# DM@LHC2016: Workshop Impressions



### Andreas Korn

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### Introduction



# DARK MATTER

Most of the universe can't even be bothered to interact with you.

M. Buckley

Andreas Korn



### Introduction

### What do we know?

- How much: Ω<sub>DM</sub> ≈ 0.26
- Likely particle with non-gravitational interactions
- Dark
  - Electrically neutral probably
  - Colour neutral (H-dibaryon...)
- Cold: nonrelativistic during structure formation
- Sufficiently long-lived
- Non-baryonic (from BBN  $\Omega_{\rm B} \approx 0.04$ )

#### Candidate within the Standard Model of particle physics?

- Neutrinos
  - Correspond to hot DM
  - Cannot account for the observed dark matter density



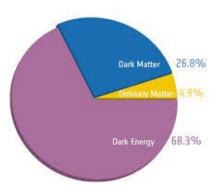
Physics beyond the Standard Model !

Kai Schmidt-Hoberg | Overview of Dark Matter models | 30 March 2016 | Page 7

Many candidates (theorists are inventive...)



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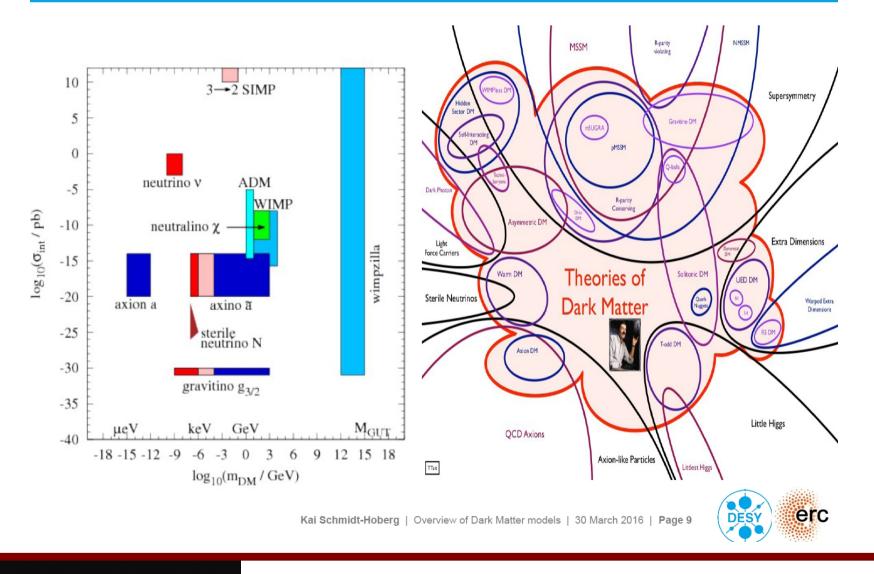




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### Introduction

### **Particle physics candidates**



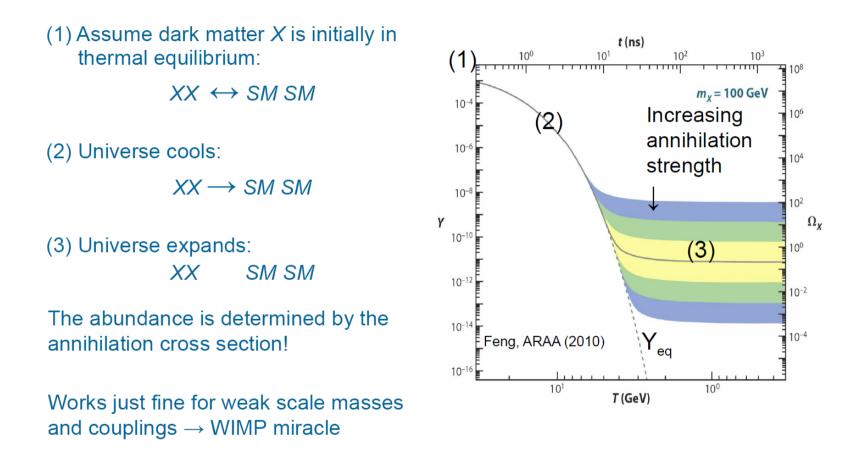
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### Introduction

### Freeze out - WIMPs



Unitarity bound:  $m_{_{DM}} \le 100 \text{ TeV}$ 



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**ICL** 

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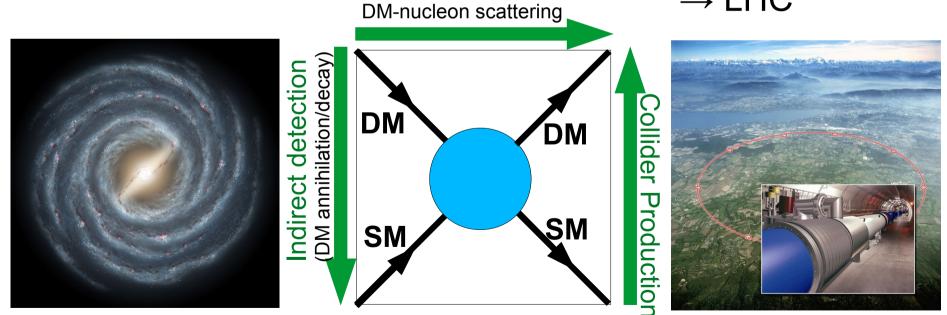


# Introduction

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- Dark Matter (DM) well established:
  - Galaxy rotation
  - CMB measurements

- Three detection methods
  - Direct
  - Indirect
  - Production in collisions
    - $\rightarrow$  LHC



direct detection



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### **Indirect Detection**

### Indirect Detection: Challenges

Astrophysical backgrounds = Large Systematic Uncertainties

Challenging to model theoretically (e.g., cosmic ray diffusion)

New surprises as experimental sensitivities continue to improve (e.g., millisecond pulsars)

How will we ever know that we have discovered dark matter?

Signal detection in more than one target

Correlate gamma-ray signal with large-scale structure that traces the DM signal

ML, Mishra-Sharma, Rodd, Safdi [in progress]

Mariaangela Lisanti

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### **InDirect Detection**



### GeV Photon Excess

Observed at the Galactic Center and Inner Galaxy (≤ 10°)

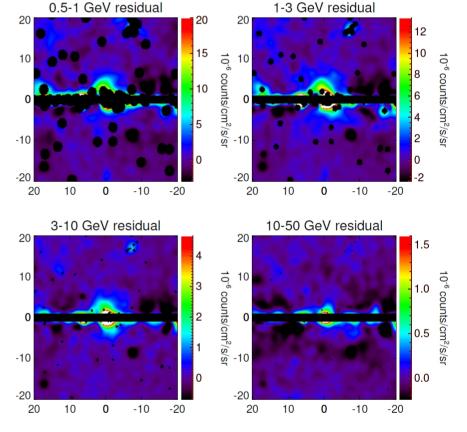
Constitutes ~10% total flux

High statistical significance

Goodenough and Hooper [0910.2998] Hooper and Goodenough [1010.2752] Boyarsky, Malyshev, Ruchayskiy [1012.5839] Hooper and Linden [1110.0006] Abazajian and Kaplinghat [1207.6047] Gordon and Macias [1306.5725] Abazajian *et al.* [1402.4090] Daylan *et al.* [1402.6703] Calore, Cholis, and Weniger [1409.0042] *Fermi* Collaboration [1511.02938]



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Daylan et al. [1402.6703]

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### Introduction

### direct detection

- vector(DM)-vector(SM)
  - stringent limits from spin-independent direct detection
  - best limit:  $\mathcal{O}(10^{-45} \text{cm}^2)$  by LUX
- axial(DM)-axial(SM)
  - not quite so stringent limits from spin-dependent direct detection
  - best limit neutron  $\mathcal{O}(10^{-40} \text{cm}^2)$  by LUX
  - best limit proton  $\mathcal{O}(10^{-39} \text{ cm}^2)$  by PICO
- vector(DM)-axial(SM)
  - $\sigma \propto v^2$  or  $q^2$  and direct detection very suppressed
  - essentially no limit (see Del Nobile, Cirelli, Panci 1307.5955 for actual limit)
- axial(DM)-vector(SM)
  - $\sigma \propto v^2$  or  $q^2$  and direct detection very suppressed
  - essentially no limit

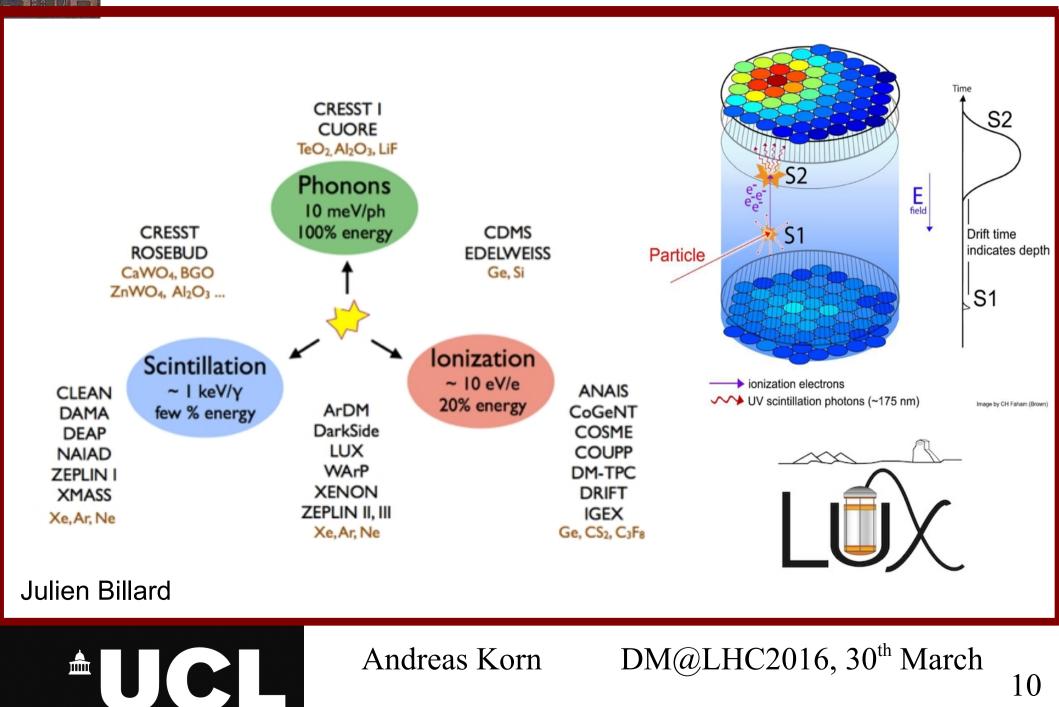
### S. Vogel

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### **Direct Detection**



# - R

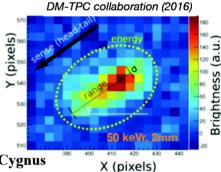
# **Direct Detection: Direction!**

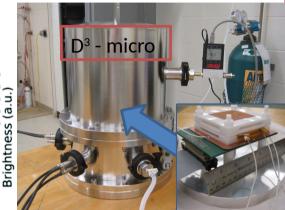
### Directional detection

Directional detection aims at measuring both the recoil energy and direction using gas TPC

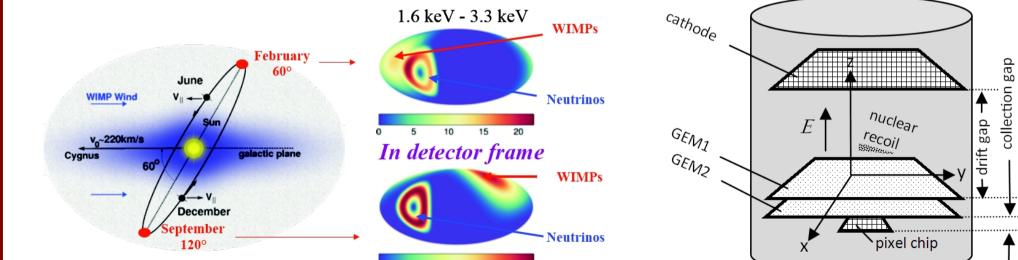
Leading experiments are: **DRIFT**, **DM-TPC**, **MIMAC** and **Newage** Great experimental challenges:

Intrinsic angular resolution: ~20 degrees RMS Thresholds: 1 keV (energy), 10 keV (directional) Exposure: ~100 g (DRIFT) Target: Fluorine (excellent spin-dependent coupling)





