
## Inner Bremsstrahlung (CC)

- (i) Hard photon released from neutrino interaction vertex
- (ii) Use events where the  $\mu$  is tagged by the decay  $e^-$   
- study misidentification using BDT algorithm.



# Analysis Strategy: External Backgrounds

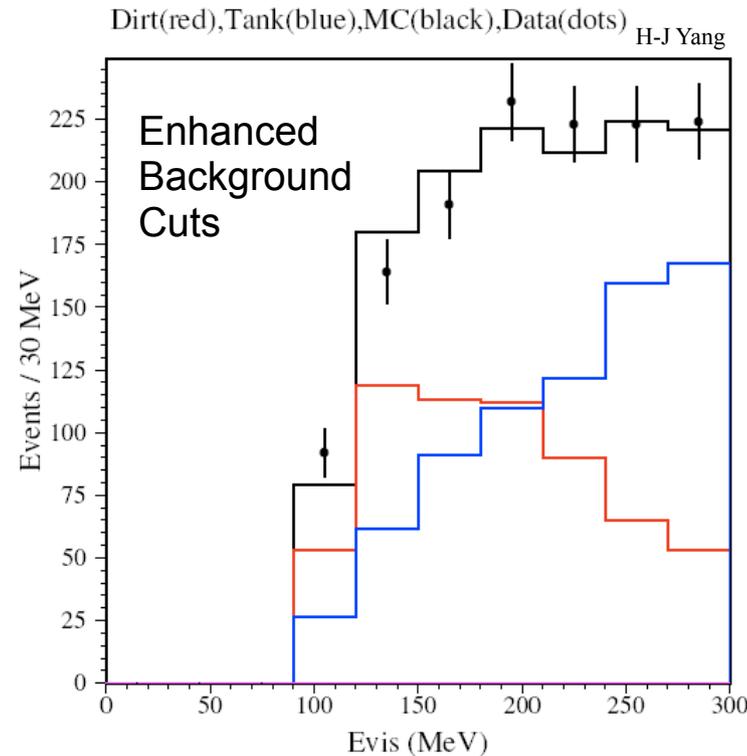
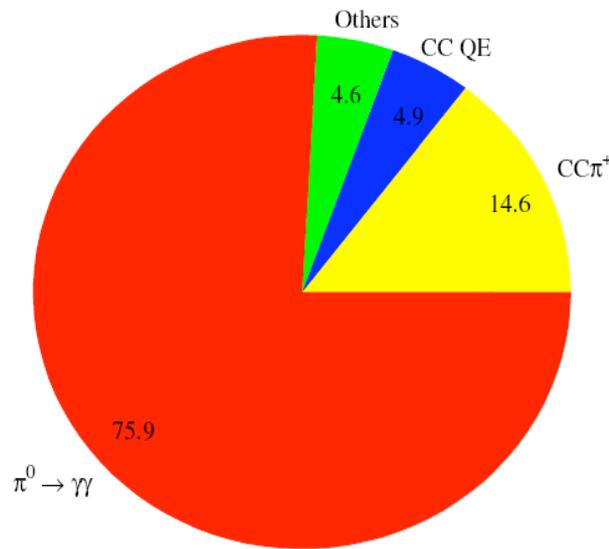
*interactions outside the detector that deposit energy in the fiducial volume and pass the veto PMT hits cut*



## 1. "Dirt" Events

$\nu$  interactions outside of the detector are measured in the "dirt box:"  $N_{\text{data}}/N_{\text{MC}} = 0.99 \pm 0.15$

Event Type of Dirt Events



## 2. Cosmic Ray Background Events

Measured from 126E6 strobe data triggers:  $2.1 \pm 0.5$  events.

# Analysis Overview: Background Summary

*Summary of predicted backgrounds for the primary MiniBooNE result  
 (Track-Based Analysis):*

Process	Number of Events
$\nu_\mu$ CCQE	10
$\nu_\mu e \rightarrow \nu_\mu e$	7
Miscellaneous $\nu_\mu$ Events	13
NC $\pi^0$	62
NC $\Delta \rightarrow N\gamma$	20
NC Coherent & Radiative $\gamma$	< 1
Dirt Events	17
$\nu_e$ from $\mu$ Decay	132
$\nu_e$ from $K^+$ Decay	71
$\nu_e$ from $K_L^0$ Decay	23
$\nu_e$ from $\pi$ Decay	3
Total Background	358
0.26% $\nu_\mu \rightarrow \nu_e$	(example signal) 163

# Step 1

Return the  $\chi^2$  of the data/MC comparison for  
a set of diagnostic variables