Thanks to the rotation of Solar System around Galactic Center, WIMPs are coming from Cygnus
Solar neutrinos are coming from ... the Sun



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Julien Billard

**UC** 

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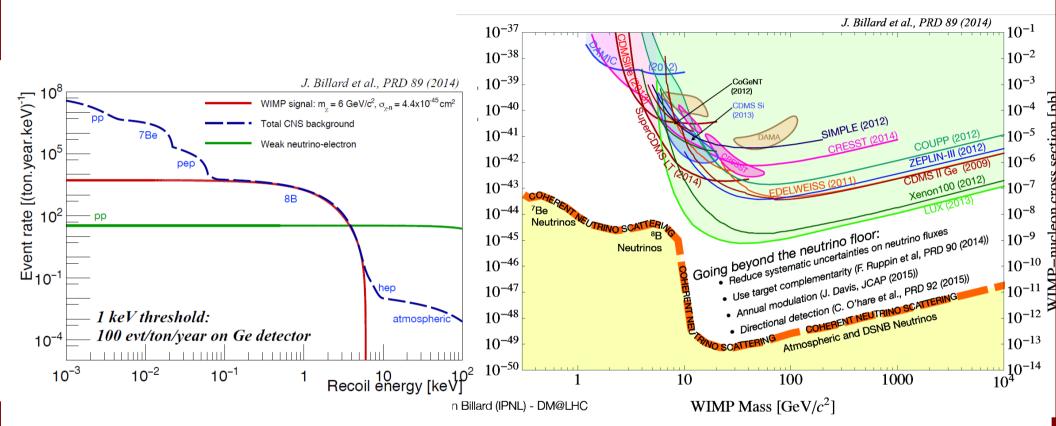
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transfer gap



### **Direct Detection**



### The neutrino background

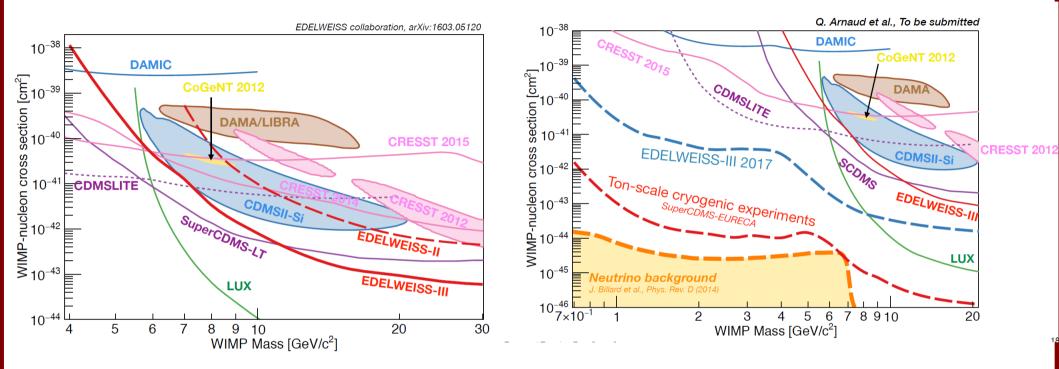
Julien Billard

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### **Direct Detection**



### Julien Billard

**L** 

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# DM at the LHC

### 1. Monojet

Strategy: reconstruction, analysis & interpretation

### 2. Other Mono-X

- ≻Mono-V/γ
- ≻Mono-H
- ≻DM+HF
- ≻H→inv
- 3. Alice+LHCb
  - ▷ Dark photons
  - Dark sectors
- 4. Cosmological constraints
  - ➢ For LHC & FCC
- 5. Resonances constraints

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Tristan du Pree

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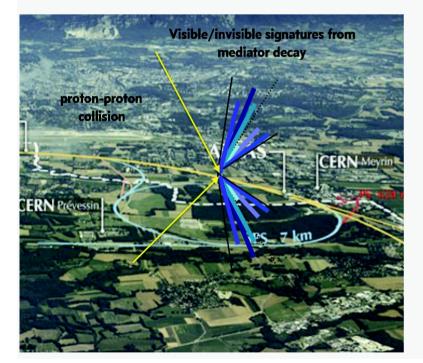


# DM at the LHC

#### Gathering the community for a view of the landscape

#### ATLAS/CMS Dark Matter Forum (2015)

Determined recommended benchmark models for LHC Run 2 searches: emphasis on mediators that could be produced and discovered in the next few years



#### Joint ATLAS/CMS/theory discussion forum

- Classify benchmark models (simplified models) according to final state signatures
- Propose a small set of simplified models for early Run-2 LHC searches
- Review tools and implementations, stateof-the-art calculations
- Maintain model repository

### A. Boveia/S. Malik



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## DM at the LHC

#### LHC Dark Matter Working Group (2015-)

http://lpcc.web.cern.ch/lpcc/index.php? page=dm\_wg

The LHC Dark Matter Working Group (LHC DM WG) brings together theorists and experimentalists to define guidelines and recommendations for the benchmark models, interpretation, and characterisation necessary for broad and systematic searches for dark matter at the LHC.

The LHC DM WG develops and maintains close connections with theorists and other experimental particle DM searches (e.g. Direct and Indirect Detection experiments) in order to help verify and constrain particle physics models of astrophysical excesses, to understand how collider searches and noncollider experiments complement one another, and to help build a comprehensive understanding of viable dark matter models.

#### arXiv:1603.04156

CERN-LPCC-2016-001

Recommendations on presenting LHC searches for missing transverse energy signals using simplified *s*-channel models

of dark matter

14 Mar 2016

[hep-ex]

arXiv:1603.04156v1

Antonio Boveia,<sup>1,\*</sup> Oliver Buchmueller,<sup>2,\*</sup> Giorgio Busoni,<sup>3</sup> Francesco D'Eramo,<sup>4</sup> Albert De Roeck,<sup>1,5</sup> Andrea De Simone,<sup>6</sup> Caterina Doglioni,<sup>7,\*</sup> Matthew J. Dolan,<sup>3</sup> Marie-Helene Genest,<sup>8</sup> Kristian Hahn,<sup>9,\*</sup> Ulrich Haisch,<sup>10,11,\*</sup> Philip C. Harris,<sup>1</sup> Jan Heisg,<sup>12</sup> Valerio Ippolito,<sup>13</sup> Felix Kahlhoefer,<sup>14,\*</sup> Valentin V. Khoze,<sup>16</sup> Suchita Kulkarni,<sup>16</sup> Greg Landsberg,<sup>17</sup> Steven Lowette,<sup>18</sup> Sarah Malik,<sup>2</sup> Michelangelo Mangano,<sup>11,\*</sup> Christopher McCabe,<sup>19,\*</sup> Stephen Mrenna,<sup>30</sup> Priscilla Pani,<sup>21</sup> Tristan du Pree,<sup>1</sup> Antonio Riotto,<sup>11</sup> David Salek,<sup>19,22</sup> Kai Schmidt-Hoberg,<sup>14</sup> William Shepherd,<sup>23</sup> Tim M.P. Tait,<sup>21,\*</sup> Lian-Tao Wang,<sup>35</sup> Steven Worm<sup>26</sup> and Kathryn Zurek<sup>27</sup>

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 <sup>3</sup>ARC Centre of Excellence for Particle Physics at the Terascale, School of Physics, University of Melbourne, 3010, Australia
 <sup>4</sup>UC, Santa Cruz, and UC, Santa Cruz, Inst. Part. Phys., USA
 <sup>5</sup>Antwerp University, B2610 Wilrijk, Belgium
 <sup>6</sup>SISSA and INFN Sectione di Trieste, via Bonomea 265, I-34136 Trieste, Italy
 <sup>7</sup>Pysiska Institutionen, Lundus universitet, Lund, Sweden
 <sup>8</sup>LPSC, Universite Grenobie-Alpes, CNRS/IN2P3, France
 <sup>9</sup>Department of Physics and Astronomy, Northwestern University, Evanston, Illinois 60208, USA
 <sup>10</sup>Rudoff Pierie Centre for Theoretical Physics, University of Oxford, Oxford, OX1 3PN, United Kingdom

**Pilot effort:** translation of LHC simplified model results into DM-nucleon cross-section plane (DD/ID)

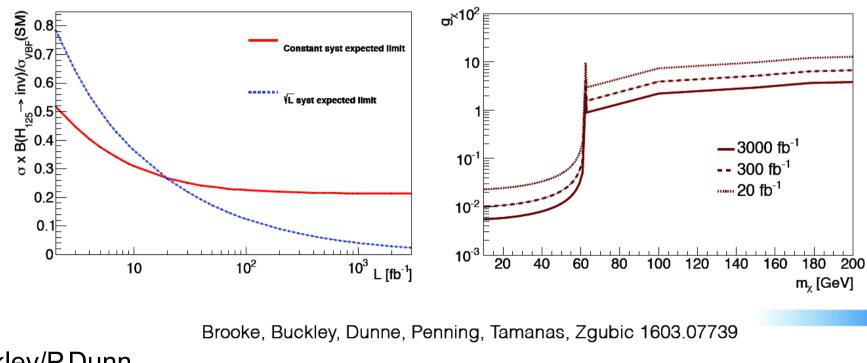
### A. Boveia/O. Buchmueller



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# Higgs $\rightarrow$ invisible

- This Higgs boson thing seems like it could be useful  $pp \to (h^{(*)} \to \chi \chi) + X$ 
  - Or through extra Higgs(es) in non-minimal models
- Combined limit on  $BR(H_{125} \rightarrow \text{ inv.}) \lesssim 0.25$



### M. Buckley/P.Dunn

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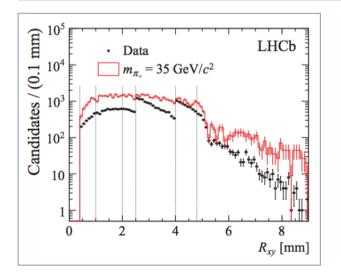
### More exotic searches

#### Tristan du Pree (CERN), DM@LHC (1 April 2016)

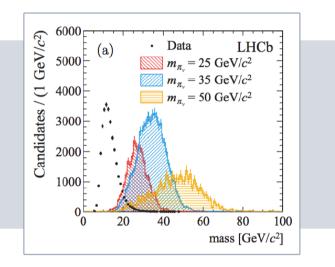
#### 38

### **Displaced jets**

- $H \rightarrow \pi_v(bb)\pi_v(bb)$ 
  - > Hidden valley long-lived particles
- Low mass
  - >25 GeV 50 GeV
- Displaced bb
  - ≥0.4 mm 4.8 mm



#### Eur.Phys.J.C75(2015)152



#### LHCb advantages

- Triggers: low-mass&p<sub>T</sub>
   > Upgrade: full software trigger
- Vertex resolution
  - > Critical for displaced searches

DM@LHC: see Swagata Mukherjee for Atlas+CMS long lived

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# DM: Latest results in the mono-jet and di-jet channels



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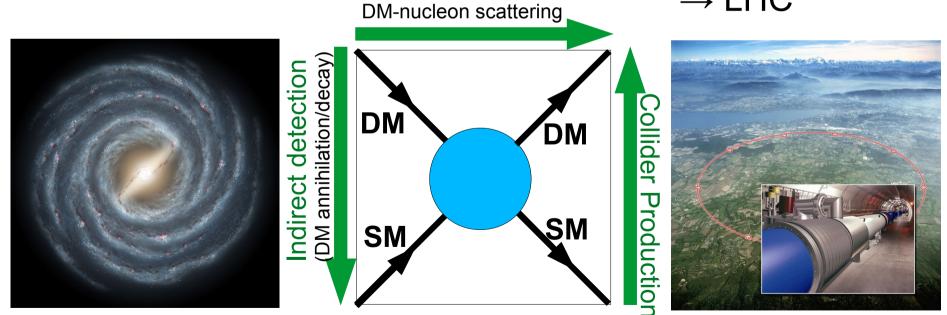


# Introduction

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    - $\rightarrow$  LHC



direct detection



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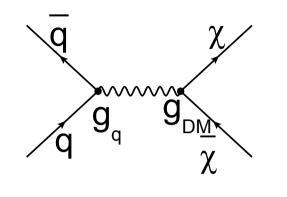


# DM at LHC:Models

complex (Majorana) Dirac	Name           C1           C5           D1           D5           D8           D9	Initial state $qq$ $gg$ $qq$ $qq$ $qq$ $qq$ $qq$ $qq$	Type scalar scalar scalar vector axial-vector tensor	Operator $\frac{\overline{m}_{q}}{M_{\star}^{2}}\chi^{\dagger}\chi\bar{q}q$ $\frac{1}{4M_{\star}^{2}}\chi^{\dagger}\chi\alpha_{s}(G_{\mu\nu}^{a})^{2}$ $\frac{\overline{m}_{q}}{M_{\star}^{3}}\bar{\chi}\chi\bar{q}q$ $\frac{1}{M_{\star}^{2}}\bar{\chi}\gamma^{\mu}\chi\bar{q}\gamma_{\mu}q$ $\frac{1}{M_{\star}^{2}}\bar{\chi}\gamma^{\mu}\gamma^{5}\chi\bar{q}\gamma_{\mu}\gamma^{5}q$ $\frac{1}{M_{\star}^{2}}\bar{\chi}\sigma^{\mu\nu}\chi\bar{q}\sigma_{\mu\nu}q$	• Common: Effective Field Theory (EFT) • Limited validity: $m_{mediator} >> E(\chi)$ • Integrate mediator out, Fermi constant $G_{F}$ like coupling: $G_{DM} = \left(\frac{\sqrt{g_q g_{DM}}}{M_{mediator}}\right) = \frac{1}{(M_*)^n}$
q	D11 EFT	$\chi$	scalar	$\frac{\frac{1}{4M_{\star}^{3}}\bar{\chi}\chi\alpha_{s}(G_{\mu\nu}^{a})^{2}}{g_{DM}}$	• Now mostly superseeded
		CL	A	ndreas Korn	DM@LHC2016, 30 <sup>th</sup> March 21



# DM at LHC:Models



$$\mathcal{L}_{\text{vector}} = g_{q} \sum_{q=u,d,s,c,b,t} Z'_{\mu} \bar{q} \gamma^{\mu} q + g_{\chi} Z'_{\mu} \bar{\chi} \gamma^{\mu} \chi$$

 $\mathcal{L}_{\text{axial-vector}} = g_{q} \sum_{q=u,d,s,c,b,t} Z'_{\mu} \bar{q} \gamma^{\mu} \gamma^{5} q + g_{\chi} Z'_{\mu} \bar{\chi} \gamma^{\mu} \gamma^{5} \chi.$ 

$$\mathcal{L}_{\text{scalar}} = -g_{\text{DM}}\phi\bar{\chi}\chi - g_q \frac{\phi}{\sqrt{2}} \sum_{q=u,d,s,c,b,t} y_q \bar{q}q ,$$
$$\mathcal{L}_{\text{pseudo-scalar}} = -ig_{\text{DM}}\phi\bar{\chi}\gamma_5\chi - ig_q \frac{\phi}{\sqrt{2}} \sum_{q=u,d,s,c,b,t} y_q \bar{q}\gamma_5q$$

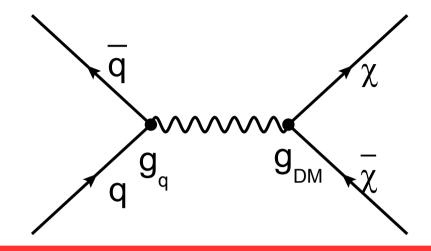
Move towards simplified models

- → recommendations from DM@LHC14 (arXiv:hep-ph/1506.03116)
- → recommendations from LHCDMWG (arXiv:hep-ex/1603.04156, arXiv:hep-ex/1507.00966)
- Parameters:  $m_{DM}$ ,  $m_{mediator}$ ,  $g_{DM}$ ,  $g_{q}$
- Benchmark models
- m<sub>DM</sub> vs m<sub>mediator</sub> plane, couplings
- Vector :  $g_{DM} = 1$ ;  $g_{q} = 0.25$
- Axial-Vector :  $g_{DM} = 1$ ;  $g_{q} = 0.25$
- Scalar :  $g_{DM} = 1$ ;  $g_{q} = 1$
- Pseudo-Scalar:  $g_{DM} = 1$ ;  $g_{q} = 1$

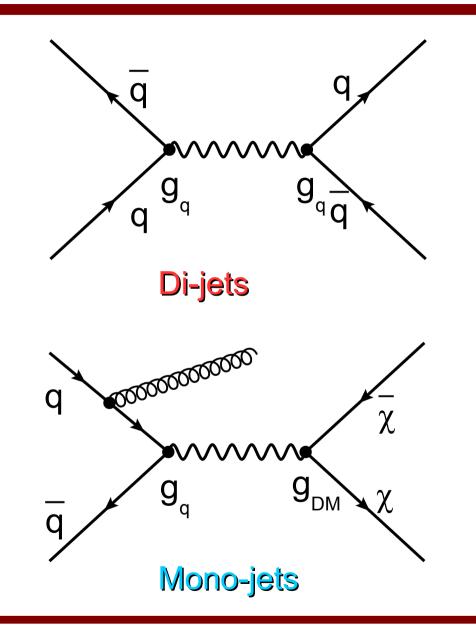
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### DM at LHC



- DM  $\chi$  couples loosely to SM particles (quarks q) through a mediator
- $\rightarrow$  mediator couples to DM  $\chi$  with g<sub>DM</sub> and to SM quark with g<sub>q</sub>
- Can't reconstruct DM in detector
   → need accompanying signature
- Mediator can decay into quark (jet) pairs

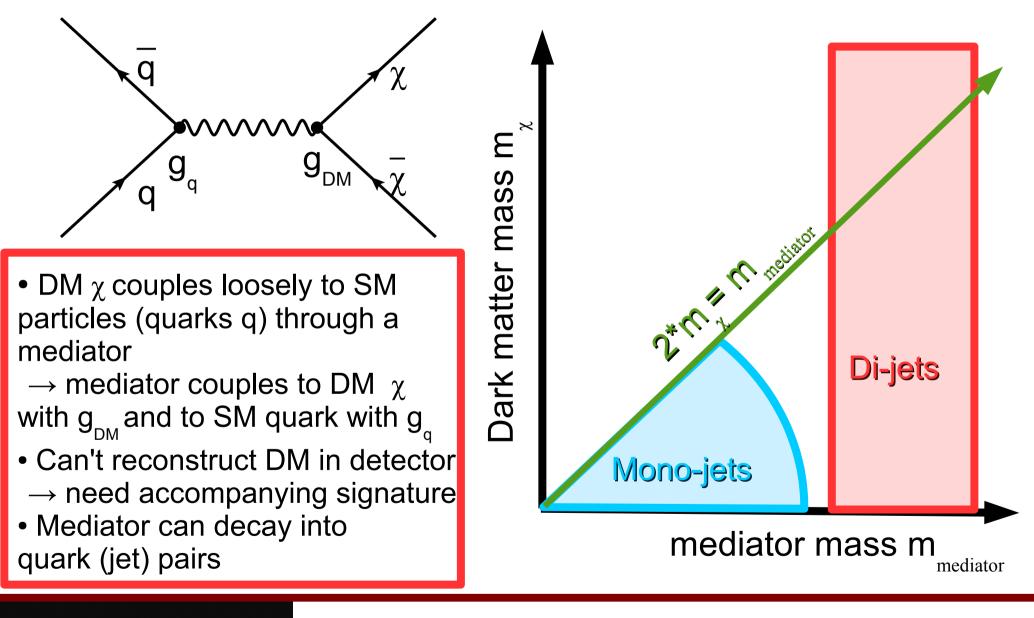


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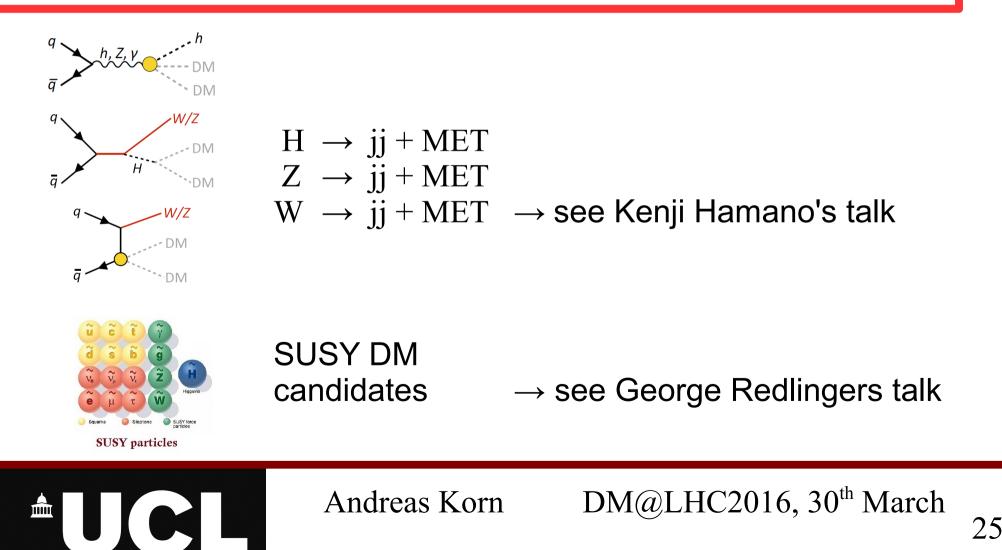
### DM at LHC



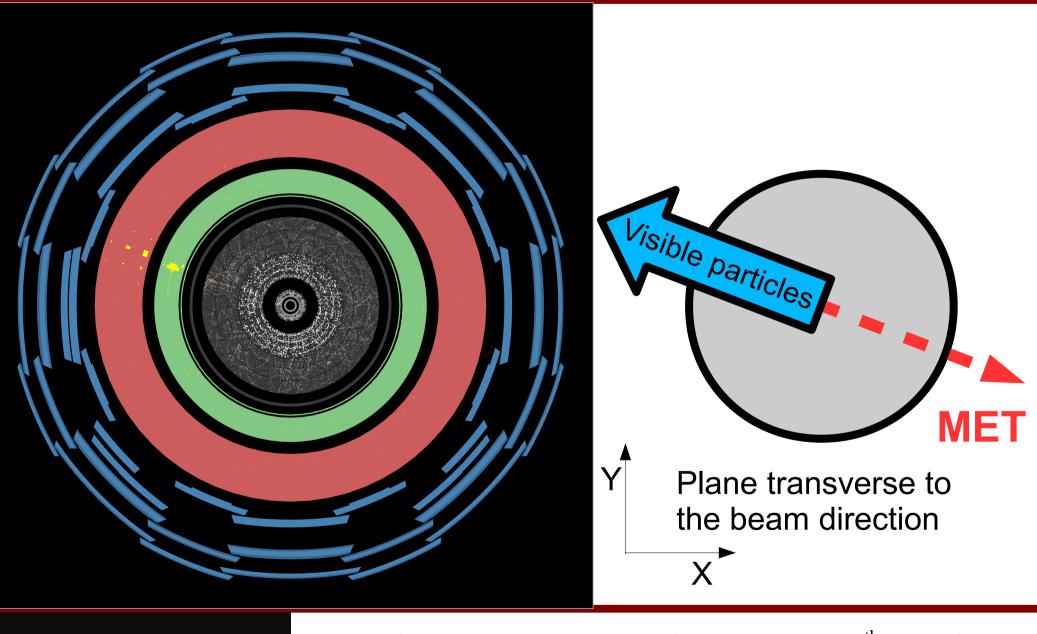


# Other DM signals at the LHC

- Can't reconstruct DM in detector
   → missing momentum → need accompanying signature
- Differentiate channels by accompanying signature

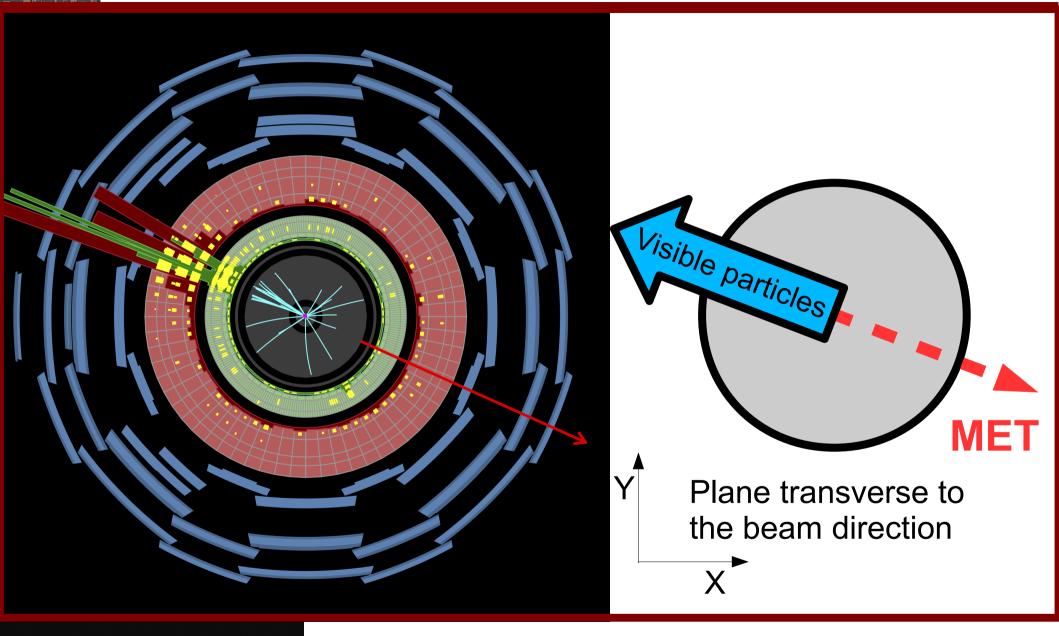


# Missing transverse Energy MET



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# Missing transverse Energy MET