12 variables are tested for TB  
46 variables are tested for BDT

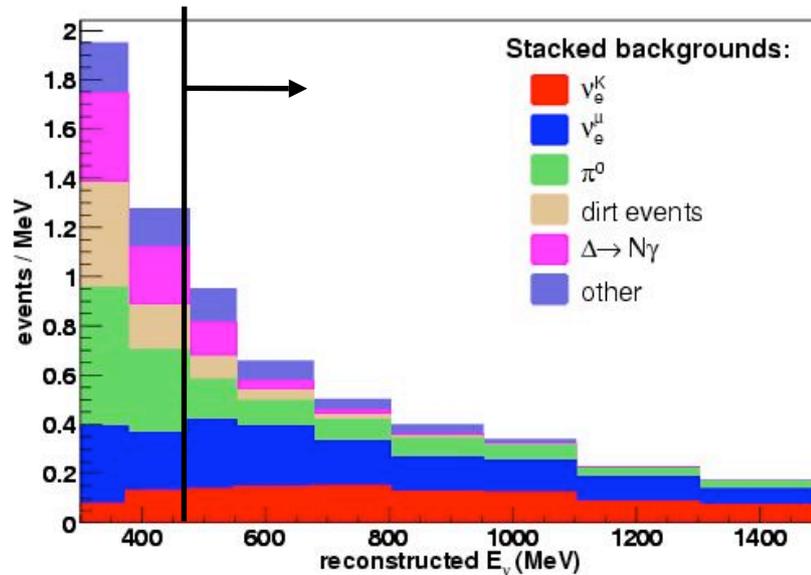
All analysis variables were returned with good  
probability except...

TB analysis  $\chi^2$  Probability of  $E_{\text{visible}}$  fit: 1%

**This probability was sufficiently low  
to merit further consideration**

## In the TB analysis

- We re-examined our background estimates using sideband studies.  
 ⇒ *We found no evidence of a problem*
- However, knowing that backgrounds rise at low energy,  
*We tightened the cuts for the oscillation fit:*



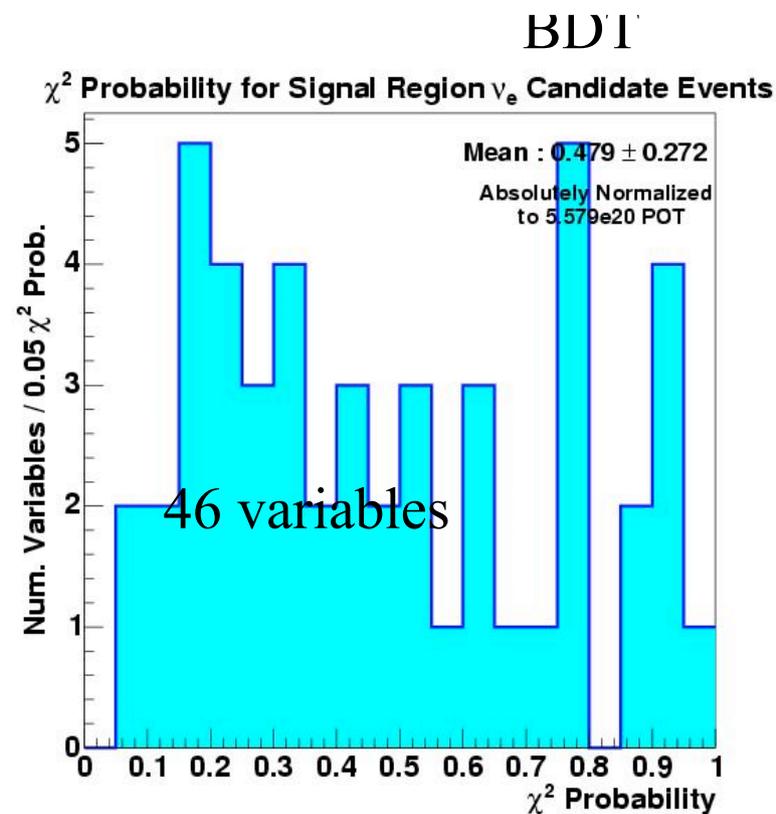
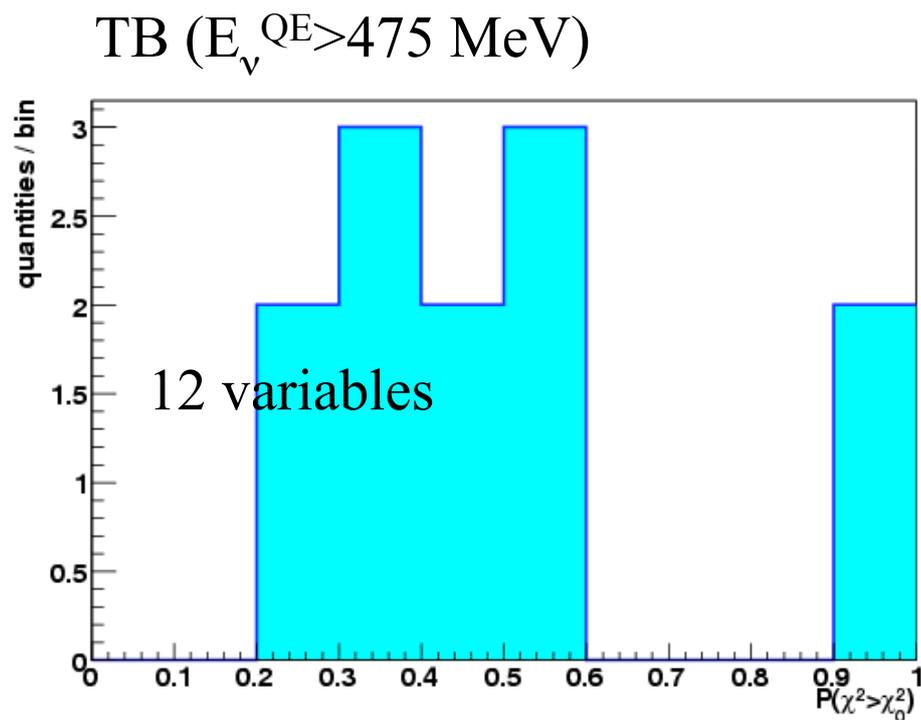
$$E_\nu^{QE} > 475 \text{ MeV}$$