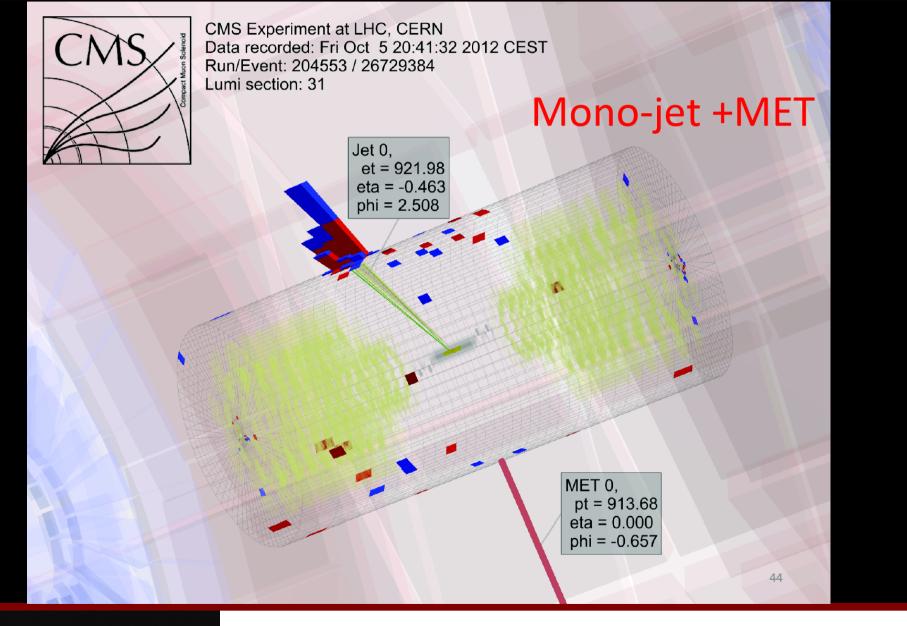




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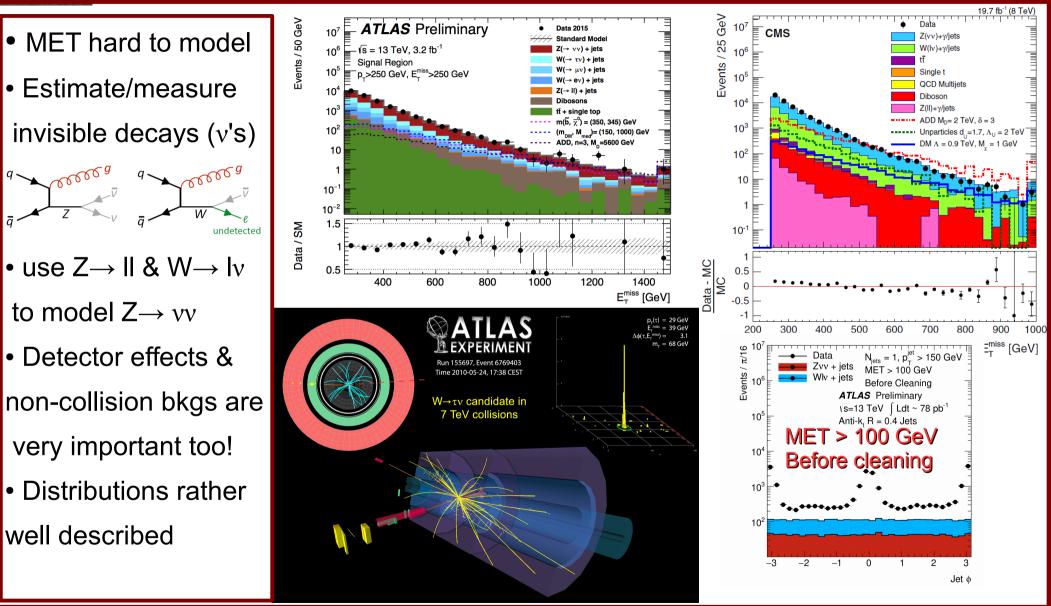
### Mono-Jet



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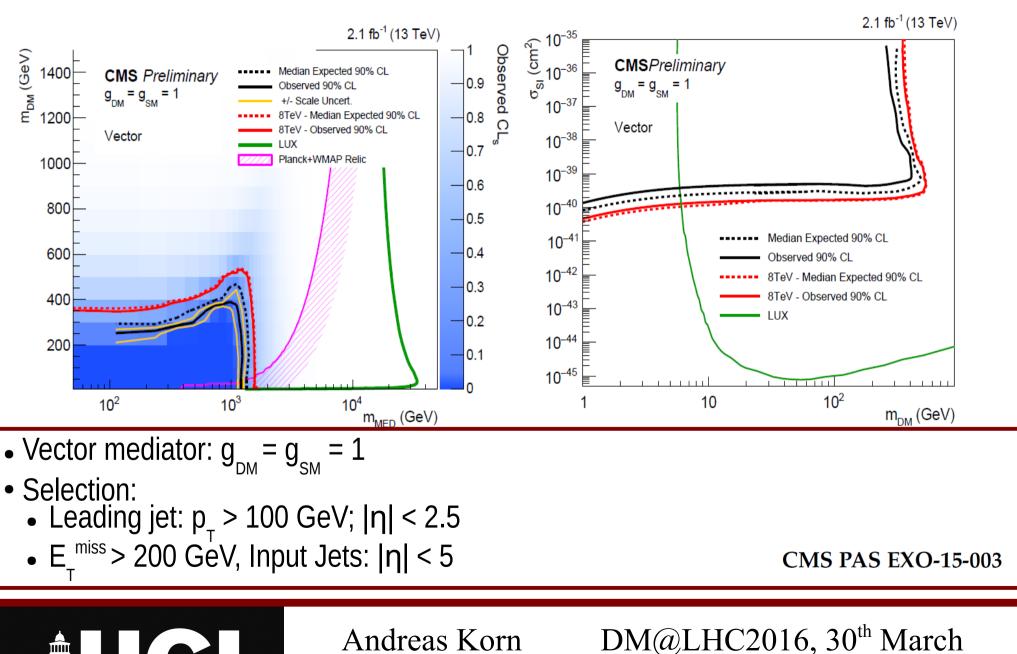
# Mono-Jet: Backgrounds



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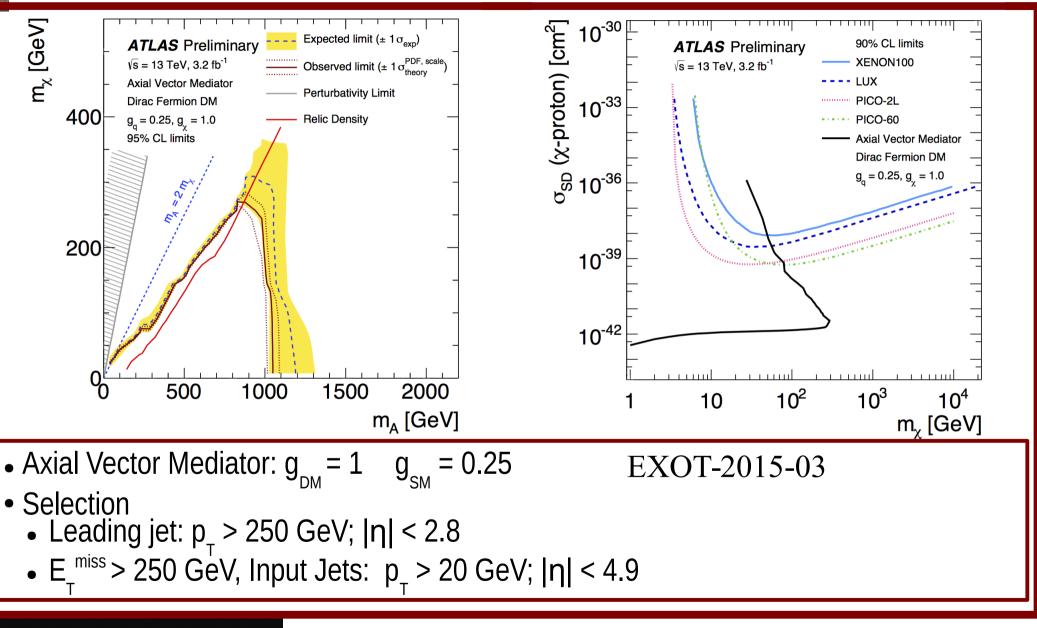
### Mono-Jet





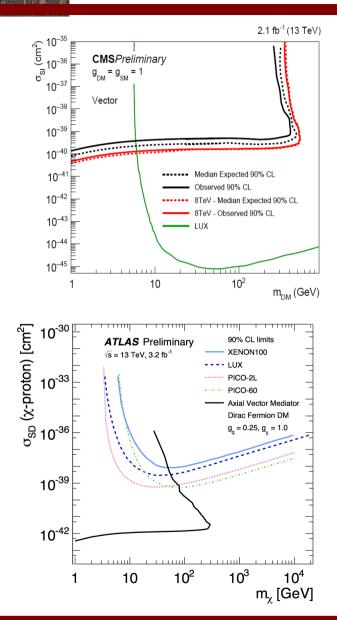
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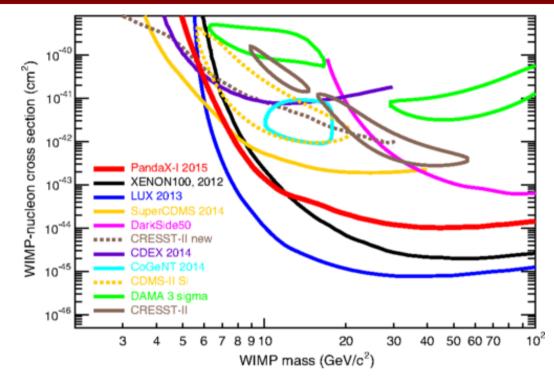
### Mono-Jet



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# **Comparison with Direct Detection**





- Translate into DM-nucleon cross sections
- Spin dependent (SD) or independent (SI) according to mediator model used
- Note: Comparisons model dependent!

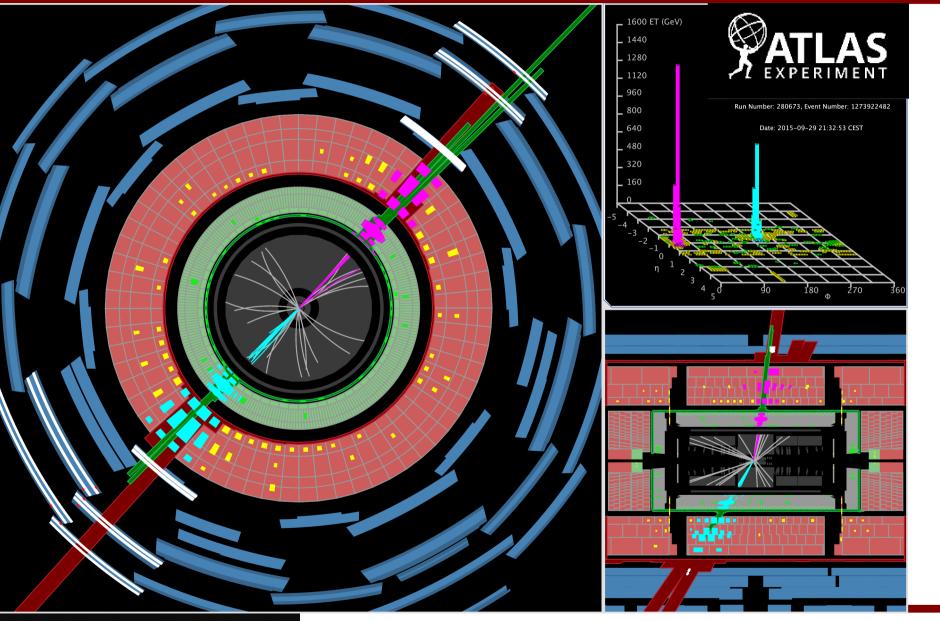
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• LHC experiments provide complementary coverage!





### **DiJet Events**



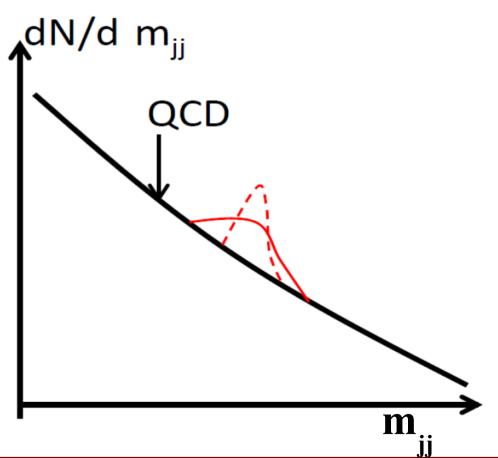


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# + BANK BANK

# Introduction: Dijet resonances

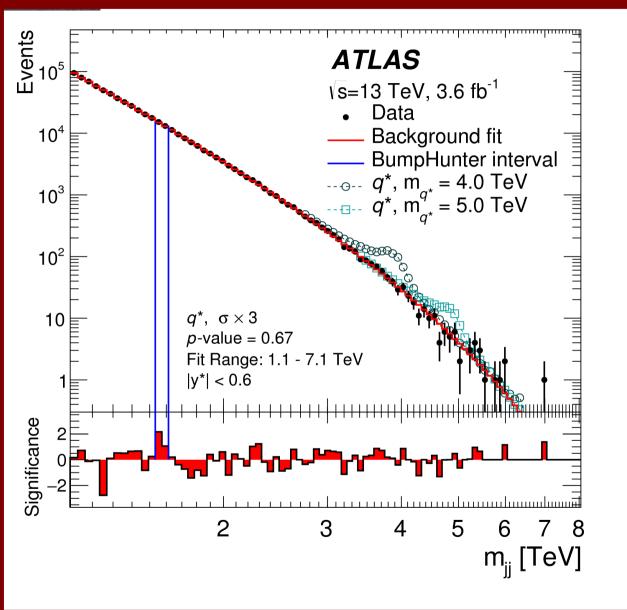
- Look for resonant qq, gq and gg states
- Benchmark search for new physics
- Construct dijet mass
- Fit smooth spectrum  $f(x) = p_1(1-x)^{p_2}x^{p_3+p_4\ln x}$ • Look for deviations
- Look for deviations
   → Bumphunter
   Toilburgtor
  - $\rightarrow$  Tailhunter
- Limits on acceptance times x-section and specific models





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### **Dijet resonances**



- Selection
  - p<sub>1</sub> > 440 GeV

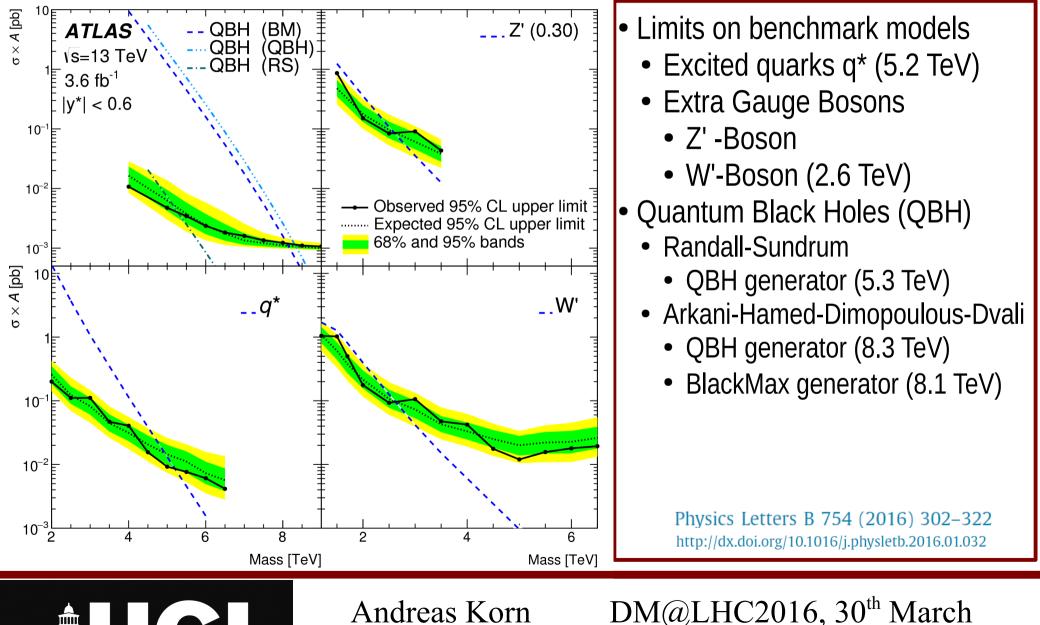
• 
$$|y^*| = \frac{1}{2} |y_1 - y_2| < 0.6$$

- Background from fit
- BumpHunter indicates most discrepant interval (not so exciting at all)

Physics Letters B 754 (2016) 302–322 http://dx.doi.org/10.1016/j.physletb.2016.01.032

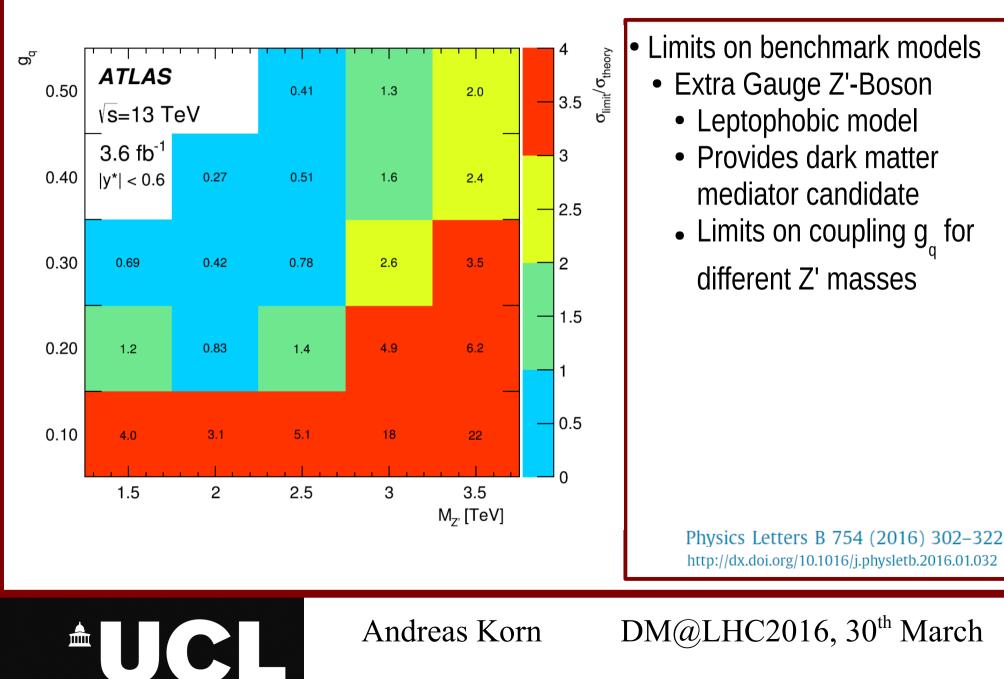
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## Limits on Dijet resonances