We agreed to report events over the original full range:  
 $E_\nu^{QE} > 300 \text{ MeV}$ ,

## Step 1: again!

Return the  $\chi^2$  of the data/MC comparison for a set of diagnostic variables

$\chi^2$  probabilities returned:

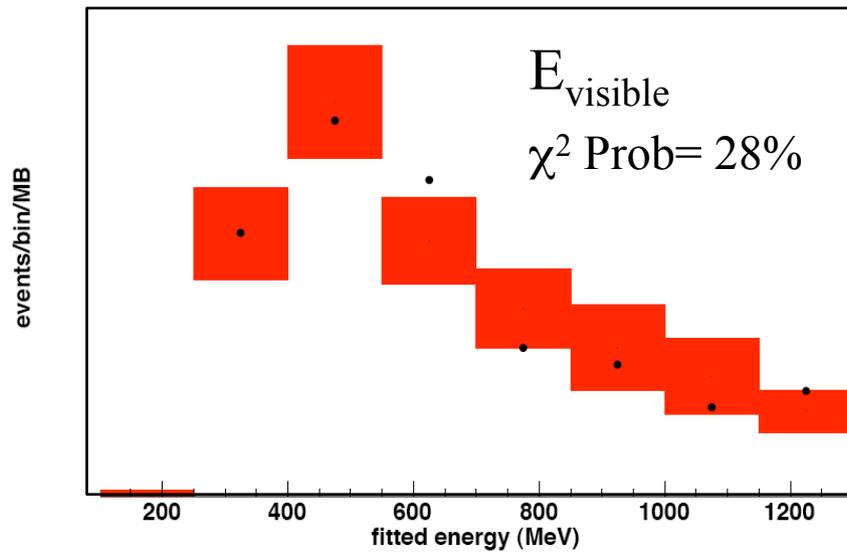


Parameters of the oscillation fit were not returned.

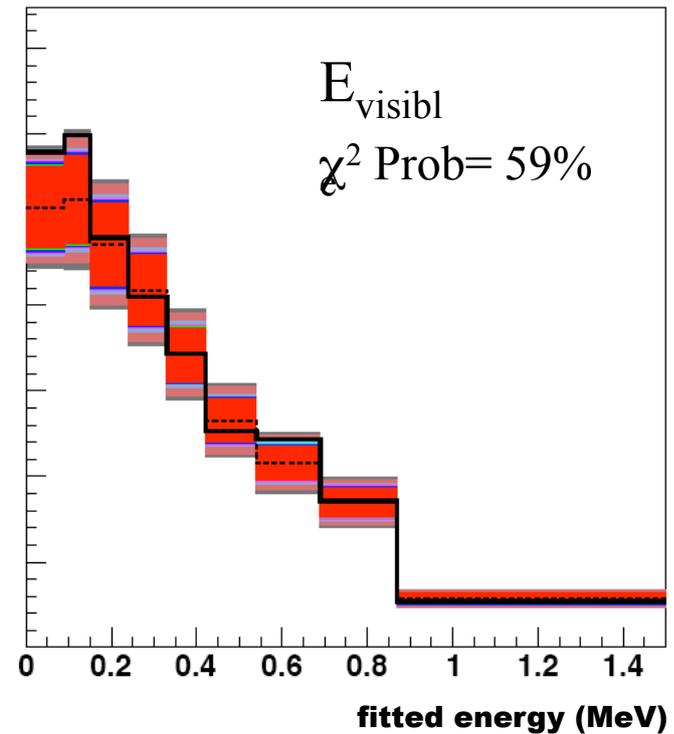
# Step 2

Open up the plots from step 1 for approval.

*Examples of  
what we saw:*



TB ( $E_{\nu}^{\text{QE}} > 475$  MeV)



BDT

MC contains fitted signal at unknown level

Step  
 3

Report the  $\chi^2$  for a fit to  $E_\nu^{\text{QE}}$  across full energy range

TB ( $E_\nu^{\text{QE}} > 475$  MeV)  $\chi^2$  Probability of fit: 99%  
 BDT analysis  $\chi^2$  Probability of fit: 52%

Leading to...

Step  
 4  
 Open the box...