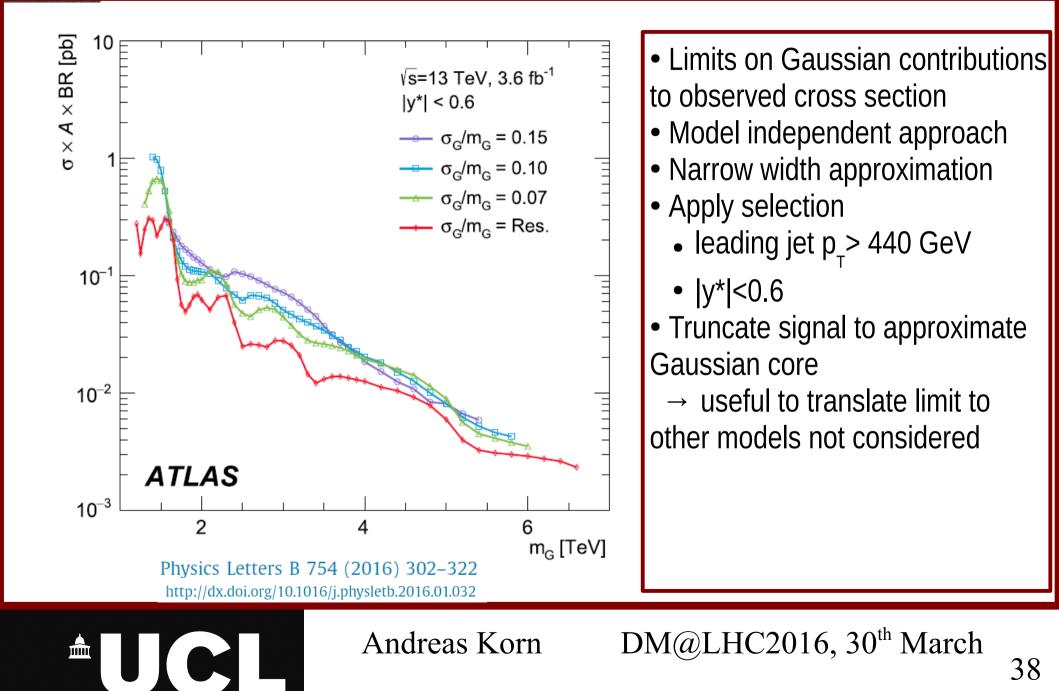




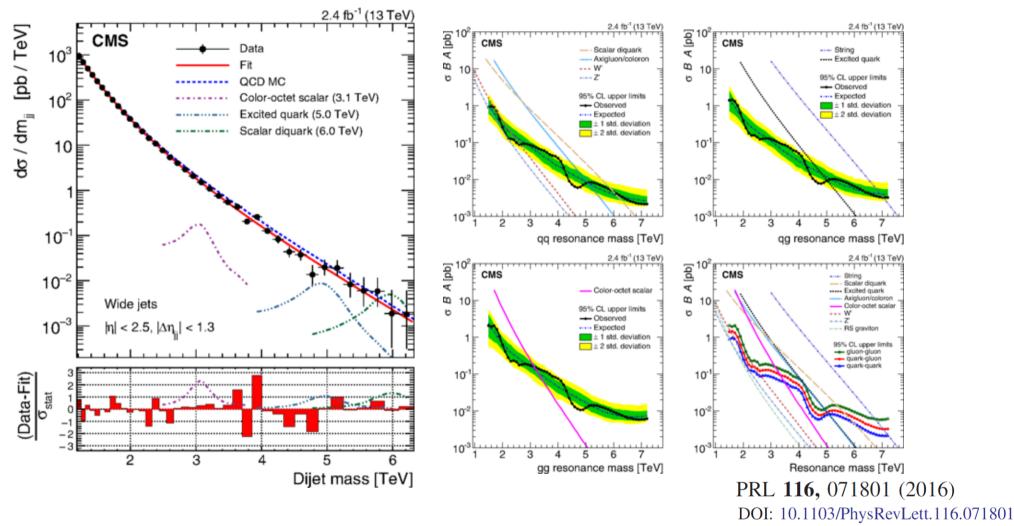
## Limits on Dijet resonances



## Limits on Dijet resonances



## Limits on Dijet resonances

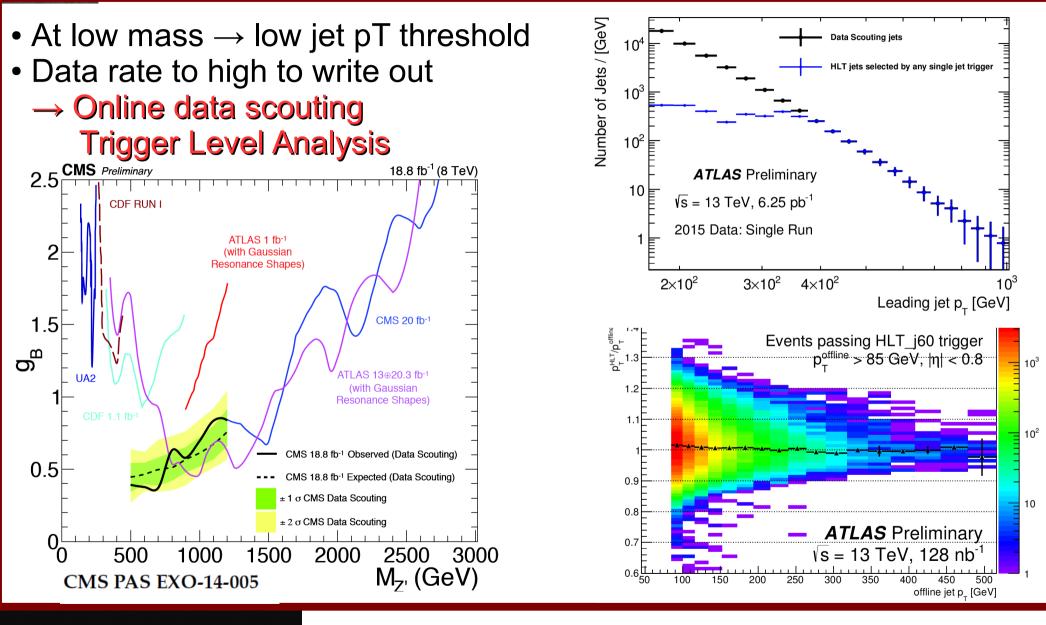


•Excluded masses: String resonance (7.0 TeV) , scalar di-quark (6.0 TeV), axigluon (5.1 TeV), excited quark q\* (5.0 TeV), Heavy W' (2.6 TeV)

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## Dijet Events at low mass

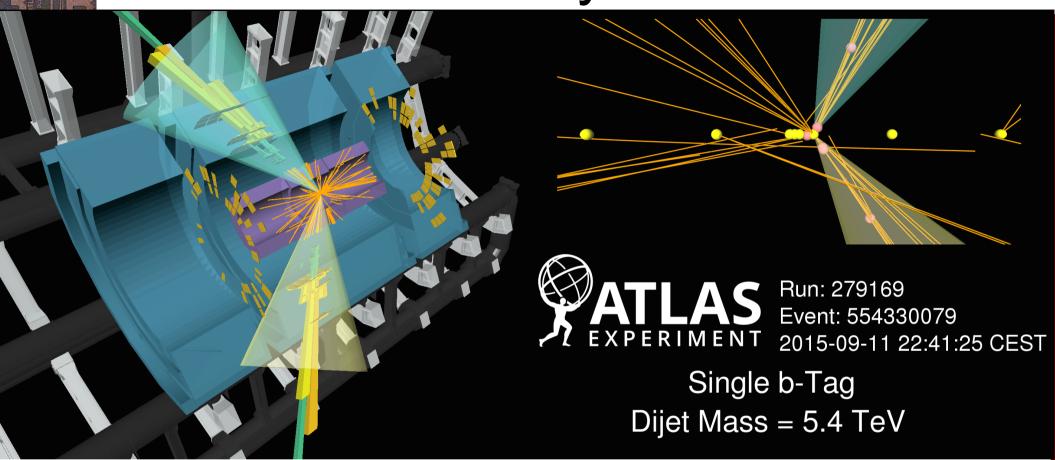


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## **Di-beauty-Jets**



• Third Generation (top & bottom) heavy, might be special  $\rightarrow$  investigate couplings to b

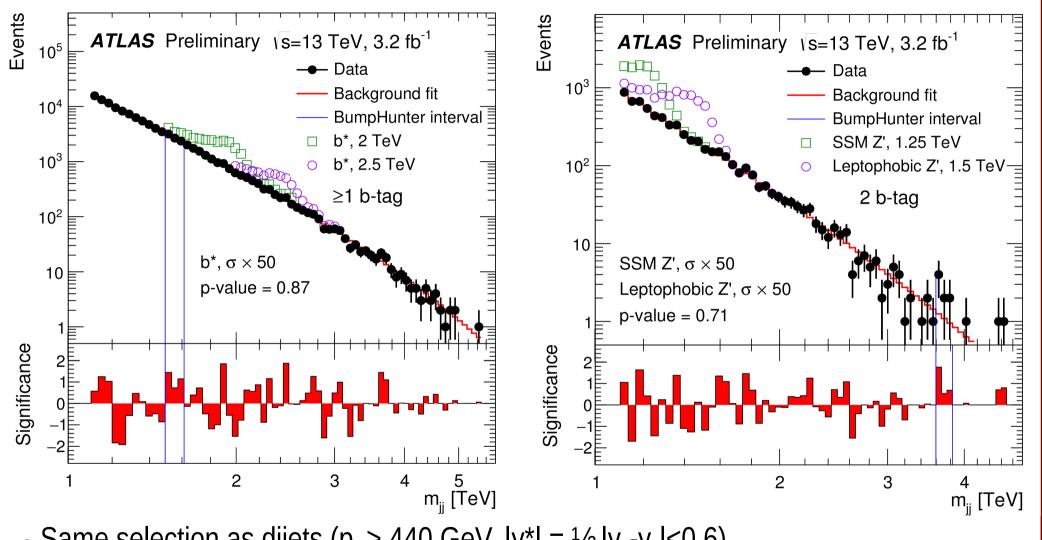
- Needs identification of jets containing bottom hadrons  $\rightarrow$  b-tagging
- Depending on decay (bb, bq, bg)  $\rightarrow$  at least 1 or 2 b-tags

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• Possible qg background reduction also for  $X \rightarrow qq$  modes

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## **Di-beauty-Jets**



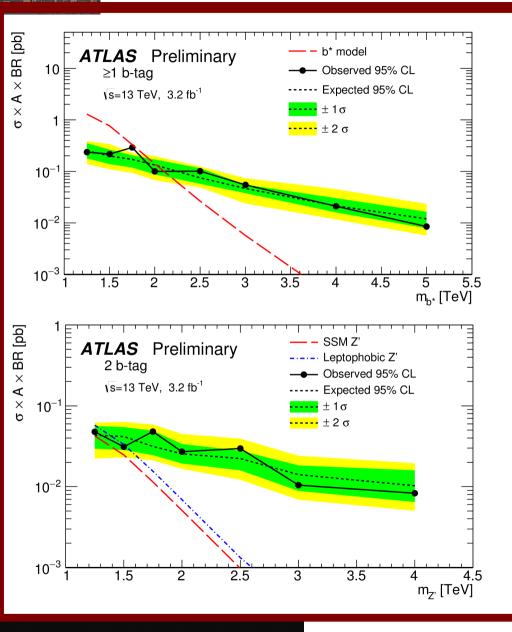
• Same selection as dijets ( $p_T > 440$  GeV,  $|y^*| = \frac{1}{2} |y_1 - y_2| < 0.6$ ) • Limit  $|\eta| < 2.5$ , to tracking coverage for b-tagging ar

arXiv:1603.08791v1 [hep-ex]

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## **Di-beauty-Jets**

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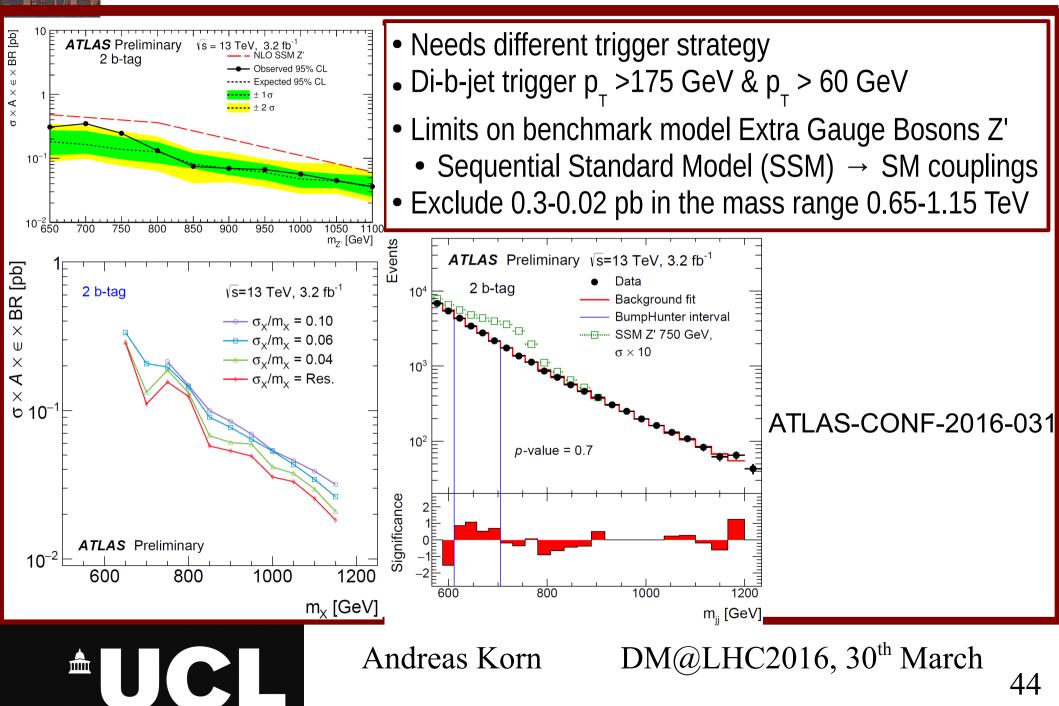
- Limits on benchmark models
- Excited quarks  $b^* \rightarrow bg$ 
  - >= 1 b-tag
  - Excluded masses 1.1-2.1 TeV

### •Extra Gauge Bosons Z'

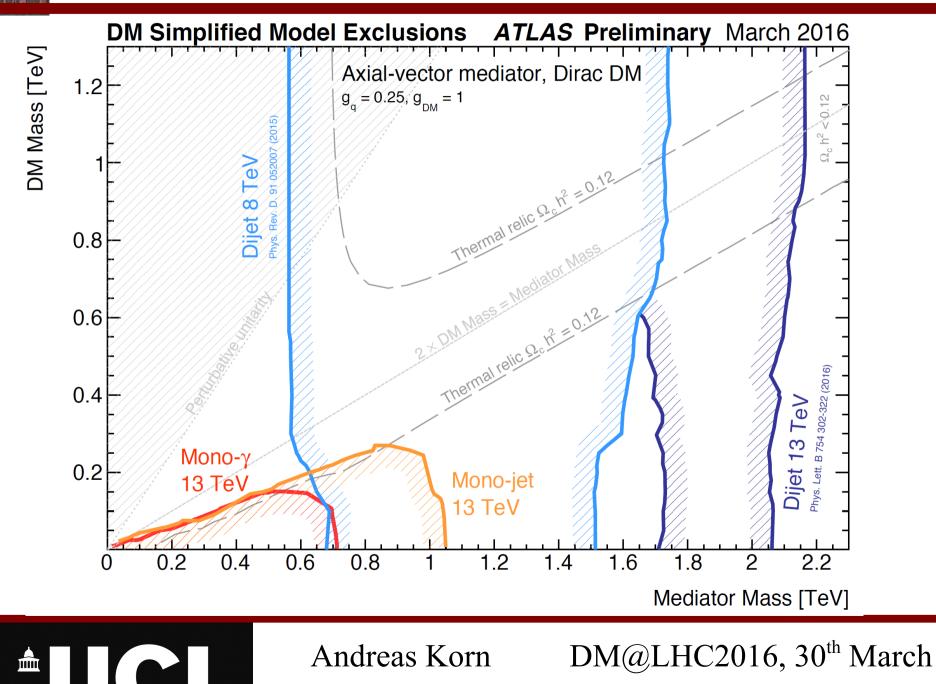
- 2 b-tag
- Leptophobic Z'
- Excluded masses 1.1-1.5 TeV
- Sequential Standard Model (SSM) → SM couplings
- Not enough data to exclude Sequential SM Z'

arXiv:1603.08791v1 [hep-ex]

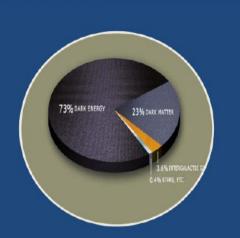
## Di-b-Jets: what about 750 GeV ?



## Summary

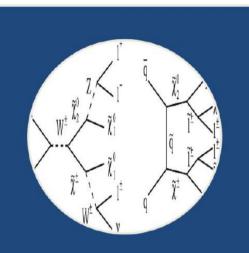


## Putting it together: Global Fits!



#### Evidence from Astroparticle physics

- Dark Matter
- Assumptions



## Theoretical connections

- Supersymmetry
- Extra Dimensions

• ... ' śś



## Consequences for LHC

- LHC phenomenology
- Model testing

S. Caron

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Andreas Korn

## Putting it together: Global Fits!

#### GAMBIT: a second-generation global fit code

GAMBIT: The Global And Modular BSM Inference Tool

Overriding principles of GAMBIT: flexibility and modularity

- General enough to allow fast definition of new datasets and theoretical models
- Plug and play scanning, physics and likelihood packages
- Extensive model database not just small modifications to constrained MSSM (NUHM, etc), and not just SUSY!
- Extensive observable/data libraries (likelihood modules)
- Many statistical options Bayesian/frequentist, likelihood definitions, scanning algorithms
- A smart and fast LHC likelihood calculator
- Massively parallel
- Full open-source code release



#### Pat Scott

Andreas Korn



## The Future?

#### The Future of Simplified Models

Joachim Kopp



#### Tables, tables, tables

ID	Х	$\alpha + \beta$	$M_s$	Spin	$(SM_1 SM_2)$	$X-DM-SM_3$	$M_s$ –X–X
SU1			(1, 1, 0)	В	$egin{aligned} &(u_R\overline{u_R}),(d_R\overline{d_R}),(\ell_R\overline{\ell_R})\ &(Q_L\overline{Q_L}),(L_L\overline{L_L}),(HH^\dagger) \end{aligned}$	H1	$\checkmark$
SU2		0		F	$(L_L H)$		
SU3			$(1, 3, 0)^{N \ge 2}$	В	$(Q_L \overline{Q_L}), (L_L \overline{L_L}), (H H^{\dagger})$	H1	$\checkmark$
SU4	$(1, N, \alpha)$		(1, 3, 0) =	F	$(L_L H)$		
SU5	$(1, 1^{\vee}, \alpha)$		(1 1 0)	В	$(d_R  \overline{u_R}), (H^\dagger  H^\dagger), (L_L  L_L)$		$\checkmark$

#### Tally

In total 161 simplified models (defined by representations of DM, X and M)

#### 49 s-channel, 105 t-channel, 7 hybrid

1	0U12		-3	(1, 2, -3)	D	$(L_L \iota_R)$	
	SU13			(1, 2, -3)	F	$(\ell_R H^{\dagger})$	
[	SU14	$(1,N\pm2,lpha)$	0	(1, 3, 0)	В	$(Q_L \overline{Q_L}), (L_L \overline{L_L}), (H H^{\dagger})$	$\checkmark(lpha=0)$
	SU15				F	$(L_L H)$	
	SU16		$^{-2}$	(1, 3, -2)	В	$(H^{\dagger} H^{\dagger}), (L_L L_L)$	$\checkmark (\alpha = \pm 1)$
	SU17				F	$(L_L H^{\dagger})$	

DM	in $(1, N, \beta)$ representation of $SU(3) \times SU(2) \times U(1)$
Х	coannihilation partner
$M_{s}$	<i>s</i> -channel mediator
$SM_1, SM_2$	SM particles in coannihilation $DM + X \rightarrow SM_1SM_2$
SM <sub>3</sub>	Possible additional vertex DM–X–SM <sub>3</sub>

#### Joachim Kopp



Andreas Korn



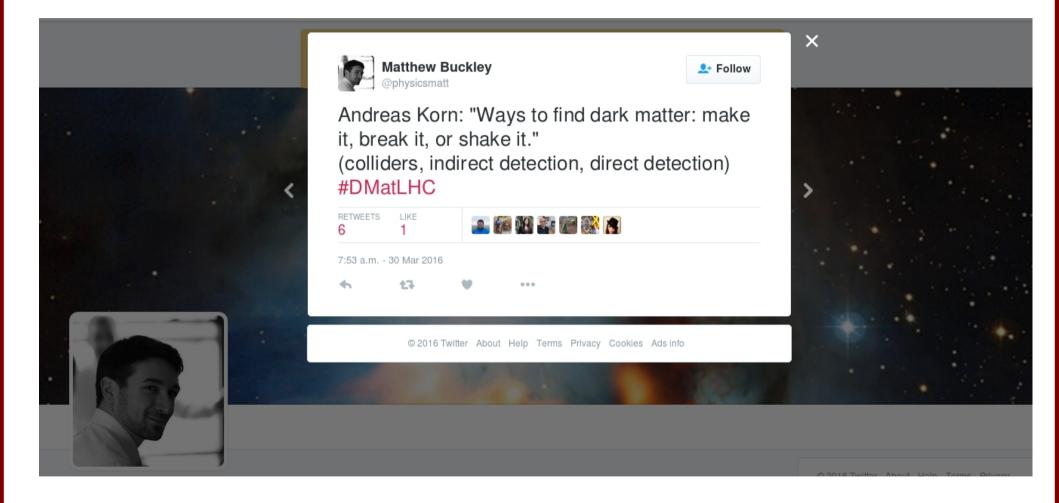
## Conclusion

- The hunt for Dark Matter continues
- Very nice and constructive workshop https://indico.cern.ch/event/342623
- ATLAS and CMS have new interesting results
- Just the beginning of 13 TeV running ...
- LHC searches complementary to direct searches
- Jet channels are particularly sensitive
- Search for both DM and mediator candidates
- Model dependence in Interpretation

   LHCDMWG Recommendations should unify approaches and help comparisons
- Hope presenting hard work on LHC measurements has the right consequences ....



## Consequences



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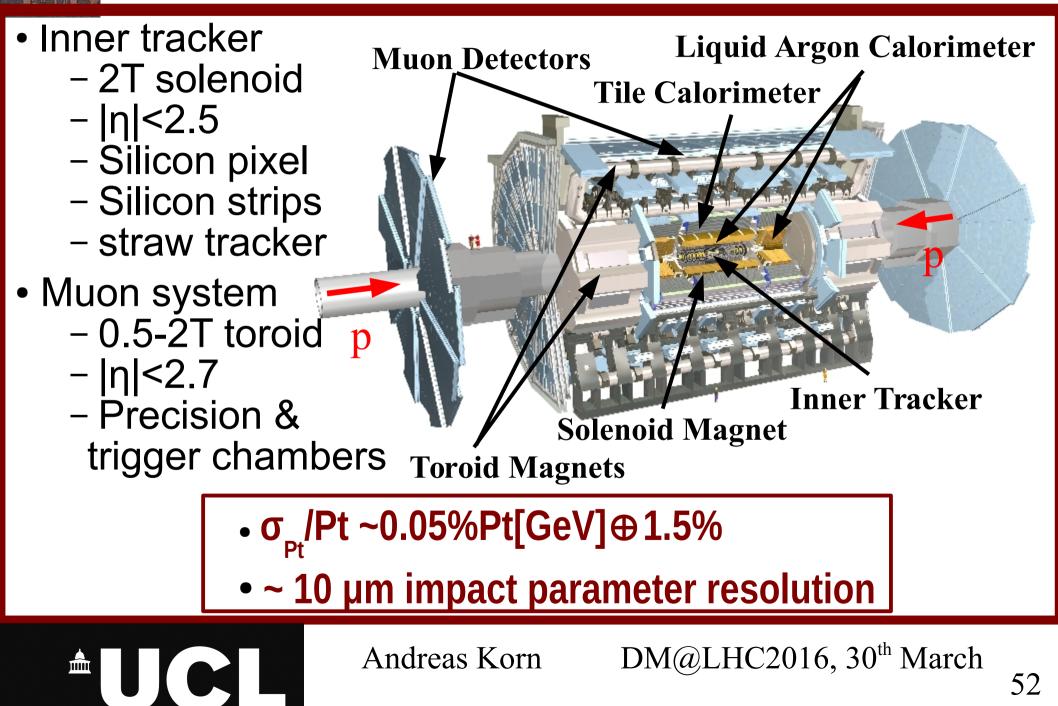


# Bonus Slides



Andreas Korn

## ATLAS: a particle detector at the LHC



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## Technicalities: narrow width

### The narrow-width approximation

width, even for resonances normally considered narrow. The extreme end of this tail due to the PDFs is sometimes suppressed in the searches by requiring the partons to be have mass close to the pole mass, within a few standard deviations on the dijet mass resolution. This is generally a reasonable solution for the shapes, as the QCD background overwhelms the signal at low dijet mass. However, the way that this tail from PDFs is handled can significantly affect the total resonance cross section quoted for specific models, as we discuss in Appendix A

#### Narrow width approximation:

Approximate the true resonance shape with a delta function

• This avoids low-mass tails as PDFs will act only in the surrounding of the peak

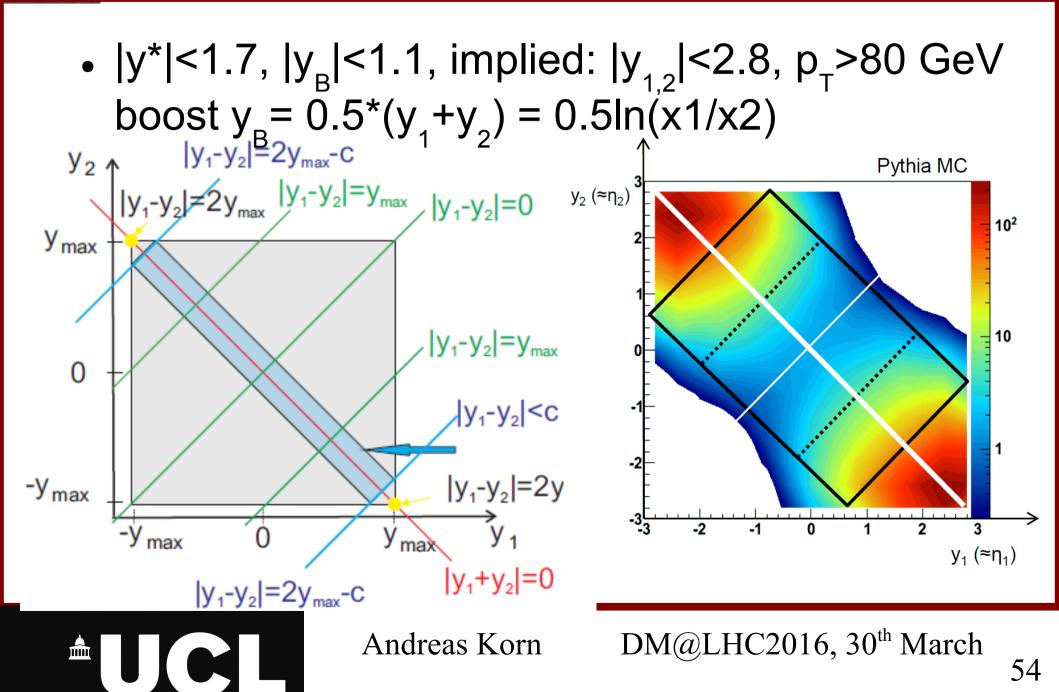
$$\sigma_{had}(m_R) = 16\pi^2 \times \mathcal{N} \times \mathcal{A}_{\cos\theta^*} \times BR \times \left[\frac{1}{s}\frac{dL(\bar{y}_{min}, \bar{y}_{max})}{d\tau}\right]_{\tau = m_R^2/s} \times \frac{\Gamma_R}{m_R}, \quad (44)$$

where the parton luminosity  $\frac{dL}{d\tau}$  is calculated at  $\tau = m_R^2/s$ , and constrained in the inematic range  $[\bar{y}_{min}, \bar{y}_{max}]$ . **Searches for Dijet Resonances at Hadron Colliders** Robert M. Harris, Konstantinos Kousouris arXiv:1110.5302

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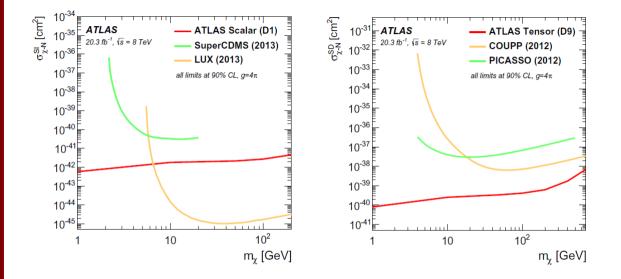
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## Rapidity distribution, Selection





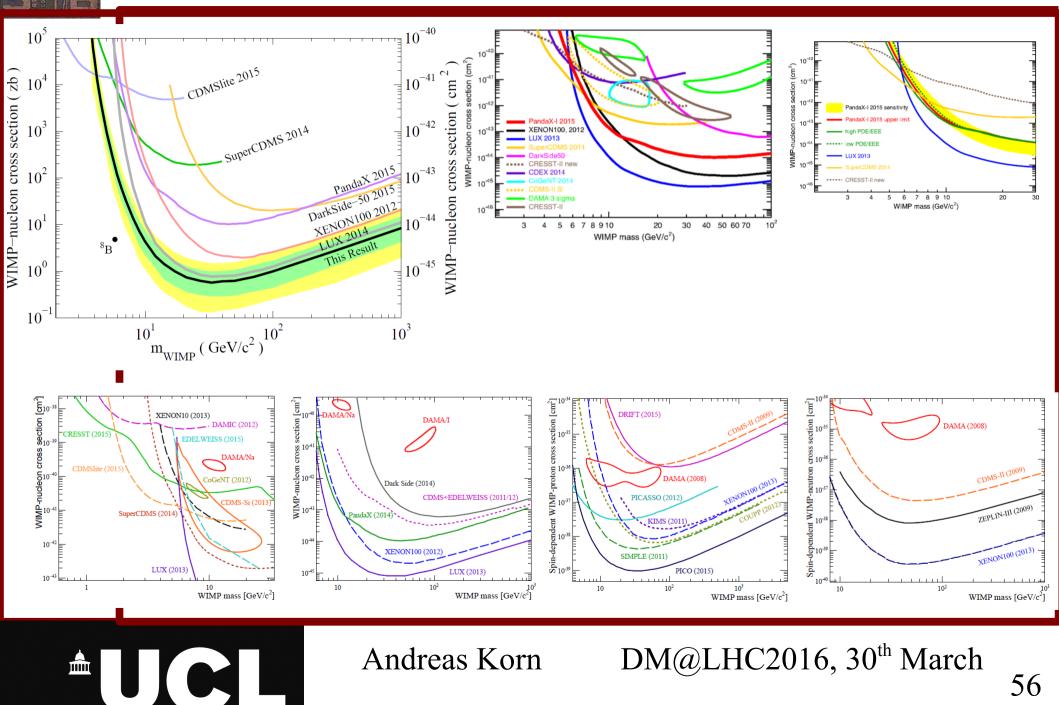
## Beauty-jets and missing E<sub>1</sub>



Eur. Phys. J. C (2015) 75:92 DOI 10.1140/epjc/s10052-015-3306-z

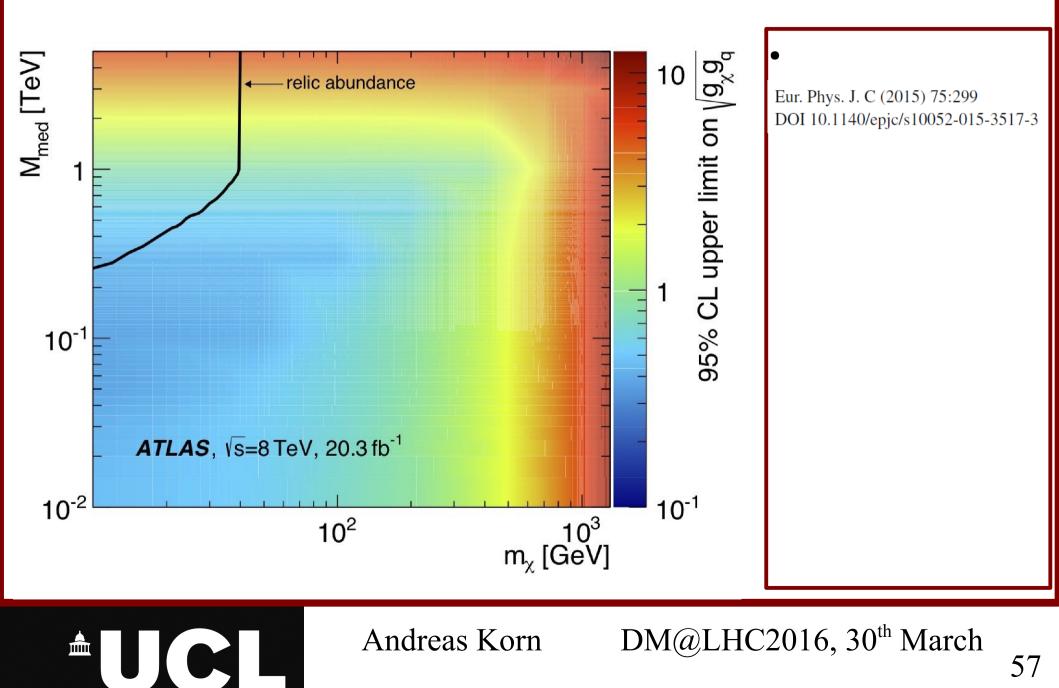
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## **Direct Detection**



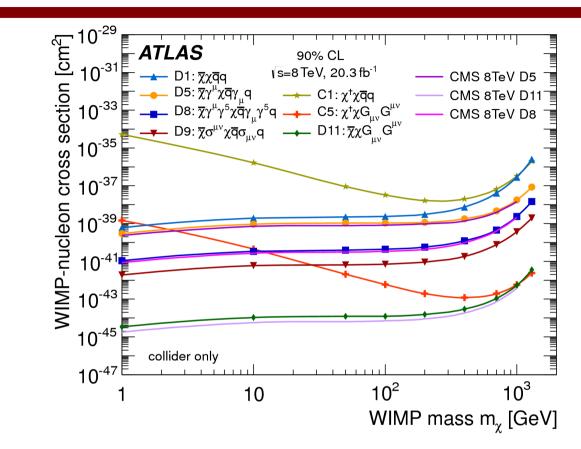


## Mono-Jet





## Mono-Jet

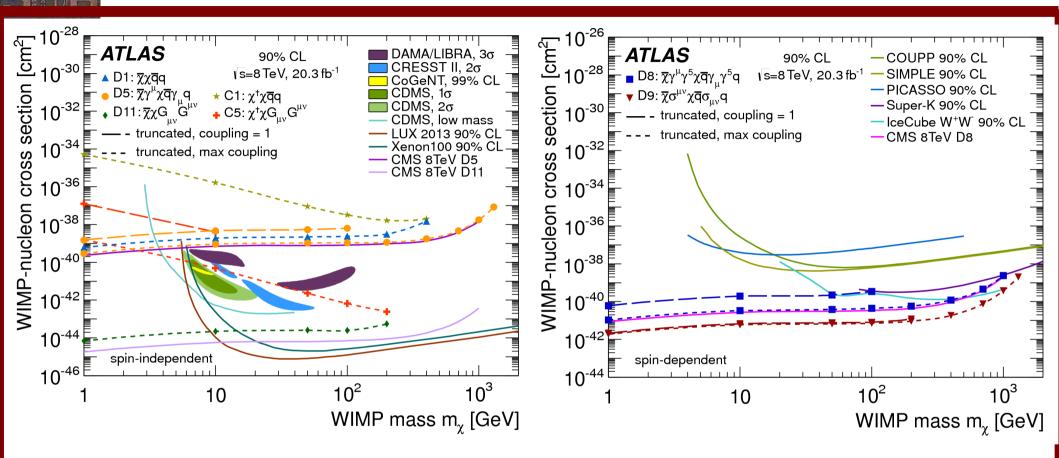


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Eur. Phys. J. C (2015) 75:299 DOI 10.1140/epjc/s10052-015-3517-3

DM@LHC2016, 30<sup>th</sup> March

## Mono-Jet



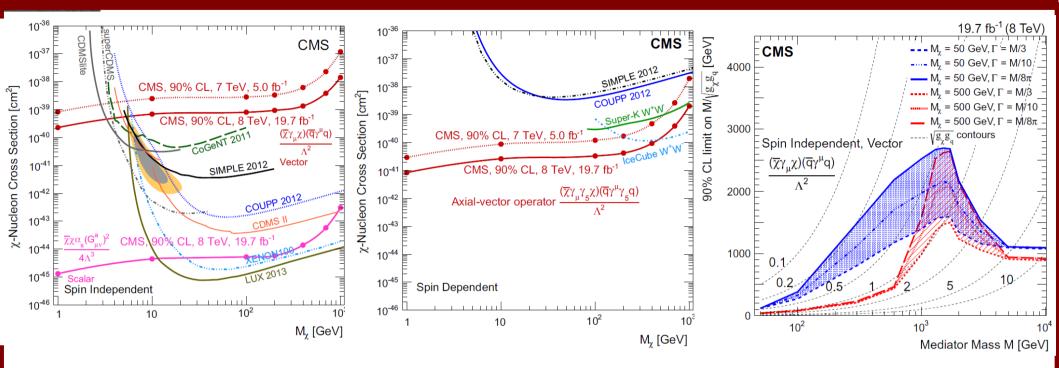
Eur. Phys. J. C (2015) 75:299 DOI 10.1140/epjc/s10052-015-3517-3

DM@LHC2016, 30<sup>th</sup> March

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## Mono-Jet



Eur. Phys. J. C (2015) 75:235 DOI 10.1140/epjc/s10052-015-3451-4

DM@LHC2016, 30<sup>th</sup> March

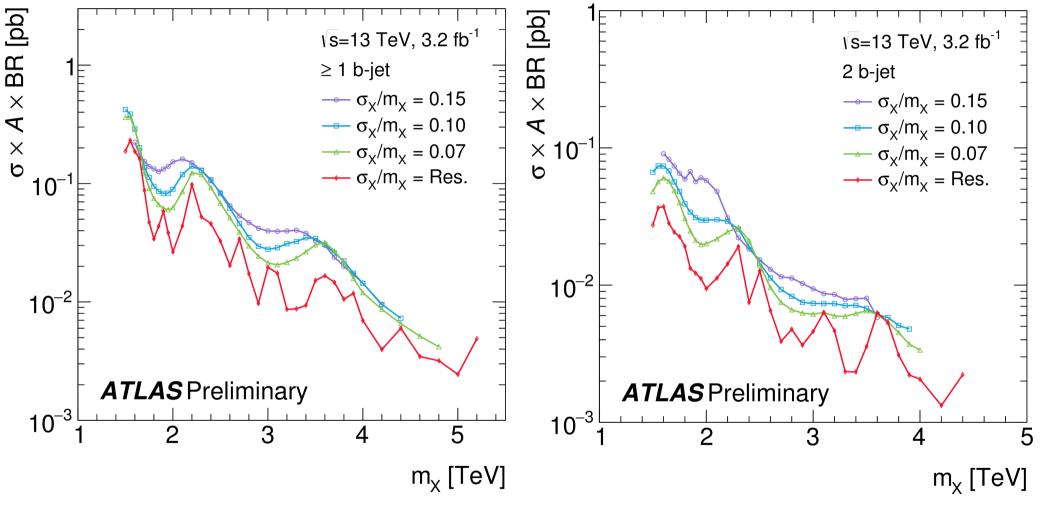


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## **Di-beauty-Jets**

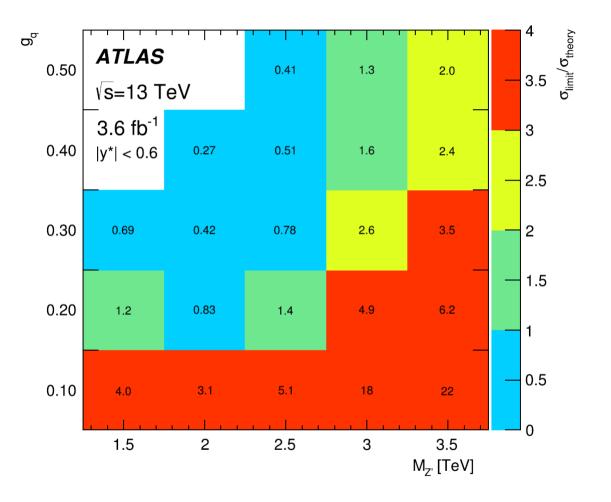


• Gaussian limits: model independent approach

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## **Di-Jet**



Physics Letters B 754 (2016) 302–322 http://dx.doi.org/10.1016/j.physletb.2016.01.032

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